IVOA Spectral Data Model
Version 0.98c Rev 1

IVOA Working Draft
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This version:
    http://www.ivoa.net/Documents/WD/SpectrumDM/SpectrumDM-2005XXXX.html
Latest version:
    http://www.ivoa.net/Documents/latest/SpectrumDM.html
Previous versions:
    http://hea-www.harvard.edu/~jcm/vo/docs/spec0.93.html
            Note: ivoa.net links are not yet active; the above are placeholders.

Editors:
    Jonathan McDowell, Doug Tody
Contributors:
    Jonathan McDowell, Doug Tody, Kelly McCusker, Tamas Budavari, Markus Dolensky, Frank Valdes, Pavlos Protopapas, Arnold Rots, and the IVOA Data Access Layer and Data Model Working Groups.

<div align="center">**Abstract**</div>

We present a data model describing the structure of spectrophotometric datasets with spectral and temporal coordinates and associated metadata. This data model may be used to represent spectra, time series data. and segments of SED (Spectral Energy Distributions).

Status of this document

This is a Working Draft, developed with the intention to support the Simple Spectral Access Protocol. The working group seeks confirmation that comments have been addressed to the satisfaction of the community.

This is an IVOA Working Draft for review by IVOA members and other interested parties. It is a draft document and may be updated, replaced, or obsoleted by other documents at any

time. It is inappropriate to use IVOA Working Drafts as reference materials or to cite them as other than "work in progress." A list of current IVOA Recommendations and other technical documents can be found at http://www.ivoa.net/Documents.

The **Virtual Observatory (VO)** is general term for a collection of federated resources that can be used to conduct astronomical research, education, and outreach.

The **International Virtual Observatory Alliance (IVOA)** (http://www.ivoa.net) is a global collaboration of separately funded projects to develop standards and infrastructure that enable VO applications.

Contents

Part 1: Spectrum Data Model

1 Introduction and Motivation

Spectra are stored in many different ways within the astronomical community. In this document we present a proposed abstraction for spectral data and serializations in VOTABLE, FITS, and XML, for use as a standard method of spectral data interchange.

We distinguish in several places between the implementation proposed in this document, referred to as Version 1, and capabilities proposed for possible later implementation.

2 Requirements

We need to represent a single 1-dimensional spectrum in sufficient detail to understand the differences between two spectra of the same object and between two spectra of different objects.

We need to represent time series photometry, with many photometry points of the same object at different times.

Finally, we need to represent spectral energy distributions (SED) which consist of multiple spectra and photometry points, usually for a single object.

3 Spectral Data Model summary

3.1 Model Components

Our model for a spectrum is a set of one or more data points (photometry) each of which share the same contextual metadata (aperture, position, etc.). Specifically, a spectrum will have arrays of the following values:

- Flux value, with upper and lower statistical (uncorrelated) errors

- Spectral coordinate (e.g. wavelength), central and bin min and max

- (Optionally) Time coordinate, convertible to MJD UTC

- Optional Quality mask

- Optional spectral resolution array

and will have associated metadata including:

- Data collection and Dataset ID

- Exposure time in seconds

- Position of aperture center, given as J2000 ICRS degrees.

- Aperture size in degrees

- Systematic (correlated) error

- Bibcode

In section 4 we elaborate these concepts in detail, including some complications that we explicitly do not attempt to handle in this version. The data model fields and possible values are listed. We distinguish between optional and required fields in the text, as well as via a "Req"

column in the tables which has values of R (Required) and O (Optional). Where appropriate we list those values of the physical units which interoperable implementations are required to recognize.

Figure 1: UML class diagram for the spectral data model.
The Coverage, Curation, DataID classes are shown in detail below in Section 4, both in diagram form and text descriptions. The text descriptions are definitive.
The minimal required content is:

- Spectrum model version

- Target name

- Characterization Coverage.Location and Coverage.Bounds.Extent descriptions of the location and extent of the data in the RA, Dec, time and spectral domains

- the Curation.Publisher field

- the descriptions of the spectral coordinate and flux fields including UCD and units

- For each point: the values of the spectral coordinate and flux.

Figure 2: Diagram for Data object

Figure 3: Diagram for Characterization object

Figure 4: Diagram for CoordSys object

Figure 5: Diagram for remaining metadata: Curation, DataID, Derived, Target objects

3.2 Units

We adopt the WCS/OGIP convention for units: Document OGIP 93-001 (http://legacy.gsfc.nasa.gov/docs/heasarc/ofw
Briefly, units are given in the form

```
10**(-14) erg/cm**2/s/Hz, 10**3 Jy Hz
```

i.e. with exponents denoted by **, division by /, multiplication by a space.

This format is mostly consistent with the AAS standards for online tables in journals (http://grumpy.as.arizona.edu/~gschwarz/unitstandards.html) except for the use of space rather than "." for multiplication and the fact that we do not require the use of SI units.

SI prefixes for units are to be recognized; for instance, the listing of "m" as a known unit for wavelength implies that "cm", "nm", and "um" (with "u" the OGIP convention for rendering "micro") are also acceptable.

Until IVOA generic unit conversion software is mature and widely deployed, it is helpful to interoperable applications to include a representation of the units in "base SI form", including only the base units kg, m, s (and possibly A, sr) with a numeric prefix. Pedro Osuna and Jesus Salgado have proposed a representation in the spirit of dimensional analysis, using the symbols M, L, T to signify kg, m, s respectively and omitting the ** for powers, so that
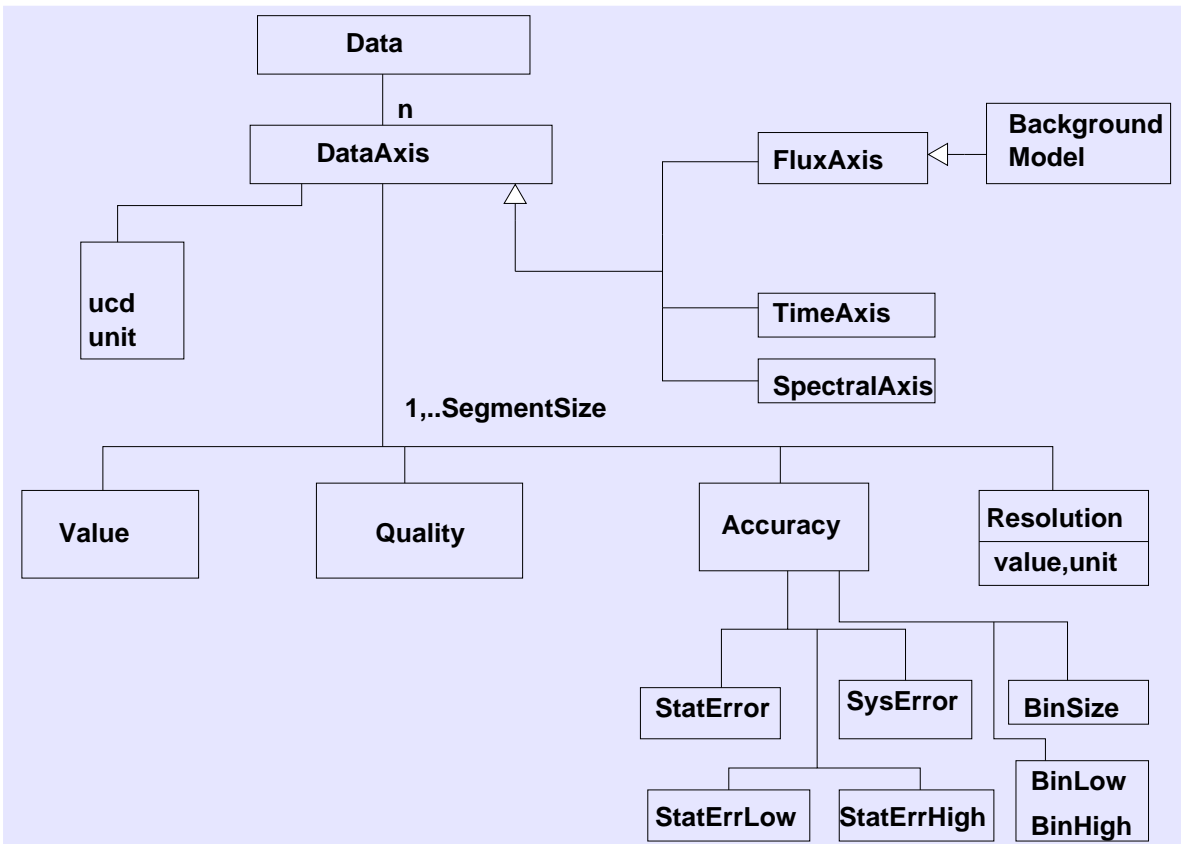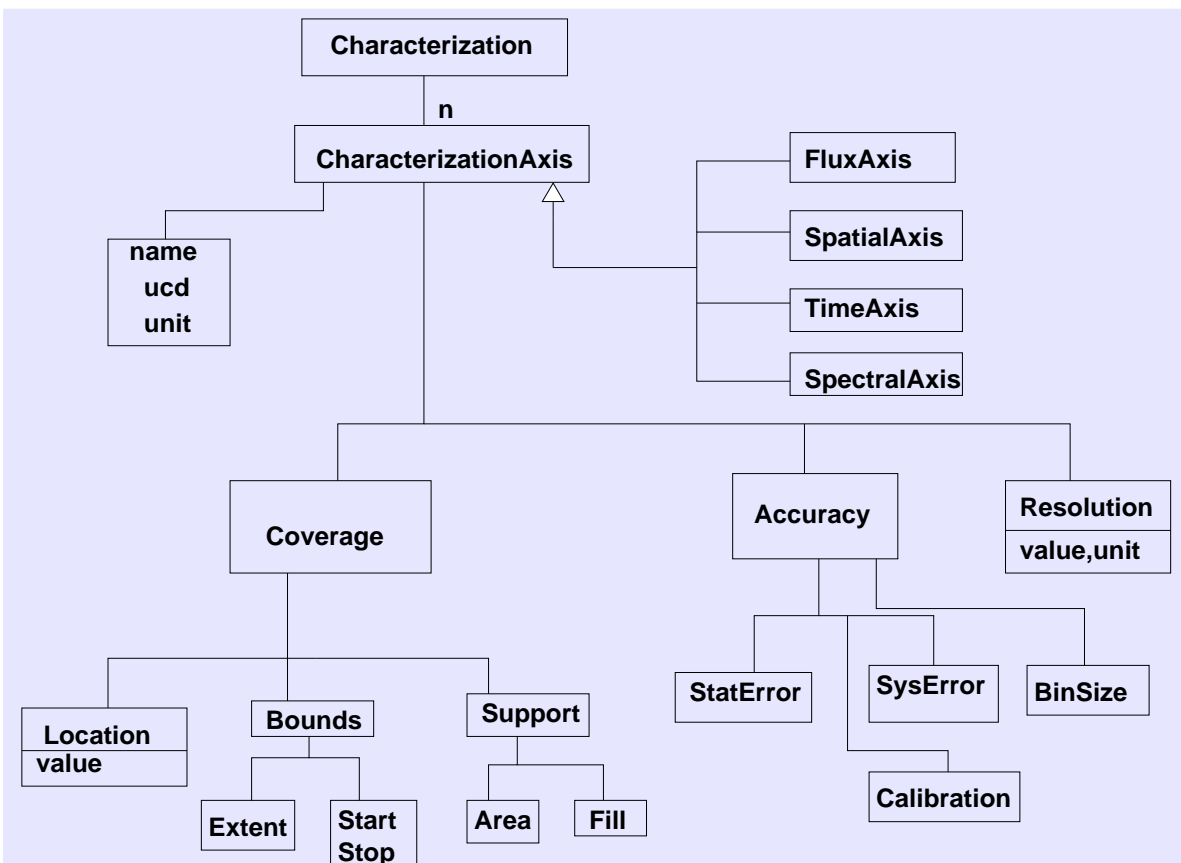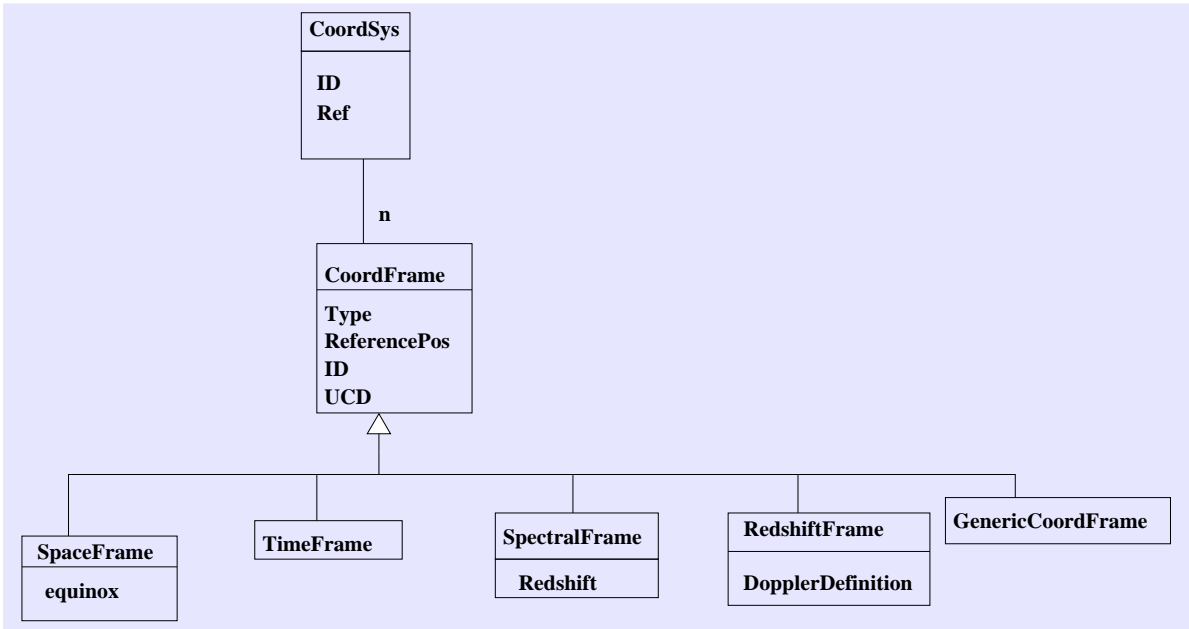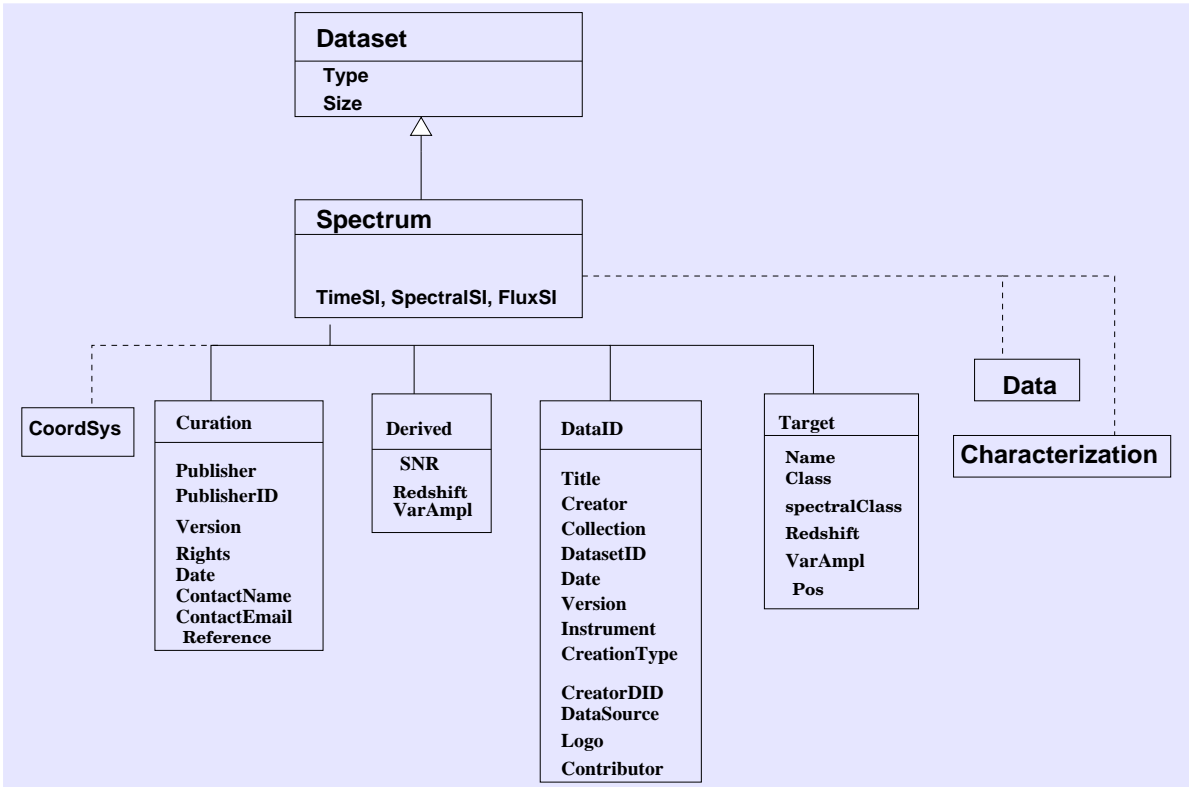
```
10**3 Jy Hz
```

which is equivalent to

```
10**-23 kg s**-2
```

is written compactly as

```
10-23MT-2
```

This alternate representation is supported for the main model fields (time, spectral coordinate and flux) only.

Although the spectral model is flexible enough to permit different units for each field, as a matter of style we strongly recommend that whenever possible the same units should be used for compatible fields (e.g. flux and error on flux).

3.3 UCDs

UCDs or Uniform Content Descriptors are the IVOA's standardized vocabulary for astronomical concepts. In this document we use UCDs as field attributes (for example, element attributes in XML) to distinguish alternate physics within the same data model roles - for example, to distinguish frequency versus wavelength on the spectral coordinate 'X-axis'.

The current list of UCDs is http://cdsweb.u-strasbg.fr/UCD/ucd1p-words.txt with syntax defined in the UCD recommendation http://www.ivoa.net/Documents/latest/UCD.html.

UCDs should be case insenstive.

3.4 UTYPEs

UTYPE was a concept introduced in VOTABLE to label fields of a hierarchical data model. The word is now used generally to mean a standard identifier for a data model field. They are also case-insensitive and are of the form a.b.c.d where the dots indicate a 'has-a' hierarchy with the leftmost element being the containing model and the rightmost element being the lowest level

element referred to. This is quite close to a simple XPATH in an XML schema, but we chose not to use slash instead of dot to emphasize that we are only specifiying the element type, not the exact position in an instance (so no sophisticated query syntax). We use the terms 'data model field' and 'UTYPE' interchangeably.

3.5 Data Model Fields

The DM fields (or UTYPEs) for the Spectrum DM are tabulated on the following papges. The field names are to be used as the UTYPE values in VOTABLE serializations. Required fields are in bold. The fields are explained in more detail in the following sections.

Field	UCD1+	Meaning	Req	Default
		Spectrum metadata fields		
Spectrum.DataModel		Version of this model	R	SPECTRUM-1.0
Spectrum.Type		Segment type	O	Spectrum
Spectrum.Length		Number of points	O	(must be derived)
Spectrum.TimeSI	time;arith.zp	SI factor and dimensions	O	
Spectrum.SpectralSI	-	SI factor and dimensions	O	
Spectrum.FluxSI	-	SI factor and dimensions	O	
Spectrum.CoordSys.ID		ID string for coordinate system	O	
Spectrum.CoordSys.SpaceFrame.Name		ICRS or FK5	O	ICRS
Spectrum.CoordSys.SpaceFrame.Equinox	time.equinox;pos.frame	Equinox	O	2000.0
Spectrum.CoordSys.TimeFrame.Name	time.scale	Time type, always equal to MJD	O	MJD
Spectrum.CoordSys.TimeFrame.Zero	time;arith.zp	Zero point of timescale in MJD	O	0.0
Spectrum.CoordSys.TimeFrame.Name	time.scale	Timescale	O	TT
Spectrum.CoordSys.TimeFrame.RefPos	time.scale	Times of photon arrival are at this location	O	TOPOCENTER
Spectrum.CoordSys.SpectralFrame.RefPos	?	Spectral frame origin	0	TOPOCENTER
Spectrum.Curation.Publisher	meta.curation	Publisher	R	
Spectrum.Curation.Reference	meta.bib.bibcode	URL or Bibcode for documentation	O	Unknown
Spectrum.Curation.PublisherID	meta.ref.url;meta.curation	URI for VO Publisher	O	Unknown
Spectrum.Curation.PublisherDID	meta.ref.url;meta.curation	Publisher's ID for the dataset ID	O	Unknown
Spectrum.Curation.Version	meta.version;meta.curation	Version info	O	Unknown
Spectrum.Curation.ContactName	meta.bib.author;meta.curation	Contact name	O	Unknown
Spectrum.Curation.ContactEmail	meta.ref.url;meta.*email*	Contact email	O	Unknown
Spectrum.Curation.Rights		Restrictions: public, proprietary, mixed	O	Public
Spectrum.Curation.Date		Date curated dataset last modified	O	Unknown
Spectrum.Target.Name	meta.id;src	Target name	R	
Spectrum.Target.Description		Target descriptive text	O	
Spectrum.Target.Class	src.class	Target or object class	O	Unknown
Spectrum.Target.spectralClass	src.spType	Object spectral class	O	Unknown
Spectrum.Target.redshift	src.redshift	Target redshift	O	Unknown
Spectrum.Target.pos	pos.eq;src	Target RA and Dec	O	Unknown or variable
Spectrum.Target.VarAmpl	src.var.amplitude	Target variability amplitude, typical	O	Unknown

Field	UCD1+	Meaning	Req	Default
Spectrum.DataID.Title	meta.title;meta.dataset	Dataset Title	O	None
Spectrum.DataID.Creator		VO Creator ID	O	Unknown
Spectrum.DataID.Collection		Collection name(s)	O	None
Spectrum.DataID.DatasetID	meta.id;meta.dataset	IVOA Dataset ID	O	Unknown
Spectrum.DataID.CreatorDID	meta.id	Creator's ID for the dataset	O	None
Spectrum.DataID.Date	time;meta.dataset	Data processing/creation date	O	Unknown
Spectrum.DataID.Version	meta.version;meta.dataset	Version of dataset	O	Unknown
Spectrum.DataID.Instrument	meta.id;instr	Instrument ID	O	Unknown
Spectrum.DataID.CreationType		dataset creation type (archive, cutout,derived)	O	Archival
Spectrum.DataID.Logo	meta.ref.url	URL for creator logo	O	Unknown
Spectrum.DataID.Contributor		Contributor (may be many)	O	Unknown
Spectrum.DataID.DataSource		Original data type: survey, pointed, theory, analytic	O	Unknown
Spectrum.Derived.SNR	stat.snr	Signal-to-noise for spectrum	O	Unknown
Spectrum.Derived.redshift.Value		Measured redshift for spectrum	O	Undefined
Spectrum.Derived.redshift.StatError	stat.error;src.redshift	Error on measured redshift	O	Unknown
Spectrum.Derived.redshift.Confidence		Confidence value on redshift	O	Undefined
Spectrum.Derived.VarAmpl	src.var.amplitude;stat.ratio	Variability amplitude as fraction of mean (0 to 1)	O	

Field	UCD1+	Meaning	Req	Default
Spectrum.Char Axis definition fields				
Spectrum.Char.FluxAxis.name		name for flux	O	Flux
Spectrum.Char.FluxAxis.ucd		ucd for flux	R	
Spectrum.Char.FluxAxis.Unit		Unit for flux	R	
Spectrum.Char.SpectralAxis.name		name for spectral axis	O	SpectralCoord
Spectrum.Char.SpectralAxis.ucd		ucd for spectral coord	R	
Spectrum.Char.SpectralAxis.Unit		Unit for spectral coord	R	
Spectrum.Char.TimeAxis.name		name for time axis	O	Time
Spectrum.Char.TimeAxis.ucd		ucd for time	O	time
Spectrum.Char.TimeAxis.Unit		Unit for time	O	d
Spectrum.Char.SpatialAxis.name		name for spatial axis	O	Sky
Spectrum.Char.SpatialAxis.ucd		ucd for spectral coord	O	pos.eq
Spectrum.Char.SpatialAxis.Unit		Unit for spectral coord	O	deg
Spectrum.Char Coverage Fields				
Spectrum.Char.SpatialAxis.Coverage.Location.Value	pos.eq	RA and Dec, deg,	R	
Spectrum.Char.SpatialAxis.Coverage.Bounds.Extent	instr.fov	Aperture angular diameter, deg	R	
Spectrum.Char.SpatialAxis.Coverage.Support.Area		Aperture region	O	circle on location ameter = extent
Spectrum.Char.SpatialAxis.Coverage.Support.Fill	stat.*fill*;pos.eq	Sampling Filling factor	O	1.0
Spectrum.Char.TimeAxis.Coverage.Location.Value	time.epoch	Midpoint of exposure, on MJD scale, d	R	
Spectrum.Char.TimeAxis.Coverage.Bounds.Extent	time.expo	Total exposure time	R	
Spectrum.Char.TimeAxis.Coverage.Bounds.Start	time.expo.start	Start time	O	
Spectrum.Char.TimeAxis.Coverage.Bounds.Stop	time.expo.end	Stop time	O	
Spectrum.Char.TimeAxis.Coverage.Support.Fill	time;stat.*fill*;time	Sampling Filling factor	-	O
Spectrum.Char.SpectralAxis.Coverage.Location.Value	instr.bandpass	Band, consistent with RSM		O
Spectrum.Char.SpectralAxis.Coverage.Bounds.Extent	instr.bandwidth	Width of spectrum in A or other spec. coord. (See text)	R	
Spectrum.Char.SpectralAxis.Coverage.Bounds.Start	em.*;stat.min	Start in spectral coordinate	(see text)	O
Spectrum.Char.SpectralAxis.Coverage.Bounds.Stop	em.*;stat.max	Stop in spectral coordinate	(see text)	O
Spectrum.Char.SpectralAxis.Coverage.Support.Fill	stat.*fill*;em.*	Sampling Filling factor	O	1.0

Field	UCD1+	Meaning	Req	Default
		Spectrum.Char Accuracy Fields - global		
Spectrum.Char.FluxAxis.Accuracy.StatError	phot.flux;em...;stat.error	error	O	Unknown
Spectrum.Char.FluxAxis.Accuracy.SysError	phot.flux;em...;stat.error.sys	Systematic error	O	Unknown
Spectrum.Char.SpectralAxis.Accuracy.BinSize	em.*;?	Wavelength bin size	O	Undefined
Spectrum.Char.SpectralAxis.Accuracy.StatError	em.*; stat.error	Spectral coord measurement error	O	0
Spectrum.Char.SpectralAxis.Accuracy.SysError	em.*; stat.error	Spectral coord measurement error	O	0
Spectrum.Char.SpectralAxis.Accuracy.Calibration		Type of coord calibration	O	CALIBRATED
Spectrum.Char.SpectralAxis.Resolution	instr.spectr.resolution	Spectral resolution FWHM	O	Accuracy.BinSize
Spectrum.Char.TimeAxis.Accuracy.BinSize	em.*;?	Time bin size	O	Undefined
Spectrum.Char.TimeAxis.Accuracy.StatError	time; stat.error	Time coord measurement statistical error	O	Unknown
Spectrum.Char.TimeAxis.Accuracy.SysError	time; stat.error.sys	Time coord measurement systematic error	O	Unknown
Spectrum.Char.TimeAxis.Accuracy.Calibration		Type of coord calibration	O	CALIBRATED
Spectrum.Char.TimeAxis.Resolution	TimeAxis.Accuracy.resolution	Temporal resolution FWHM	O	Accuracy.BinSize
Spectrum.Char.SpatialAxis.Accuracy.StatError	pos.eq;stat.error	Astrometric statistical error	O	Unknown
Spectrum.Char.SpatialAxis.Accuracy.SysError	pos.eq;stat.error.sys	Systematic error	O	Unknown
Spectrum.Char.SpatialAxis.Accuracy.Calibration		Type of coord calibration	O	CALIBRATED
Spectrum.Char.SpatialAxis.Resolution	instr.ang-res	Spatial resolution of data	O	

Field	UCD1+	Meaning	Req	Default
		Data Fields		
Spectrum.Data.FluxAxis.Value		Flux values for points	R	
Spectrum.Data.FluxAxis.ucd		ucd for flux	O	Char.FluxAxis.ucd
Spectrum.Data.FluxAxis.Unit		Unit for flux	O	Char.FluxAxis.Unit
Spectrum.Data.SpectralAxis.Value		Spectral coordinates for points	R	(Char.SpectralAxis.Location)
Spectrum.Data.SpectralAxis.ucd		ucd for spectral coord	O	Char.SpectralAxis.ucd
Spectrum.Data.SpectralAxis.Unit		Unit for spectral coord	O	Char.SpectralAxis.Unit
Spectrum.Data.TimeAxis.Value		Time coordinates for points	O	Char.TimeAxis.Location
Spectrum.Data.TimeAxis.ucd		ucd for time	O	Char.TimeAxis.ucd
Spectrum.Data.TimeAxis.Unit		Unit for time	O	Char.TimeAxis.Unit
Spectrum.Data.BackgroundModel.Value		Flux values for points	O	No background model
Spectrum.Data.BackgroundModel.ucd		ucd for background flux	O	Points.FluxAxis.ucd
Spectrum.Data.BackgroundModel.Unit		Unit for background flux	O	Points.FluxAxis.Unit

Field	UCD1+	Meaning	Req	Default
Accuracy Fields - per data point (default to corresponding Spectrum.Char values)				
Spectrum.Data.FluxAxis.Accuracy.StatError	phot.flux;em...;stat.error	symmetric error	O	(Char)
Spectrum.Data.FluxAxis.Accuracy.StatErrLow	phot.flux;em...;stat.error;stat.min	Lower error	O	(Char)
Spectrum.Data.FluxAxis.Accuracy.StatErrHigh	phot.flux;em...;stat.error;stat.max	Upper error	O	(Char)
Spectrum.Data.FluxAxis.Accuracy.SysError	phot.flux;em...;stat.error.sys	Systematic error	O	(Char)
Spectrum.Data.FluxAxis.Quality	meta.code.qual;phot.flux,em...	Quality mask	O	0
Spectrum.Data.FluxAxis.Quality.n		String value, for n = 0,1,2..; meaning of quality value	O	Undefined
Spectrum.Data.SpectralAxis.Accuracy.BinSize	em.*;?	Wavelength bin size	O	(Char)
Spectrum.Data.SpectralAxis.Accuracy.BinLow	em.*;stat.min	Spectral coord bin lower end	O	Midpoint of values
Spectrum.Data.SpectralAxis.Accuracy.BinHigh	em.*;stat.max	Spectral coord bin upper end	O	Midpoint of values
Spectrum.Data.SpectralAxis.Accuracy.StatError	em.*; stat.error	Spectral coord measurement error	O	(Char)
Spectrum.Data.SpectralAxis.Accuracy.StatErrLow	em.*; stat.error; stat.min	Spectral coord measurement lower error	O	(Char)
Spectrum.Data.SpectralAxis.Accuracy.StatErrHigh	em.*; stat.error; stat.max	Spectral coord measurement upper error	O	(Char)
Spectrum.Data.SpectralAxis.Accuracy.SysError	em.*; stat.error	Spectral coord systematic error	O	(Char)
Spectrum.Data.SpectralAxis.Resolution	instr.spectr.resolution	Spectral resolution FWHM	O	(Char)
Spectrum.Data.TimeAxis.Accuracy.BinSize	em.*;?	Time bin size	O	(Char)
Spectrum.Data.TimeAxis.Accuracy.BinLow	em.*;stat.min	Time bin start	O	Midpoint of values
Spectrum.Data.TimeAxis.Accuracy.BinHigh	em.*;stat.max	Time bin stop	O	Midpoint of values
Spectrum.Data.TimeAxis.Accuracy.StatError	time; stat.error	Time coord measurement statistical error	O	(Char)
Spectrum.Data.TimeAxis.Accuracy.StatErrLow	time; stat.error, stat.min	Time coord measurement lower error	O	(Char)
Spectrum.Data.TimeAxis.Accuracy.StatErrHigh	time; stat.error, stat.max	Time coord measurement upper error	O	(Char)
Spectrum.Data.TimeAxis.Accuracy.SysError	time; stat.error.sys	Time coord measurement systematic error	O	(Char)
Spectrum.Data.TimeAxis.Resolution	time.resolution	Temporal resolution FWHM	O	(Char)
Spectrum.Data.BackgroundModel.Accuracy.StatError	phot.flux;em...;stat.error	Symmetric error	O	(Char)
Spectrum.Data.BackgroundModel.Accuracy.StatErrLow	phot.flux;em...;stat.error;stat.min	Lower error	O	(Char)
Spectrum.Data.BackgroundModel.Accuracy.StatErrHigh	phot.flux;em...;stat.error;stat.max	Upper error	O	(Char)
Spectrum.Data.BackgroundModel.Accuracy.SysError	phot.flux;em...;stat.error.sys	Systematic error	O	(Char)
Spectrum.Data.BackgroundModel.Quality	meta.code.qual;phot.flux,em...	Quality mask	O	0

4 Spectral data model Measurement objects

4.1 Spectral coordinate

Astronomers use a number of different spectral coordinates to label the electromagnetic spectrum. The cases enumerated by Greisen et al. (2003) are listed below with their UCDs.

REQUIRED: Exactly one Spectrum.Char.SpectralAxis field should be present, with units and one of the UCD values listed below. We distinguish between the VO data model field name (which might be used for VOTABLE UTYPE), the FITS WCS name (provided for comparison only), and the UCD1+ names.

Note 1: For this version, only the first four entries, Wavelength, Frequency, Energy, and spectral channel, should be used for interoperable transmission of data - implementations are not required to understand (convert) the other UCD values.

Note 2: For the velocity cases, the UCD uses a spect.dopplerVeloc tree rather than a *src.veloc* tree, because the velocity here is really a labelling of a spectral coordinate, and the link to the physical radial velocity of the different emission sources contributing to the spectrum is rather indirect.

Field	FITS WCS	UCD1+	Meaning	Units
PREFERRED CHOICES				
SpectralAxis.ucd	**WAVE**	em.wl	Wavelength	Angstrom, m
SpectralAxis.ucd	**FREQ**	em.freq	Frequency of photon	Hz
SpectralAxis.ucd	**ENER**	em.energy	Photon energy	erg, eV, J
SpectralAxis.ucd	-	instr.pixel;em.wl	Instrumental spectral bin	chan
ALTERNATIVE CHOICES				
SpectralAxis.ucd	**WAVN**	em.wavenumber	Wavenumber	m**(-1)
SpectralAxis.ucd	**AWAV**	em.wl; obs.atmos	Air wavelength	Angstrom, m
SpectralAxis.ucd	**WAVE-LOG**	em.wl	Log wavelength	
SpectralAxis.ucd	**FREQ-LOG**	em.freq	Log frequency of photon	
SpectralAxis.ucd	**ENER-LOG**	em.energy	Log photon energy	
SpectralAxis.ucd	**VELO**	spect.dopplerVeloc	Apparent radial velocity	m/s
SpectralAxis.ucd	**VRAD**	spect.dopplerVeloc.radio	Radio velocity	m/s
SpectralAxis.ucd	**VOPT**	spect.dopplerVeloc.opt	Optical velocity	m/s
SpectralAxis.ucd	**BETA**	spect.dopplerVeloc	Velocity (c=1)	-

4.2 Flux (Spectral Intensity) Object

Two instances of the Flux object are supported: Flux and BackgroundModel. The Flux may be either the background-subtracted net flux or the total flux (the source+background), in the latter case hopefully with the BackgroundModel (see below). Net and total flux are distinguished by the 'src.net' UCD adjective.

For each of these cases, there are many slightly different physical quantities covered by the general concept of Flux; we distinguish them by their UCD.

Note in particular the distinction between the unit **count** (an instrumental value) and the unit **photon** (used in the photon number flux, i.e. the number of photons incident; photon number flux = energy flux divided by photon energy).

Note: The concept of the "nu L-nu" or "lambda L-lambda" flux, or equivalently the luminosity per logarithmic energy interval L(log nu), is a distinct concept in the world of spectral

energy distributions - and it's a different concept from the bolometric luminosity, which has the same units. The UCD board has not yet approved a UCD expressing this concept; we have to use phys.luminosity and infer the concept from the units. My solution for brightness temperature is also rather questionable.

Flux Value choices			
Field	UCD1+	Meaning	Unit
FluxAxis.ucd	phot.fluDens.uncal;em.wl	Flux density per unit wave, not absolutely calibrated.	
FluxAxis.ucd	phot.fluDens;em.wl	Flux density per unit wave.	erg cm$^{-2}s^{-1}\AA^{-1}$ Wm^{-2}m^{-1}, keVcm$^{-2}s^{-1}\AA^{-1}$
FluxAxis.ucd	phot.fluDens;em.freq	Flux density per unit freq.	erg cm$^{-2}s^{-1}$Hz^{-1} Jy, Wm^{-2}Hz^{-1}
FluxAxis.ucd	phot.fluDens;em.energy	Flux density per energy interval	keV cm$^{-2}s^{-1}$keV^{-1}
FluxAxis.ucd	phot.fluDens;em.energy; meta.number	Photons per unit area, time, energy	photon cm$^{-2}s^{-1}$keV^{-1}
FluxAxis.ucd	phot.fluDens;em.wl	Flux density per log wave interval (nu fnu)	Jy Hz
FluxAxis.ucd	phot.fluDens.sb;em.wl	Surface brightness per unit wavelength	erg cm$^{-2}s^{-1}\AA^{-1}$arcsec^{-2}
FluxAxis.ucd	phot.fluDens.sb;em.freq	Surface brightness per unit frequency	Jy sr^{-1}
FluxAxis.ucd	phot.count	Counts in spectral channel	count
FluxAxis.ucd	arith.rate;phot.count	Count rate in spectral channel	count/s
FluxAxis.ucd	arith.ratio; phot.fluDens	Flux ratio of two spectra	-
FluxAxis.ucd	phys.luminosity;em.wl	Luminosity per unit wave	erg$s^{-1}\AA^{-1}$, W/m
FluxAxis.ucd	phys.luminosity;em.freq	Luminosity per unit freq	erg$s^{-1}Hz^{-1}$, W/Hz
FluxAxis.ucd	phys.luminosity;em.energy	Luminosity per unit energy	ergs^{-1}keV^{-1}
FluxAxis.ucd	phys. luminosity;em.energy	Luminosity per log frequency	erg s^{-1}, W
FluxAxis.ucd	phys.energDensity	Radiation energy density per unit volume, per unit wave etc.	erg cm^{-3}, W m^{-3}
FluxAxis.ucd	phys.luminosity; phys.angArea;em.wl	Flux per unit solid angle (at source)	erg cm$^{-2}s^{-1}sr^{-1}\AA^{-1}$
FluxAxis.ucd	instr.antenna-temp	Antenna temperature	K
FluxAxis.ucd	phot.fluDens;phys.temperature	Brightness temperature	K
FluxAxis.ucd	phot.mag	Magnitude in defined band	mag
FluxAxis.ucd	phot.mag	AB (spectrophotometric) magnitude	mag
FluxAxis.ucd	phot.fluDens; instr.beam	Flux per resolution element (e.g. Jy/beam)	Jy/beam
FluxAxis.ucd	phot.mag.sb	Surface brightness in magnitudes	mag arcsec^{-2}

4.3 BackgroundModel Object

We optionally allow a BackgroundModel value for each Flux value. We define NetFlux = TotalFlux - BackgroundModel. The name BackgroundModel, rather than Background, reminds us that it is an estimate: often, the BackgroundModel will be generated by taking a flux measurement at another location and rescaling it for any difference in exposure time or extraction aperture.

The BackgroundModel array is required to have the same UCD and units as the Flux array. It represents a model for the expected flux values if the Target had zero flux.

OPTIONAL: There may be at most one BackgroundModel.Value field present. It must have the same UCD as the Flux.

4.4 Time coordinate

For data with a time-series component, whether regularly sampled or sparse photometry points, the time coordinate is given by an elapsed time in some physical units (e.g. seconds or days) relative to a reference time.

This reference time is given in MJD as the field Spectrum.Char.TimeAxis.Coverage.Location, as described in the Characterization section. For a simple spectrum with no time-resolved data, this is the time of the observation.

For time-resolved data, the time coordinate refers to the midpoint of the sample interval. See the Space-Time Coordinates document for details of time coordinate complexity.

The time unit is specified by a string, and the only valid values for this unit are 's' (seconds) and 'd' (days).

4.5 Position coordinate

In general we may consider position coordinates as part of the measurement (and possibly varying from point to point), but this capability is not included in the current document. The celestial position of the aperture for the spectrum is given in the spatial Spectrum.Char.SpatialAxis.Coverage.Location field. The Spectrum.Char.SpatialAxis.Coverage.Location.Value field is an RA,Dec pair expressed in decimal degrees.

4.6 Accuracy Fields

We include accuracy models for both the coordinates (spectral, spatial and temporal) and the fluxes. The accuracy can appear in two places: in the global characterization, where it represents typical accuracy for the dataset, and in the data points themselves, providing a way to provide per-data-point errors. All the Accuracy fields are optional, both in the per-data-point fields and in the Characterization instances; the per-data fields default to the values in Characterization.

4.6.1 Coordinate bins

We express the bandpass for each spectral bin as a low and high value for the spectral coordinate, or as a width. The same is done for photometry points, which amounts to approximating a filter by a rectangular bandpass. Time bins are also given as low and high values or as a width. Note that the width values are suitable for Spectrum.Char (the global accuracy) while the bin low/high values only have meaning for Spectrum.Data (the per-data-point values).

Only one of BinSize, or both BinLow and BinHigh, must be present (possibly as a header parameter implying a constant value for each flux point). If absent, the bin limits are assumed to be halfway between the coordinate values and bounded by the range given in Char.*.Coverage.Extent.

4.6.2 Uncertainties

In addition to the binning, we allow the model to express uncertainties (which may be larger than the bin width), both statistical and systematic. We allow one or two-sided statistical errors

but only one-sided systematic errors. You can specify StatErr, or StatErrHigh/StatErrLow, but not both.

For position we have a single statistical error - a two-sided error doesn't make sense for a 2D coordinate. Eventually we may want a full error ellipse, but this is too complicated for the present model.

We also use a very simple error model for the fluxes: we include plus and minus flux errors, and a quality flag. The errors are understood as 1 sigma gaussian errors which are uncorrelated for different points in the spectrum. If the data provider has only upper limit information, it should be represented by setting the flux value and the lower error value equal to the limit, and the upper error value equal to zero (e.g. 5 (+0,-5)). In general applications may choose to render measurements as upper limits if the flux value is less than some multiple (e.g. 3) of the lower error. We also allow a systematic error value, assumed constant across a given spectrum and fully correlated (so that, e.g. it does not enter into estimating spectral slopes).

CLARIFICATION: the two-sided errors StatErrLow and StatErrHigh are the plus/minus ERRORS, not the (value+error, value-error). In other words, if Value = 10 and there is a symmetric uncertainty of 3, the ErrorLow and ErrorHigh are both +3.0, and NOT 7.0, 13.0. This is different from the sampling description BinLow and BinHigh, which give the VALUES at the low and high end of the bin. Thus if the central wavelength of the bin is 4200.0, and the bin size is 10, then the BinLow and BinHigh values are 4195.0, 4205.0 and NOT 10.0, 10.0. Note that because of this, 0.0 is NOT an acceptable default for BinLow and BinHigh, while it IS acceptable (albeit unlikely) for StatErrLow and StatErrHigh.

The StatErrLow, StatErrHigh, SysError fields for SpectralCoord, Time, Sky and Flux are optional; however, omitting these fields indicates that the errors are unknown. Data providers are STRONGLY encouraged to provide explicit error measures whenever possible.

4.6.3   Resolution

We also include a trivial resolution model: a single number nominally representing a FWHM spectral or time resolution expressed in the same units as the spectral or time coordinate. The default is to assume that the resolution is equal to the BinSize if defined. The spatial (sky) resolution may be useful to know if it exceeds the aperture size; the default is to assume it is equal to the aperture size.

4.6.4   Quality

The Quality model represents quality by an integer, with the following meanings: 0 is good data, 1 is data which is bad for an unspecified reason (e.g., no data in the sample interval), and other positive integers greater than 1 may be used to flag data which is bad or dubious for specific reasons.

The data provider may also define scalar string-valued metadata fields Quality.2, Quality.3... to define specific quality flags on a per-spectrum basis. Bitmasks, used in some archives such as SDSS, should be remapped to such independent Quality fields.

Quality defaults to zero, i.e. good data.

4.6.5   Calibration

We also introduce a Calibration field which can have the values ABSOLUTE, RELATIVE or UNCALIBRATED. This is expected to be particularly useful to describe the flux. ABSOLUTE indicates that the values in the data are expected to be correct within the given uncertainty.

RELATIVE indicates that although an unknown systematic error is present, the ratio of any two values will be correct. UNCALIBRATED indicates that although the values reflect a measurement of the given UCD, they are modified by an unspecified coordinate-dependent correction. Such values may be useful in the case of a spectrum with ABSOLUTE calibration on the wavelengths but UNCALIBRATED fluxes; the wavelengths of discontinuous features such as spectral lines can be measured on the assumption that the missing calibration function has no sharp discontinuities in the region of interest.

The Calibration fields are present in the global characterization accuracy but not in the per-data-point accuracy.

5   Associated Metadata Fields

Most of the associated metadata are generic observational metadata that can be applied to future data models, and are not specific to spectra.

5.1   CoordSys Fields

The CoordSys object is a simplified instance of the STC CoordSystem object. For XML serializations, it can be replaced by an actual STC CoordSystem instance.

A CoordSys consists of 1 or more CoordFrame objects, each of which defines the coordinates for a particular axis. The CoordSys has an overall ID string, which is user-defined and arbitrary. Each CoordFrame also has a type, a UCD and a ReferencePosition; the Reference Position gives the origin of the coordinate system (and thus also its rest frame).

For the space, time, and spectral axes we define specialized CoordFrames for convenience: SpatialFrame, TimeFrame and SpectralFrame. The CoordFrame names (types) for SpatialFrame and TimeFrame must be from a controlled list; for other frames, the type is an arbitrary string.

Token	Meaning	Note
UNKNOWN	Unknown origin	
RELOCATABLE	Relative origin	Suitable for simulations
CUSTOM	Origin specified wrt another system	
TOPOCENTER	Location of the observing device	(telescope)
BARYCENTER	Solar system barycenter	
HELIOCENTER	Center of the Sun	
GEOCENTER	Center of the Earth	
EMBARYCENTER	Earth-Moon barycenter	
MOON	Center of the Moon	
MERCURY	Center of Mercury	
VENUS	Center of Venus	
MARS	Center of Mars	
JUPITER	Center of Jupiter	
SATURN	Center of Saturn	
URANUS	Center of Uranus	
NEPTUNE	Center of Neptune	
PLUTO	Center of Pluto	
LSR	Kinematic local standard of rest	Redshift frame only
LSRD	Dynamic local standard of rest	Redshift frame only
GALACTIC_CENTER	Center of the Galaxy	
LOCAL_GROUP_CENTER	Barycenter of the Local Group	

Table 1: Allowed values for CoordFrame.ReferencePosition.Type

5.1.1   SpatialFrame

The SpatialFrame has an optional Equinox attribute which is used if the frame name is FK4 or FK5. The allowed frame names for SpatialFrame are listed below.

5.1.2 TimeFrame

The TimeFrame is defined by the frame name and the ReferencePosition. Allowed values of the name are given below.

One standard reference time in astronomy is the origin of Julian Day Number on the TT (Terrestrial Time) timescale, BC 4713 Nov 24 at 11:59:27.81 (Gregorian). Using TT is preferable to UTC because it does not contain leap seconds, so the elapsed time in days is just equal to the difference in JD values.

The ISO-8601 calendar format standard does not support dates before AD 1, so cannot express this reference time. Therefore, it is not a suitable format for internal representations of such reference times. However, non-default choices of reference time may be specified in external serializations by a date in ISO-8601 format, e.g. "2004-11-30T11:59:00.01".

In this version of the model we require use of MJD as the time type for absolute times. (ISO dates and JD are other possibilities covered by the STC document).

(Note that in the FITS serialization, the MJDREF keyword allows definition of reference times in decimal days relative to MJD 0.0 = JD 2400000.5.)

5.1.3 SpectralFrame

The spectral frame has an optional Redshift attribute to specify a rest frame; it is used only if the the frame's ReferencePosition is "CUSTOM".

5.1.4 RedshiftFrame

When you convert the spectral coordinate to velocity or redshift (relative to some assumed rest-frame spectral feature) you need to record some other metadata. Other coordinate system information needed for velocity spectral coordinates include the observation-fixed spectral frame, the observatory location, the source redshift, and the velocity zero point (in Greisen et al, SSYSOBS, OBSGEO, VELOSYS, RESTFRQ/RESTWAV). However, we omit these in the current model. The only metadata we provide is the Doppler Definition - optical, radio or pseudo-relativistic.

Note that you only use RedshiftFrame if you're measuring things in velocities; a rest-frame spectrum of a redshifted quasar whose spectral axis is in Angstroms will be described by a SpectralFrame.

5.1.5 STC

Notes on compatibility with, and differences from, STC 1.0:

- We include the types (e.g. MJD, Cartesian) with the CoordSys (STC's CoordSys), while STC puts them with the individual Coords instances. Implicitly, this requires that within a spectrum, all coordinate values for a particular coordinate be of the same type.

- We add the extra Redshift attribute in the SpectralFrame, instead of the more complex CustomReferencePosition approach used in STC.

- In STC's XML serialization, the frame types and reference positions are enumerated defined elements. Here they are strings, and we require that the frame name is the frame type.

- We don't explicitly include the coordinate flavor.

OPTIONAL: All CoordSys values are optional , but data providers should take special care to check whether or not the defaults are appropriate for their data. The implications of the defaults are:

- Positions are given in ICRS RA,Dec in degrees and are heliocentric values (i.e. corrected for annual parallax and aberration, as normally found in source catalogs).

- Times are given in MJD days and represent times of photon arrival at the telescope.

- Spectral coordinates are as observed at the telescope, and not corrected for redshift, the motion of the Earth, etc.

Token	Meaning	Parameter(s)
UNKNOWN	Unknown frame	
CUSTOM	Custom frame	Pole, axis
AZ_EL	Azimuth and elevation	
BODY	Generic body (eg planet)	
ICRS	The ICRS frame	
FK4	FK4	Equinox
FK5	FK5	Equinox
ECLIPTIC	Ecliptic l,b	Equinox
GALACTIC_I	Old galactic LI,BI	
GALACTIC_II	Galactic LII,BII	
SUPER_GALACTIC	SGL, SGB	
MAG	Geomagnetic ref frame	
GSE	Geocentric Solar Ecliptic	
GSM	Geocentric Solar Magnetic	
SM	Solar Magnetic	
HGC	Heliographic	
HEE	Heliocentric Earth Ecliptic	
HEEQ	Heliocentric Earth Equatorial	
HCI	Heliocentric Inertial	
HCD	Heliocentric of Date	
GEO_C	Geocentric corotating	
GEO_D	Geodetic ref frame	Spheroid
MERCURY_C	Corotating planetocentric	
VENUS_C	Corotating planetocentric	
LUNA_C	Corotating planetocentric	
MARS_C	Corotating planetocentric	
JUPITER_C_III	Corotating planetocentric	
SATURN_C_III	Corotating planetocentric	
URANUS_C_III	Corotating planetocentric	
NEPTUNE_C_III	Corotating planetocentric	
PLUTO_C	Corotating planetocentric	
MERCURY_G	Corotating planetographic	
VENUS_G	Corotating planetographic	
LUNA_G	Corotating planetographic	
MARS_G	Corotating planetographic	
JUPITER_G_III	Corotating planetographic	
SATURN_G_III	Corotating planetographic	
URANUS_G_III	Corotating planetographic	
NEPTUNE_G_III	Corotating planetographic	
PLUTO_G	Corotating planetographic	

Table 2: Allowed values for CoordFrame.RefFrame.Name

Token	Meaning	Note
LOCAL	Relocatable (simulation) time	
TT	Terrestrial Time	
UTC	Coordinated Universal Time	
ET	Ephemeris Time	
TDB	Barycentric dynamical time	
TCG	Terrestrial Coordinate Time	
TCB	Barycentric Coordinate Time	
TAI	International Atomic Time	
LST	Local Sidereal Time	

Table 3: Allowed values for CoordFrame.TimeScale

5.2 Characterization

The Characterization metadata in this document are based on, but not identical to, the IVOA Note published by the Characterization data model subgroup in March 2006. The Characterization model has a set of CharacterizationAxis objects. Each CharacterizationAxis has an AxisFrame describing the axis, a Coverage describing the scope of the data, and optionally a Resolution and a Sampling object. We do not include Sampling in the current Spectrum model.

The CharacterizationAxis is identified by its UCD attribute. Spectrum instances should have Spatial, Time and Spectral characterization axes. To simplify things for the common axes, we define SpatialAxis, SpectralAxis, TimeAxis objects as special cases of CharacterizationAxis.

The Characterization fields will have a constant value for a given spectrum.

5.2.1 Coverage Fields

The coverage fields will have a constant value for a given spectrum. They describe the region of space, time and spectrum from which the data were taken.

We give a value giving the effective exposure time (useful for selecting among multiple spectra from the same instrument). The aperture is important to determine what part of an extended object is contributing to the spectrum; we allow a simple aperture description consisting of a single number representing the aperture diameter in decimal degrees.

The units of the spectral Coverage.Bounds.Extent and Coverage.Support should be the same as those of SpectralCoord.

For time, the Coverage.Bounds.Range is a pair of values giving the start and stop time; in principle we could also have Coverage.Support with a whole array of start-stop pairs indicating data accumulated over a series of intervals, but we don't support that yet.

The SpatialAxis.Coverage.Location and SpatialAxis.Coverage.Bounds.Extent, TimeAxis.Coverage.Location and TimeAxis.Coverage.Bounds.Extent fields are required. We don't require the spectral equivalents as these are easily derived from the data.

5.2.2 Region definitions

In the optional Char.SpatialAxis.Coverage.Support.Area we describe the detailed aperture shape on the sky. However, we don't allow a full STC region description. Our simplified region model allows for (1) a circle and (2) a polygon in a string representation: either

> circle x0 y0 r

or

> polygon x1 y1 x2 y2 x3 y3 ...

for example

```
circle 233.70 -13.32 0.00043
polygon 233.70 -13.32 233.71 -13.30 ...
```

where the positions and radii are required to be in degrees, in the coordinate system defined by CoordSys.

5.3 Derived Data Fields

The Derived Data object has useful, and optional, summary information about the spectrum. For now, we include the option of adding signal-to-noise and variability indicators and a measurement of the redshift.

5.3.1 Redshift measurement model

One common piece of derived data for a spectrum is the source redshift. We provide fields for both the redshift measured value and statistical error. We add a further measure of accuracy, the Confidence, which expresses a probability between 0 and 1 that the quoted errors do apply. This measure is used in the Sloan spectral service to provide a way of describing the estimated probability that the redshift is completely in error because the lines have been misidentified. Its default value is 1.0.

In general, such a Confidence could be useful for any measurement where the error probability distribution has multiple peaks in parameter space, and could later be added to the standard Accuracy model.

5.4 Curation model

The Curation is an object consistent with the Curation information in the document "Resource Metadata for the Virtual Observatory Version 1.01", although we have added a Reference field for a bibliographic reference.

REQUIRED: Curation.Publisher. As with the VO Resource Metadata document, this is the only required field in the Curation object. All other fields are optional.

5.5 Data Identification model

The Data Identification model gives the dataset ID for a particular spectrum, and its membership of larger collections. All DataId fields are optional.

The DataID.DatasetID is the IVOA/ADEC/ADS dataset identifier. By agreement between the AAS journals, the ADS and the ADEC (NASA data centers), dataset identifiers, described in http://vo.ads.harvard.edu/dv/, will be used to link journal articles back to the archival datasets containing the relevant observational data.

We introduce the concept of an dataset creation type, which can have one of the following three values:

- Archival, indicating that it is one of a collection of datasets (in this case spectra) generated in a systematic, homogeneous way and stored statically (or at least versioned). It will be possible to regenerate this dataset at a later date.

- Dynamic, indicating that the dataset was created 'on-the-fly', possibly by assembling the latest available data and applying a possibly changeable processing algorithm. A future attempt to generate this dataset in the same way may give different results.

- Custom, indicating that the dataset was created by manual analysis and processing, whether by data center staff or as part of a scientist's research project. Traceability of the processing may be incomplete.

The dataset is associated with one or more Collections (instrument name, survey name. etc.) indicating some degree of compatibility with other datasets sharing the same Collection properties. Examples of possible Collection values are: "WFC", "Sloan", "BFS Spectrograph", "MSX Galactic Plane Survey".

5.6   Target model

In spectral data it is particularly important to be able to specify the target of the observation, which may be an astronomical source or some other target (calibration, diffuse background, etc.). By explicitly including a target model we can not only facilitate searches on particular types of target but also support archives of model spectra for which the Coverage fields may not be relevant. The Target.Name field is required; all other Target fields are optional.

The Target.pos field gives a nominal RA and Dec for the target, for example the catalog position of the source; the Coverage.Location fields in the spectrum indicate the actual telescope pointing position for that spectrum. (An SED might have a single Target object with a known position, but many Spectrum objects with slightly different telescope pointings). Similarly, the Target.redshift is the assumed actual cosmological redshift of the astronomical object, if applicable (again, usually from a catalog, NED, etc.), while the redshifts in the DerivedData objects in the spectrum (segment) indicates a redshift measured from that spectrum.

At the moment there is no international standard list of valid values for Target class and spectral class. Nevertheless an initial deployment of the VO would gain some benefit from using archive-specific classes, and provide a framework for converging on a standard list.

5.7   Spectrum top level object

The Spectrum object contains the Data object with the actual data; the Target and DerivedData objects; and the standard dataset metadata of CoordSys, Characterization, Curation and DataID. We also add a CustomParams field to allow for propagation of unmodelled application-specific metadata.

In addition, we add an SIDim field for each axis giving the SI units of the values in the Osuna-Salgado dimensional format.

In SED applications, the Spectrum model is reused for both Spectrum and TimeSeries and is renamed Segment. The Spectrum object is expected to be generalized to a higher level Dataset object.

Each Spectrum (or Segment) may have a Length attribute giving the number of flux points in the data (in some serializations this value is deduced from the size of the data arrays, while in others it is made explicit).

Each Spectrum (or Segment) may also have a Type attribute indicating whether the data is intended as a TimeSeries (data are same spectral coord, varying times), Photometry (data are different spectral coords with irregular gaps), Spectrum (data are different spectral coords in contiguous bins), or Mixed (some mixture of the above).

This attribute is optional and defaults to Spectrum.

Segments are discussed in more detail in the SED document.

5.8 Packaging model

The simple Packaging model for SSA describes the format of the associated dataset. Allowed values for the format (detailed serialization for formats 4 to 7 to be specified in a separate document; only 1-3 are discussed here.)

These packaging values will be part of the SSA protocol response, and are implicit in the individual serializations.

- (1) FITS (standard BINTABLE for SED, to be defined)

- (2) VOTABLE

- (3) XML (native XML for web services and XML tools)

- (4) text (simple text table with columns of data and no markup)

- (5) text/html

- (6) graphics; a JPG, GIF etc. representation of the data

- (7) metadata; only the XML metadata.

6 Relationship to general VO data models

The Spectrum model involves objects addressed by the proposed VO Observation and Quantity data models. Although these models have not yet been fully worked out, we may note that a single Spectrum maps to the Observation model, which will include the Curation and Characterization objects. The Flux and the spectral coordinate entries together with their associated errors and quality will be special cases of the Quantity model, as will the simpler individual parameters. The field structure presented here is consistent with current drafts of the models.

6.1 Extensibility

The model and serializations defined in this document are extensible in the following sense:

- Future versions of the abstract (UML) model may add attributes or fields, and may deprecate the 'optional' property of existing fields.

- For the FITS serialization, implementors may add arbitrary additional keywords or table columns; readers must be able to handle files containing extra keywords and columns, and are encouraged to propagate such extra information when copying files. This permits local conventions to be layered on the basic definition.

- For the VOTABLE serialization, implementors may add arbitrary additional GROUP, PARAM or FIELDref/FIELD elements, with the restriction that the layering of existing elements should not be changed. (e.g. within the spectrum:DataID GROUP one may add a new GROUP containing newly defined PARAMs, but one may not move the existing Title PARAM inside the new group because that would change its indentation level). Readers must be able to handle files containing extra elements and are encouraged to propagate such extra information when copying files. This permits local conventions to be layered on the basic definition.

- For the XML object-based serialization, the CustomParams element at the top level of Spectrum is intended to allow extensibility and is equivalent to the ability to add a new GROUP at the top level. Future versions of the schema could use type extension to back-compatibly include the current schema as a special case, but apart from the CustomParams we have not currently provided for local extensibility within the current schema. (We could improve the schema by allowing the Group element to have arbitrary extra Param elements.)

References

Greisen, EW, Valdes F G, Calabretta M R and Allen S L 2003, in prep.,
http:www.aoc.nrao.edu/ egreisen/scs_112103.ps.

Hanisch, R., (ed)., Resource Metadata for the VO, Version 1.01, 2004 Apr 26.
http://www.ivoa.net/Documents/latest/RM.html

Derriere, S. et al (eds.), UCD, Moving to UCD 1+, 2004 Oct 26.
http://www.ivoa.net/Documents/latest/UCD.html

Part 2: XML schema serialization

7 XML schema serialization

7.1 XML schema

In the following XML schema, we implement the model fairly directly.

Within a spectrum the data points are kept together in objects called Point, with their general structure specified in a Fields object. Also, we have included a CustomParams element to allow site-specific metadata to be added.

The Coverage.Location fields have been collapsed to simple values rather than SEDCoord elements; this should perhaps be extended in a future version.

The Flux object is defined as an example of a more general SEDQuantity object, which is also used for the Sloan spectral service's redshift information.

A SED aggregation model is also included in the schema, as the top level element. This may be ignored until the SED model has been approved by IVOA.

The old Fields element is subsumed in the Char.Axis elements.

```xml
<?xml version="1.0" encoding="utf-8"?>
<xs:schema xmlns="http://www.ivoa.net/xml/Spectrum/Spectrum-0.98b.xsd"
xmlns:xs="http://www.w3.org/2001/XMLSchema"
xmlns:jxb="http://java.sun.com/xml/ns/jaxb"
xmlns:xlink="http://www.w3.org/1999/xlink"
targetNamespace="http://www.ivoa.net/xml/Spectrum/Spectrum-0.98b.xsd"
elementFormDefault="qualified" jxb:version="1.0">
<xs:import namespace="http://www.w3.org/1999/xlink" schemaLocation="http://www.ivoa.net/xml/Xlink/xlink.xsd"/>


<!-- A single segment corresponding to a spectrum or single point -->

<xs:element name="BaseSegment" type="segmentType"/>
<xs:element name="Spectrum" type="spectrumType" substitutionGroup="BaseSegment"/>
<xs:element name="Segment" type="spectrumType" substitutionGroup="BaseSegment"/>
<xs:element name="TimeSeries" type="timeSeriesType" substitutionGroup="BaseSegment"/>

<xs:complexType name="spectrumType">
<xs:complexContent mixed="false">
<xs:extension base="segmentType"/>
</xs:complexContent>
</xs:complexType>
<xs:complexType name="timeSeriesType">
<xs:complexContent mixed="false">
<xs:extension base="segmentType"/>
</xs:complexContent>
</xs:complexType>

<xs:complexType name="segmentType">
<xs:complexContent mixed="false">
<xs:extension base="Group">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="Target" type="targetType" />
<xs:element minOccurs="0" maxOccurs="1" ref="Data"/>
<xs:element minOccurs="0" maxOccurs="1" name="Char" type="characterizationType" />
<xs:element minOccurs="0" maxOccurs="1" name="CoordSys" type="coordSysType" />
<xs:element minOccurs="0" maxOccurs="1" name="Curation" type="curationType" />
<xs:element minOccurs="0" maxOccurs="1" name="DataID" type="dataIDType" />
<xs:element minOccurs="0" maxOccurs="1" name="Derived" type="derivedDataType" />
<xs:element minOccurs="0" maxOccurs="1" name="CustomParams" type="arrayOfParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Type" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Length" type="intParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="TimeSI" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="SpectralSI" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="FluxSI" type="textParamType" />
</xs:sequence>
</xs:extension>
</xs:complexContent>
</xs:complexType>


<!-- The top level element: an SED with one target and many segments -->
<xs:element name="SED" nillable="true" type="sedType" />

<xs:complexType name="sedType">
<xs:sequence>
<xs:element minOccurs="0" maxOccurs="1" name="Date" type="timeParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Target" type="targetType" />
<xs:element minOccurs="0" maxOccurs="1" name="CustomParams" type="arrayOfParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="Type" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="NSegments" type="intParamType" />
<xs:element minOccurs="0" maxOccurs="unbounded" ref="BaseSegment"/>
<xs:element minOccurs="0" maxOccurs="1" name="Creator" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="CreatorDID" type="textParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="SpectralMinWavelength" type="doubleParamType" />
<xs:element minOccurs="0" maxOccurs="1" name="SpectralMaxWavelength" type="doubleParamType" />
</xs:sequence>
</xs:complexType>
```

7.2 Instance example

```xml
<?xml version="1.0" encoding="UTF-8"?>
<Spectrum xmlns="http://www.ivoa.net/xml/Spectrum/Spectrum-0.98.xsd"
xmlns:xsi="http://www.w3.org/2001/XMLSchema-instance"
xsi:schemaLocation="http://www.ivoa.net/xml/Spectrum/Spectrum-0.98.xsd">

<!-- xml instance example -->
  <Target>
   <name>Arp 220 </name>
   <pos>233.737917 23.503330</pos>
   <redshift>0.0018</redshift>
  </Target>
  <SegmentType ucd="meta.code">Spectrum</SegmentType>
<!-- STC metadata, with references to defined fields -->
 <CoordSys id="ID000001">
        <TimeFrame>
                <Name>UTC</Name>
<ReferencePosition>BARYCENTER</ReferencePosition>
        </TimeFrame>
        <SpaceFrame>
                <Name>ICRS</Name>
<ReferencePosition>BARYCENTER</ReferencePosition>
        </SpaceFrame>
        <SpectralFrame ucd="em.wavelength">
                <Name>Wavelength</Name>
<ReferencePosition>HELIOCENTER</ReferencePosition>
        </SpectralFrame>
        <GenericCoordFrame id="FD" ucd="phot.fludens;em.wl">
                <Name>Flux density</Name>
        <GenericCoordFrame>
 </CoordSys>
```

```xml
<Characterization>
  <FluxAxis>
    <name>Flux density</name>
    <ucd>phot.flux;em.wavelength</ucd>
    <unit>erg cm**(-2) s**(-1) Angstrom**(-1)"</unit>
    <Accuracy>
      <StatErrLow  unit="erg cm**(-2) s**(-1) Angstrom**(-1)" id="Field5"/>
      <StatErrHigh unit="erg cm**(-2) s**(-1) Angstrom**(-1)" id="Field6"/>
      <SysErr>0.05</SysErr>
    </Accuracy>
  </FluxAxis>
  <SpatialAxis>
    <name>Sky</name>
    <ucd>pos.eq</ucd>
    <unit>deg</unit>
    <coordsystem ID="ID000001">
    <Coverage>
     <Location><Value ucd="pos.eq" unit="deg">132.4210 12.1232</Value> </Location>
     <Bounds>
      <Extent ucd="pos.region.diameter" unit="arcsec">20    </Extent>
     </Bounds>
    </Coverage>
    <Accuracy>
     <Calibration>CALIBRATED</Calibration>
    </Accuracy>
  </SpatialAxis>
  <TimeAxis>
     <name>Time</name>
     <ucd>time</ucd>
     <unit>d</unit>
     <coordsystem ID="ID000001">
    <Coverage>
    <Location><value ucd="time.obs" unit="d">52148.3252</value></Location>
     <Bounds>
      <Extent ucd="time.expo;phot.spectrum" unit="s">1500.0</Extent>
     </Bounds>
    </Coverage>
    <Accuracy>
     <Calibration>CALIBRATED</Calibration>
    </Accuracy>
  </TimeAxis>
  <SpectralAxis>
     <name>SpectralCoord</name>
     <ucd>em.wl</ucd>
     <unit>Angstrom</unit>
     <coordsystem ID="ID000001">
    <Coverage>
     <Bounds>
      <Extent ucd="instr.bandwidth" unit="Angstrom">3000.0</Extent>
     </Bounds>
    </Coverage>
    <Accuracy>
      <BinLow ucd="stat.min;em.wavelength" unit="Angstrom"/>
      <BinHigh ucd="stat.max;em.wavelength" unit="Angstrom"/>
      <Calibration>CALIBRATED</Calibration>
    </Accuracy>
  </SpectralAxis>
</Characterization>
```

```
<Curation>
 <Publisher ucd="meta.organization;meta.curation">SAO</Publisher>
 <PubID ucd="meta.curation.pubid">ivoa://cfa.harvard.edu</PubID>
 <Logo ucd=meta.curation.logo">http://cfa-www.harvard.edu/nvo/cfalogo.jpg</Logo>
 <Contact>
    <Name ucd="meta.human;meta.curation">Jonathan McDowell</Name>
    <Email ucd="meta.email">jcm@cfa.harvard.edu</Email>
 </Contact>
</Curation>
<DataID>
  <Title>Arp 220 SED</Title>
  <Creator ucd="meta.curation.creator">SAO/FLWO</Creator>
  <Date ucd="time;soft.dataset;meta.curation">2003-12-31T14:00:02Z</Date>
  <Version ucd="soft.dataset.version;meta.curation">1</Version>
  <Instrument ucd="inst.id">BCS</Instrument>
  <Collection>Archival</Collection>
</DataID>
<Derived>
  <SNR>3.0</SNR>
</Derived>
```

```
  <!-- Use table structure -->
  <Data>
   <Point>
    <SpectralCoord>
     <Value>3200.0</Value>
     <Accuracy><BinLow>3195.0</BinLow><BinHigh>3205.0</BinHigh></Accuracy>
    </SpectralCoord>
    <Flux>
      <Value>1.38E-12</Value>
      <Accuracy><StatErrLow>5.2E-14</StatErrLow><StatErrHigh>6.2E-14</StatErrHigh></Accuracy>
      <Quality>0</Quality>
    </Flux>
   </Point>

   <Point>
    <SpectralCoord>
     <Value>3210.5</Value>
     <Accuracy><BinLow>3205.0</BinLow><BinHigh>3216.0</BinHigh></Accuracy>
    </SpectralCoord>
    <Flux>
      <Value>1.12E-12</Value>
      <Accuracy><StatErrLow>1.12E-12</StatErrLow><StatErrHigh>0</StatErrHigh></Accuracy>
      <Quality>0</Quality>
    </Flux>
   </Point>

   <Point>
    <SpectralCoord>
     <Value>3222.0</Value>
     <Accuracy><BinLow>3216.0</BinLow><BinHigh>3228.0</BinHigh></Accuracy>
    </SpectralCoord>
    <Flux>
      <Value>1.42E-12</Value>
      <Accuracy><StatErrLow>1.3E-14</StatErrLow><StatErrHigh>0.2E-14</StatErrHigh></Accuracy>
      <Quality>3</Quality>
    </Flux>
   </Point>
...
  </Data>



OR

  <Data>
  <FlatPoint SP=3200.0 SP_BinL=3195.0 SP_BinH=3205.0
                  F=1.48E-12 F_ErrL=2.0E-14  F_ErrH=2.0E-14/>
  <FlatPoint SP=3210.0 SP_BinL=3205.0 SP_BinH=3215.0
                  F=1.48E-12 F_ErrL=3.2E-14  F_ErrH=3.8E-14/>
  <FlatPoint SP=3220.0 SP_BinL=3215.0 SP_BinH=3225.0
                  F=1.48E-12 F_ErrL=1.48E-12  F_ErrH=0.0 />
...
 </Data>
</Spectrum>
```

Part 3: VOTABLE serialization

8 VOTABLE serialization

8.1 Mapping Schema to VOTABLE

We reproduce below the XML schema instance example as a VOTABLE instance example. To go from the XML instance to the VOTABLE instance, we:

- - map the top level element to a RESOURCE

- - map all elements with simple content to PARAM

- - map all elements with complex content to GROUP

- - map the element names (with appropriate path) to values of the utype attribute,

- - but, handle the FIELDS and Data elements in a special way. The FIELDS element is used to define the table fields and the Data element is used to define the table data.

- - but, also, all the second level elements below RESOURCE except SPECTRUM map to an initial TABLE, while we map SPECTRUM to a second TABLE.

- - most of the elements extend the Param element, to which I have added an optional name attribute that I have not used in the instance. If this attribute is used, it can hold the name attributes of the PARAM and FIELD; otherwise the relevant attributes could be filled with the same value as the utype (without namespace prefix).

How can this be generalized to mapping an arbitrary data model schema to VOTABLE? The only tricky parts are

- **Spotting where the tabledata parts are.** We could require any DM schema that maps to VOTABLE to include elements called FIELDS and Data (perhaps ROWS would be a better name), otherwise you would get a VOTABLE with no data section.

- **Spotting where to start the main TABLE (i.e. the fact that SPECTRUM is special).** We could change the schema to have an explicit attribute, annotation or other marker to tell us this.

These issues will require further discussion for future models.

8.2 A VOTABLE instance

The VOTable version of Spectrum uses a single VOTable <TABLE>. (Note that this may appear as one of many tables within an SED VOTable). The data model fields described above as arrays map to VOTable FIELDs, while the remaining fields map to PARAM.

We use nested GROUP constructs to delimit data model objects within the main object, and PARAM and FIELD tags for attributes. Names of fields and parameters are left to the data provider. The utype and ucd attributes are used to denote data model and UCD tags. The schema and namespace for the utypes is the XML schema given in section 8.4. I have made up arbitrary NAME attributes for the PARAM and these are not to be considered standard; the name fields are free to be whatever the data provider wants, allowing compatibility with local archive nomenclature. The NAME attributes for the FIELD elements are also not standardized (of course they must be the same as in the matching FIELDrefs); it is the utype attribute which is standardized.

The one departure from the XML schema below is that the 'ArrayOfPoints' element and the individual 'Point' elements are implicitly represented by the table structure itself. Perhaps a UTYPE attribute to the TABLEDATA element could be used to make this explicit.

The example below describes a single SPECTRUM.

```xml
<?xml version="1.0" encoding="UTF-8"?>
<VOTABLE version="1.1"
  xmlns:xsi="http://www.w3.org/2001/XMLSchema-instance"
  xsi:noNamespaceSchemaLocation="http://www.ivoa.net/xml/VOTable/v1.1"
  xmlns:spec="http://www.ivoa.net/xml/SpectrumModel/v0.98"
  xmlns="http://www.ivoa.net/xml/VOTable/v1.1">
<RESOURCE utype="spec:Spectrum">

<TABLE utype="spec:Spectrum">
<GROUP utype="spec:Target">
 <PARAM name="Target" utype="spec:Target.Name" datatype="char" arraysize="*" value="Arp 220"/>
 <PARAM name="TargetPos" utype="spec:Target.pos" unit="deg" datatype="double"
                               arraysize="2" value="233.737917 23.503330"/>
 <PARAM name="z" utype="spec:Target.redshift" datatype="float" value="0.0018"/>
</GROUP>

<!-- SegmentType can be Photometry, TimeSeries or Spectrum -->
<PARAM name="Segtype" utype="spec:Spectrum.SegmentType" datatype="char" arraysize="*"
                                        value="Photometry" ucd="meta.code"/>
<GROUP name="CoordSys" utype="spec:Spectrum.CoordSys">
 <GROUP utype="spec:Spectrum.CoordSys.SpaceFrame">
   <PARAM name="System" utype="spec:Spectrum.CoordSys.SpaceFrame.Name" ucd="frame.pos.system"
                               datatype="char" arraysize="*" value="ICRS"/>
   <PARAM name="Equinox" utype="spec:Spectrum.CoordSys.SpaceFrame.Equinox" ucd="time.equinox;pos.eq"
                               datatype="float" value="2000.0" />
 </GROUP>
 <GROUP utype="spec:Spectrum.CoordSys.TimeFrame">
  <PARAM name="TimeFrame" utype="spec:Spectrum.CoordSys.TimeFrame.Name" ucd="frame.time.scale" datatype="char"
                               arraysize="*" value="UTC"/>
 </GROUP>
 <GROUP utype="spec:Spectrum.CoordSys.SpectralFrame">
  <PARAM name="SpectralFrame" utype="spec:Spectrum.CoordSys.SpectralFrame.RefPos" ucd="frame.em.system"
                               datatype="char" arraysize="*" value="BARYCENTER"/>
 </GROUP>
</GROUP>
</GROUP>
```

```
<GROUP utype="spec:Char">
 <GROUP utype="spec:Char.SpatialAxis">
   <PARAM name="SpatialAxisName" utype="name" ucd="pos.eq" unit="deg" value="Sky"/>
   <GROUP utype="spec:Char.Coverage">
    <GROUP utype="spec:Char.Coverage.Location">
     <PARAM name="SkyPos" utype="Value"
                                   ucd="pos.eq" unit="deg"
                                   datatype="double" arraysize="2" value="132.4210 12.1232"/>
   </GROUP>
   <GROUP utype="Bounds">
     <PARAM name="SkyExtent" utype="Extent" ucd="pos.region.diameter"
                    datatype="double" unit="arcsec" value="20"/>
   </GROUP>
  </GROUP>
 </GROUP>


 <GROUP utype="spec:Char.TimeAxis">
  <PARAM name="TimeAxisName" utype="name" ucd="time" unit="d" value="Time"/>
  <GROUP utype="Coverage">
   <GROUP utype="Location">
    <PARAM name="TimeObs" utype="Value" ucd="time.obs"
                                  datatype="double" value="52148.3252"/>
   </GROUP>
   <GROUP utype="Bounds">
    <PARAM name="TimeExtent" utype="Extent" ucd="time.expo;phot.spectrum"
                    unit="s" datatype="double" value="1500.0" />
    <GROUP utype="Range">
     <PARAM name="TimeStart" utype="Start" ucd="time" unit="s"
                                   datatype="double" value="52100.000" />
     <PARAM name="TimeStop" utype="Stop" ucd="time" unit="s"
                                   datatype="double" value="52300.000" />
    </GROUP>
   </GROUP>

  </GROUP>
 </GROUP>

 <GROUP utype="spec:Char.SpectralAxis">
  <PARAM name="SpectralAxisName" utype="name" ucd="em.wl" unit="Angstrom" value="Wavelength"/>
  <GROUP utype="Coverage">
   <GROUP utype="Bounds">
    <PARAM name="SpectralExtent" utype="Extent" ucd="instr.bandwidth"
                    unit="Angstrom" datatype="double" value="3000.0"/>
   </GROUP>
  </GROUP>
 </GROUP>
</GROUP>
```

```
<GROUP utype="spec:Spectrum.Curation">
 <PARAM name="Publisher" utype="spec:Spectrum.Curation.Publisher" ucd="meta.organization;meta.curation"
                        datatype="char" arraysize="*" value="SAO"/>
 <PARAM name="PubID" utype="spec:Spectrum.Curation.PubID" ucd="meta.curation.pubid" datatype="char"
                                arraysize="*" value="ivoa://cfa.harvard.edu"/>
 <PARAM name="Logo" utype="spec:Spectrum.Curation.Logo" ucd="meta.curation.logo" datatype="char"
                                arraysize="*" value="http://cfa-www.harvard.edu/nvo/cfalogo.jpg"/>
 <PARAM name="Contact" utype="spec:Spectrum.Curation.ContactName" ucd="meta.human;meta.curation"
                        datatype="char" arraysize="*" value="Jonathan McDowell"/>
 <PARAM name="email" utype="spec:Spectrum.Curation.ContactEmail" ucd="meta.email" datatype="char"
                                arraysize="*" value="jcm@cfa.harvard.edu"/>
</GROUP>


<GROUP utype="spec:Spectrum.DataID">
 <PARAM name="Title" utype="spec:Spectrum.DataID.Title" datatype="char" arraysize="*" value="Arp 220 SED"/>
 <PARAM name="Creator" utype="spec:Segmenet.DataID.Creator" ucd="meta.curation.creator" datatype="char"
                                arraysize="*" value="SAO/FLWO"/>
 <PARAM name="DataDate" utype="spec:Spectrum.DataID.Date" ucd="time;soft.dataset;meta.curation"
                        datatype="char" arraysize="*" value="2003-12-31T14:00:02Z"/>
 <PARAM name="Version" utype="spec:Spectrum.DataID.Version" ucd="soft.dataset.version;meta.curation"
                        datatype="char" arraysize="*" value="1"/>
 <PARAM name="Instrument" utype="spec:Spectrum.DataID.Instrument" ucd="inst.id" datatype="char"
                                arraysize="*" value="BCS"/>
 <PARAM name="Filter" utype="spec:Spectrum.DataID.Collection" ucd="inst.filter.id" datatype="char"
                                arraysize="*" value="G300"/>
 <PARAM name="CreationType" utype="spec:Spectrum.DataID.CreationType" datatype="char" arraysize="*" value="Archiv
</GROUP>

<GROUP utype="spec:Spectrum.Derived">
 <PARAM name="SNR" utype="spec:Spectrum.Derived.SNR" datatype="float" value="3.0"/>
</GROUP>

<GROUP utype="spec:Spectrum.Data">

<GROUP utype="spec:Spectrum.Data.SpectralAxis">
 <FIELDref ref="Coord"/>

 <GROUP utype="spec:Spectrum.Data.SpectralAxis.Accuracy">
  <FIELDref ref="BinLow"/>
  <FIELDref ref="BinHigh"/>
 </GROUP>
<!-- In this case Resolution is demoted from Field to Param since it is constant -->
 <PARAM name="Resolution" utype="spec:Spectrum.Data.SpectralAxis.Resolution"
                        unit="Angstrom" datatype="float" value="14.2"/>
</GROUP>
```

```
<GROUP utype="spec:Spectrum.Data.FluxAxis">
 <FIELDref ref="Flux1"/>
 <GROUP utype="spec:Spectrum.Data.FluxAxis.Accuracy">
  <FIELDref ref="ErrorLow"/>
  <FIELDref ref="ErrorHigh"/>
  <PARAM name="SysErr" utype="SysErr" unit="" datatype="float" value="0.05"/>
 </GROUP>
 <FIELDref ref="Quality"/>
</GROUP>
</GROUP>


<FIELD name="Coord" ID="Coord" utype="spec:Spectrum.Data.SpectralAxis.Value" ucd="em.wavelength"
             datatype="double" unit="Angstrom"/>
<FIELD name="BinLow" ID="BinLow" utype="spec:Spectrum.Data.SpectralAxis.BinLow"
             ucd="stat.min;em.wavelength"
             datatype="double" unit="Angstrom"/>
<FIELD name="BinHigh" ID="BinHigh" utype="spec:Spectrum.Data.SpectralAxis.BinHigh"
              ucd="stat.max;em.wavelength"
             datatype="double" unit="Angstrom"/>
<FIELD name="Flux" ID="Flux1" utype="spec:Spectrum.Data.FluxAxis.value" ucd="phot.flux;em.wavelength"
             datatype="double" unit="erg cm**(-2) s**(-1) Angstrom**(-1)"/>
<FIELD name="ErrorLow" ID="ErrorLow" utype="spec:Spectrum.Data.FluxAxis.Accuracy.StatErrLow"
             datatype="double" unit="erg cm**(-2) s**(-1) Angstrom**(-1)"/>
<FIELD name="ErrorHigh" ID="ErrorHigh" utype="spec:Spectrum.Data.FluxAxis.Accuracy.StatErrHigh"
             datatype="double" unit="erg cm**(-2) s**(-1) Angstrom**(-1)"/>
<FIELD name="Quality" ID="Quality" datatype="int" utype="spec:Spectrum.Data.FluxAxis.Quality"/>
<DATA>
<TABLEDATA>
<!-- Note slightly nonlinear wavelength solution -->
<!-- Second row is upper limit -->
<!-- Third row has quality mask set -->
<TR><TD>3200.0</TD><TD>3195.0</TD><TD>3205.0</TD><TD>1.38E-12</TD><TD>5.2E-14</TD><TD>6.2E-14</TD>
                                                    <TD>0</TD></TR>
<TR><TD>3210.5</TD><TD>3205.0</TD><TD>3216.0</TD><TD>1.12E-12</TD><TD>1.12E-12</TD>
                                              <TD>0</TD><TD>0</TD></TR>
<TR><TD>3222.0</TD><TD>3216.0</TD><TD>3228.0</TD><TD>1.42E-12</TD><TD>1.3E-14</TD>
                                              <TD>0.2E-14</TD><TD>3</TD></TR>
</TABLEDATA>
</DATA>
</TABLE>
</RESOURCE>
</VOTABLE>
```

Part 4: FITS serialization

9 FITS serialization

9.1 Mapping Spectrum to FITS

We define a reference serialization of this data model as a FITS binary table. The table represents a spectrum or photometry point as a single row of a table.

This serialization is a special case of an SED serialization which uses one row per spectral segment; in that case, variable-length arrays may be used to contain the array quantities. In each case below where a 'variable length array' is specified, fixed length arrays are suitable for a single spectrum or for multiple spectra where all the arrays are the same length, but readers should be prepared to handle the variable length case.

For SEDs, another approach would be to have one FITS HDU per spectrum or photometry point. However this was rejected as unworkable, as the overhead of 5760 bytes (2 FITS blocks) per photometry point would inflate the data for the photometry-only SED case by factors of around 50-100.

Here we give the mapping of data model fields to FITS columns and keywords. For each column, the standard keywords TTYPEn, TUNITn, TFORMn should be provided. Order of keywords and columns is not significant, except that it is strongly recommended that RA and Dec be in adjacent columns or keywords.

The SPCO_UCD and FLUX_UCD keywords give the spectral coordinate and flux UCDs for the spectrum. The units of spectral coordinate and flux are given in the TUNITn keys of the corresponding data columns. There is no separate provision for units of Char.SpectralAxis or Char.FluxAxis; these are required to be the same as for the data.

We add a new keyword VOCLASS to describe the VO object represented by the FITS table. The value of VOCLASS should be 'SPECTRUM 1.00'.

The spectral coordinate may also be identified by optional 1Sn_1 and 1CTYPn keywords as per WCS Paper 3. Table 9 of that paper implies that each data column which is a function of the spectral coord needs a pair of such keywords. Applications which implement the spectrum data model may ignore the WCS keys and interpret the file by recognizing 'by spec' that SPCO is the spectral coordinate and that FLUX, etc. are functions of it, but the WCS keys give a general FITS application a chance at making sense of the file.

Data model field	FITS keyword or column	Value if fixed
Spectrum.DataModel	VOCLASS	SPECTRUM 1.0
Spectrum.Length	SIZE	
Spectrum.Type	VOSEGT	
Spectrum.CoordSys.ID	VOCSID	
Spectrum.CoordSys.SpaceFrame.Name	RADECSYS	either ICRS or FK5
Spectrum.CoordSys.SpaceFrame.Equinox	EQUINOX	2000.0 (required)
Spectrum.CoordSys.TimeFrame.Name	TIMESYS	TT (required)
Spectrum.CoordSys.TimeFrame.Zero (?)	MJDREF	(required)
Spectrum.CoordSys.TimeFrame.RefPos		
Spectrum.CoordSys.SpectralFrame.RefPos	SPECSYS	(see Greisen et al)
Spectrum.Curation.Publisher	VOPUB	
Spectrum.Curation.Reference	VOREF	
Spectrum.Curation.PublisherID	VOPUBID	
Spectrum.Curation.Version	VOVER	
Spectrum.Curation.ContactName	CONTACT	
Spectrum.Curation.ContactEmail	EMAIL	
Spectrum.Curation.Rights	CURIGHTS	
Spectrum.Curation.Date	CUDATE	
Spectrum.Curation.PublisherDID	VOPUBDID	
Spectrum.Target.Name	OBJECT	
Spectrum.Target.Description	OBJDESC	
Spectrum.Target.Class	SRCCLASS	
Spectrum.Target.spectralClass	SPECTYPE	
Spectrum.Target.redshift	REDSHIFT	
Spectrum.Target.pos	RA_TARG, DEC_TARG	
Spectrum.Target.VarAmpl	TARGVAR	
Spectrum.DataID.Title	TITLE	
Spectrum.DataID.Creator	CREATOR	
Spectrum.DataID.Collection	COLLECTn	
Spectrum.DataID.DatasetID	DS_IDENT	
Spectrum.DataID.CreatorDID	CR_IDENT	
Spectrum.DataID.Date	DATE	
Spectrum.DataID.Version	VERSION	
Spectrum.DataID.Instrument	INSTRUME	
Spectrum.DataID.CreationType	CRETYPE	
Spectrum.DataID.Logo	VOLOGO	
Spectrum.DataID.Contributor	CONTRIBn	
Spectrum.DataID.DataSource	DSSOURCE	
Spectrum.Derived.SNR	DER_SNR	
Spectrum.Derived.redshift.value	DER_Z	
Spectrum.Derived.redshift.statError	DER_ERR	
Spectrum.Derived.redshift.Confidence	DER_CONF	
Spectrum.Derived.VarAmpl	DER_VAMP	
Spectrum.TimeSI	TIMESDIM	(required)
Spectrum.SpectralSI	SPECSDIM	(required)
Spectrum.FluxSI	FLUXSDIM	(required)

Data model field	FITS keyword or column	Value if fixed
Spectrum.Char.FluxAxis.Name	-	
Spectrum.Char.FluxAxis.Unit	-	
Spectrum.Char.FluxAxis.ucd	FLUX_UCD	
Spectrum.Char.FluxAxis.Accuracy.StatError	STAT_ERR	
Spectrum.Char.FluxAxis.Accuracy.SysError	SYS_ERR	
Spectrum.Char.FluxAxis.Accuracy.Calibration	FLUXCAL	
Spectrum.Char.FluxAxis.Quality	QUALITY	
Spectrum.Char.SpectralAxis.Name	-	
Spectrum.Char.SpectralAxis.Unit	-	
Spectrum.Char.SpectralAxis.ucd	SPCO_UCD	
Spectrum.Char.SpectralAxis.Accuracy.BinSize		
Spectrum.Char.SpectralAxis.Accuracy.StatError	SPCO_STE	
Spectrum.Char.SpectralAxis.Accuracy.SysError	SPCO_SYE	
Spectrum.Char.SpectralAxis.Accuracy.Calibration	SPECCAL	
Spectrum.Char.SpectralAxis.Resolution	SPCO_RES	
Spectrum.Char.SpectralAxis.Coverage.Location.Value	SPECBAND	
Spectrum.Char.SpectralAxis.Coverage.Bounds.Extent	SPECBWID	
Spectrum.Char.SpectralAxis.Coverage.Support.Fill	SPCO_FIL	
Spectrum.Char.TimeAxis.Name		
Spectrum.Char.TimeAxis.Unit	TIMEUNIT	(HEA convention)
Spectrum.Char.TimeAxis.ucd	-	time
Spectrum.Char.TimeAxis.Accuracy.BinSize		
Spectrum.Char.TimeAxis.Accuracy.StatError	TIME_STE	
Spectrum.Char.TimeAxis.Accuracy.SysError	TIME_SYE	
Spectrum.Char.TimeAxis.Accuracy.Calibration	TIMECAL	
Spectrum.Char.TimeAxis.Resolution	TIME_RES	
Spectrum.Char.TimeAxis.Coverage.Location.Value	DATE-OBS	
Spectrum.Char.TimeAxis.Coverage.Bounds.Extent	EXPOSURE	
Spectrum.Char.TimeAxis.Coverage.Bounds.Start	TSTART	
Spectrum.Char.TimeAxis.Coverage.Bounds.Stop	TSTOP	
Spectrum.Char.TimeAxis.Coverage.Support.Fill		
Spectrum.Char.SpatialAxis.Name	-	
Spectrum.Char.SpatialAxis.Unit	-	deg
Spectrum.Char.SpatialAxis.ucd	-	pos.eq
Spectrum.Char.SpatialAxis.Accuracy.StatErr		
Spectrum.Char.SpatialAxis.Accuracy.SysError		
Spectrum.Char.SpatialAxis.Accuracy.Calibration	SKY_CAL	
Spectrum.Char.SpatialAxis.Resolution		
Spectrum.Char.SpatialAxis.Coverage.Location.Value	RA_NOM, DEC_NOM	
Spectrum.Char.SpatialAxis.Coverage.Bounds.Extent	APERTURE	
Spectrum.Char.SpatialAxis.Coverage.Support.Area		
Spectrum.Char.SpatialAxis.Coverage.Support.Fill		

Data model field	FITS keyword or column	Value if fixed
	Per-data-point values	
Spectrum.Data.FluxAxis.Value	FLUX	
Spectrum.Data.FluxAxis.Unit	(TUNITn)	
Spectrum.Data.FluxAxis.ucd	-	(same as Char)
Spectrum.Data.FluxAxis.Accuracy.StatError		
Spectrum.Data.FluxAxis.Accuracy.StatErrLow	ERR_LO	
Spectrum.Data.FluxAxis.Accuracy.StatErrHigh	ERR_HI	
Spectrum.Data.FluxAxis.Accuracy.SysError	SYS_ERR	
Spectrum.Data.FluxAxis.Quality	QUALITY	
Spectrum.Data.SpectralAxis.Value	SPCO	
Spectrum.Data.SpectralAxis.Unit	(TUNITn)	(same as Char)
Spectrum.Data.SpectralAxis.ucd	-	(same as Char)
Spectrum.Data.SpectralAxis.Accuracy.BinSize		
Spectrum.Data.SpectralAxis.Accuracy.BinLow	SPCO_LO	
Spectrum.Data.SpectralAxis.Accuracy.BinHigh	SPCO_HI	
Spectrum.Data.SpectralAxis.Accuracy.StatError	SPEC_SE	
Spectrum.Data.SpectralAxis.Accuracy.StatErrLow	SPEC_SEL	
Spectrum.Data.SpectralAxis.Accuracy.StatErrHigh	SPEC_SEH	
Spectrum.Data.SpectralAxis.Accuracy.SysError	SPEC_SYE	
Spectrum.Data.SpectralAxis.Resolution	SPEC_RES	
Spectrum.Data.TimeAxis.Value	TIME	
Spectrum.Data.TimeAxis.Unit	(TUNITn)	(same as Char)
Spectrum.Data.TimeAxis.ucd	-	time
Spectrum.Data.TimeAxis.Accuracy.BinLow	TIME_LO	
Spectrum.Data.TimeAxis.Accuracy.BinHigh	TIME_HI	
Spectrum.Data.TimeAxis.Resolution	TIME_RES	
Spectrum.Data.TimeAxis.Accuracy.StatError	TIME_SE	
Spectrum.Data.TimeAxis.Accuracy.StatErrLow	TIME_SEL	
Spectrum.Data.TimeAxis.Accuracy.StatErrHigh	TIME_SEH	
Spectrum.Data.TimeAxis.Accuracy.SysError	TIME_SYE	
Spectrum.Data.SpatialAxis.Value	RA,DEC	
Spectrum.Data.SpatialAxis.Unit	-	deg
Spectrum.Data.SpatialAxis.ucd	-	pos.eq
Spectrum.Data.SpatialAxis.Resolution	SKY_RES	
Spectrum.Data.SpatialAxis.Accuracy.StatError	SKY_SE	
Spectrum.Data.SpatialAxis.Accuracy.SysError	SKY_SYE	
Spectrum.Data.BackgroundModel.Value	BGFLUX	
Spectrum.Data.BackgroundModel.Unit	(TUNITn)	(same as FluxAxis)
Spectrum.Data.BackgroundModel.ucd	BG_UCD	(same as FluxAxis)
Spectrum.Data.BackgroundModel.Accuracy.StatErrLow	BGERR_LO	
Spectrum.Data.BackgroundModel.Accuracy.StatErrHigh	BGERR_HI	
Spectrum.Data.BackgroundModel.Accuracy.SysError	BSYS_ERR	
Spectrum.Data.BackgroundModel.Quality	BQQUAL	

9.2 An instance example

We summarize this with a full sample FITS extension header.

```
XTENSION= 'BINTABLE'           / binary table extension
BITPIX  =                    8 / 8-bit bytes
NAXIS   =                    2 / 2-dimensional binary table
NAXIS1  =                57344 / width of table in bytes
NAXIS2  =                    1 / number of rows in table
PCOUNT  =                    0 / size of special data area
GCOUNT  =                    1 / one data group (required keyword)
TFIELDS =                    7 / number of fields in each row
EXTNAME = 'SPECTRUM '            / name of this binary table extension
VOCLASS = 'Spectrum V0.98'      / VO Data Model
CREATOR = 'MMT Archive'        / VO Creator
VOPUBID = 'ivoa://cfa.harvard.edu' / VO Publisher ID URI
DATE    = '2004-08-30T14:18:17' / Date and time of file creation
DATE-OBS= '2004-06-03T21:18:17' / Date and time of observation
SPECBAND= ' '    / SED.Bandpass
RA_NOM       =      233.72789197    / [deg] Nominal RA
DEC_NOM      =       23.49792615    / [deg] Nominal Dec
OBJECT  = 'ARP 220 '          / Source name
SRCCLASS= 'Galaxy'             /
SPECTYPE= 'ULIRG'              /
RA_TARG      =      233.73791700    / Observer's specified target RA
DEC_TARG     =       23.50333300    / Observer's specified target Dec
REDSHIFT=              0.01812  / Emission redshift
DS_IDENT= 'cfa://whatever'    /
TITLE   = 'Observations of Merging Galaxies' /
VERSION =  2                     / Reprocessed 2004 Aug
VOPUB   = 'CfA Archive'        / VO Publisher authority
VOLOGO  = 'http://cfa.harvard.edu/vo/cfalogo.jpg' / VO Creator logo
SPCO_UCD= 'em.wl'                    /
FLUX_UCD= 'phot.fluDens;em.wl'      /
EQUINOX =  2.0000000000000E+03 / default
RADECSYS= 'ICRS    '            / default
TIMESYS = 'TT      '           / Time system
TIMEUNIT= 's'                  / Time unit
SPECSYS = 'TOPOCENT'            / Wavelengths are as observed
TIMESDIM= 'T'                  / Time SIDim
SPECSDIM= '10-10 L'             / Spectral SIDim
FLUXSDIM= '10+7 ML-1T-3'        / Flux SDim
MJDREF  =  0.0                 / MJD zero point for times
SPEC_RES=             5.0  / [Angstrom] Spectral resolution
SKY_RES =             1.0  / [arcsec] Spatial.Resolution

TELESCOP= 'MMT '              / Telescope
INSTRUME= 'MMT/BCS '          / Instrument
FILTER  = 'G220    '          / Grating
RA       =      233.73791     / [deg] Pointing position
DEC      =       23.50333     / [deg] Pointing position
RA_NOM  =       233.73791     / [deg] Pointing position
DEC_NOM =        23.50333     / [deg] Pointing position
APERTURE=        2.0          / [arcsec]  Aperture diameter/Slit width
TIME    = 52984.301203        / [d] MJD
EXPOSURE =  1500.015          / [s] Effective exposure time
SYS_ERR  =        0.15        / Fractional systematic error
SIZE     =           2048     / Number of data points
COMMENT  --------------------------
```

```
COMMENT   WCS Paper 3 Keywords
1S11_1  = 'SPCO'                / Column name with spectral coord
1CTYP11 = 'WAVE-TAB'           / Spectral coord is WAVE
1S12_1  = 'SPCO'                / Column name with spectral coord
1CTYP12 = 'WAVE-TAB'           / Spectral coord is WAVE
1S13_1  = 'SPCO'                / Column name with spectral coord
1CTYP13 = 'WAVE-TAB'           / Spectral coord is WAVE
1S14_1  = 'SPCO'                / Column name with spectral coord
1CTYP14 = 'WAVE-TAB'           / Spectral coord is WAVE
COMMENT   ------------------------
TTYPE1 = 'SPCO'   / Wavelength
TFORM1 = '2048E'
TUNIT1 = 'Angstrom'
TTYPE2 = 'SPCO_LO' /
TFORM2 = '2048E'
TUNIT2 = 'Angstrom'
TTYPE3 = 'SPCO_HI' /
TFORM3 = '2048E'
TUNIT3 = 'Angstrom'
TTYPE4 = 'FLUX' /
TFORM4 = '2048E'
TUNIT4 = 'erg cm**(-2) s**(-1) Angstrom**(-1)'
TTYPE5 = 'ERR_LO' /
TFORM5 = '2048E'
TUNIT5 ='erg cm**(-2) s**(-1) Angstrom**(-1)'
TTYPE6 = 'ERR_HI' /
TFORM6 = '2048E'
TUNIT6 ='erg cm**(-2) s**(-1) Angstrom**(-1)'
TTYPE7 = 'QUALITY' /
TFORM7 = '2048I'
```

The data would look like

```
SPCO    SPCO_LO SPOCO_HI FLUX   ERR_LO  ERR_HI  QUALITY
3200.0 3195.0 3205.0 1.48E-12 2.0E-14 2.0E-14   0
3210.0 3205.0 3215.0 1.52E-12 3.0E-14 3.0E-14   0
3220.0 3215.0 3225.0 0.38E-12 0.38E-12 0.0      0
3230.0 3225.0 3235.0 1.62E-12 3.0E-14 3.0E-14   0
```