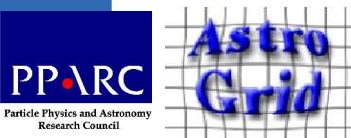


Radio data access: on-demand interferometry imaging

Anita Richards and members of MERLIN, RadioNet and AstroGrid teams.



- **'Old' radio archive access**
 - Source lists, fixed images or visibility data
 - Most maps only inner few arcsec, fixed resolⁿ
- **MERLIN archive on-demand imaging**
 - via AstroGrid, using ParseITongue
- **Next-generation interferometry**
 - Archive development strategies
 - VO requirements
- ***Enable all astronomers to use multi- λ data***

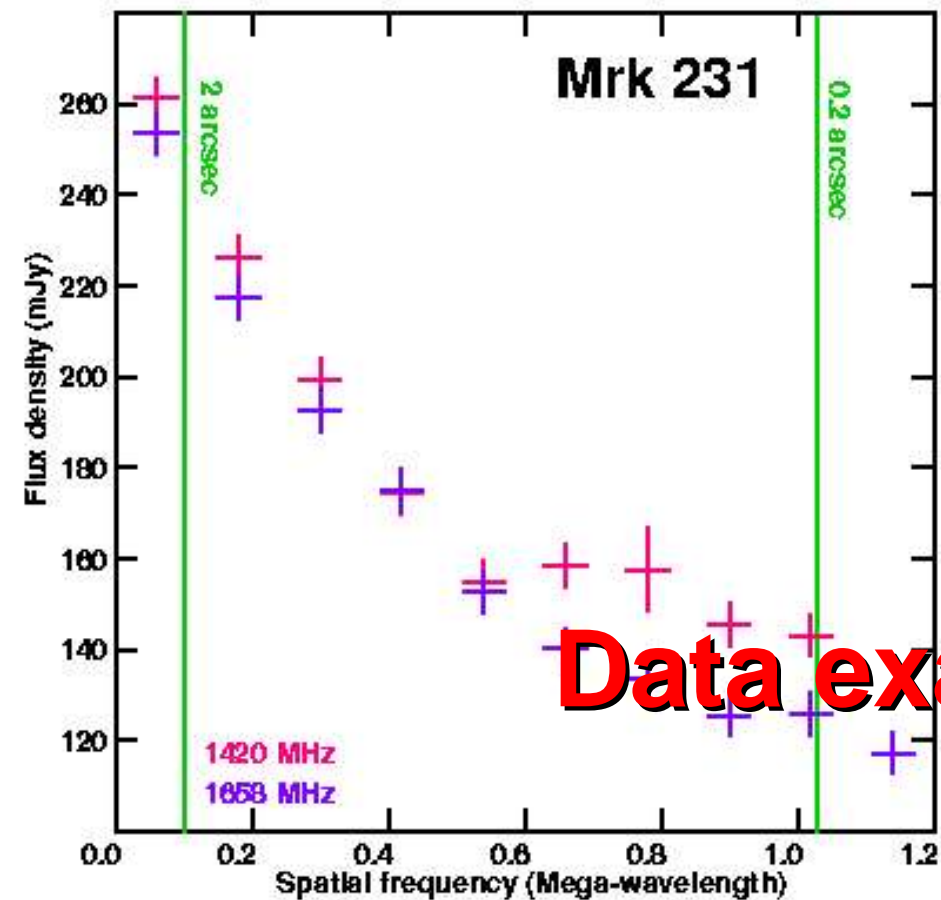
Interferometry data

- ★ GB – TB raw data per experiment (mins - weeks)
- ✿ Visibility data
 - ★ Cal sources. antennas etc.
 - ✿ Domain-specific software
- ★ No unique product
- ✿ Resolution v. sensitivity tradeoff by baseline weighting
- ✿ 3D/polarisation cubes, spectra, time series...
- ★ Field of view e.g. 8', image pixels 10 mas
- ✿ Map out regions of interest, usually not all 10^9 pixels
- ✿ Most PI results too varied to archive
- ★ Pipeline data for VOs
 - ✿ Quick Look static images
 - ✿ Flexible products to order
 - ✿ Raw data option rarely wanted if alternatives!

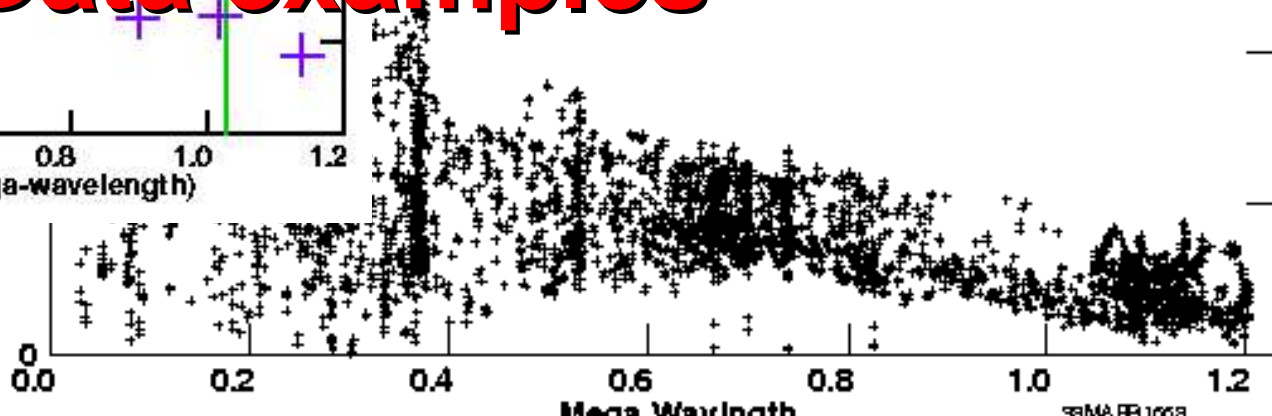
Calibrated,
binned visibility
amplitudes

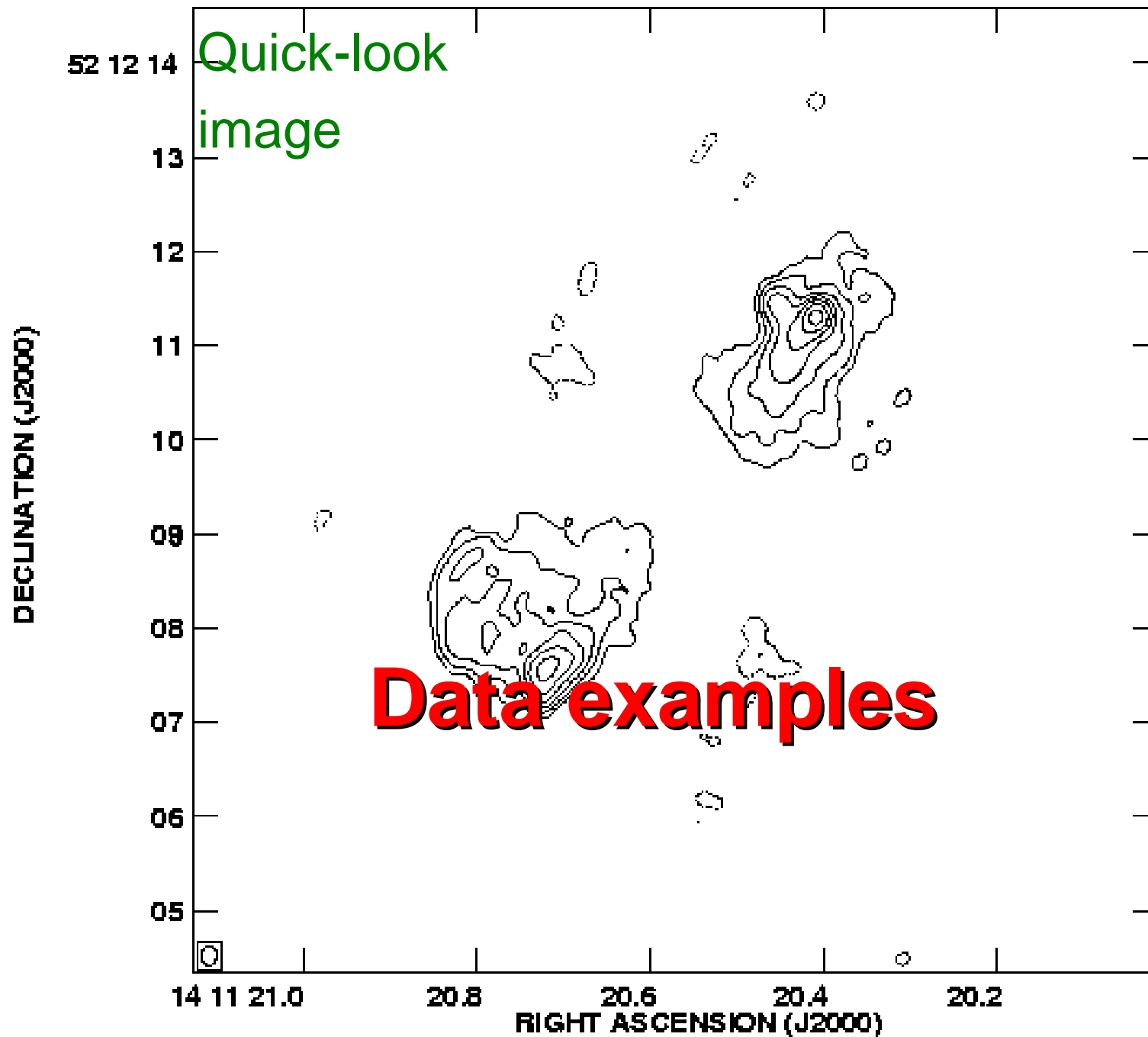
3C295 MERLIN 1658 MHz Mar B 1998 1 created 30-APR-2001 19:54:38
Amplitude vs UV dist for 3C295.UVSRT.1 Source:3C295
Ants * - * Stokes I IF# 1 Chan# 1

Interferometry
visibility
amplitude v.
baseline length



Data examples

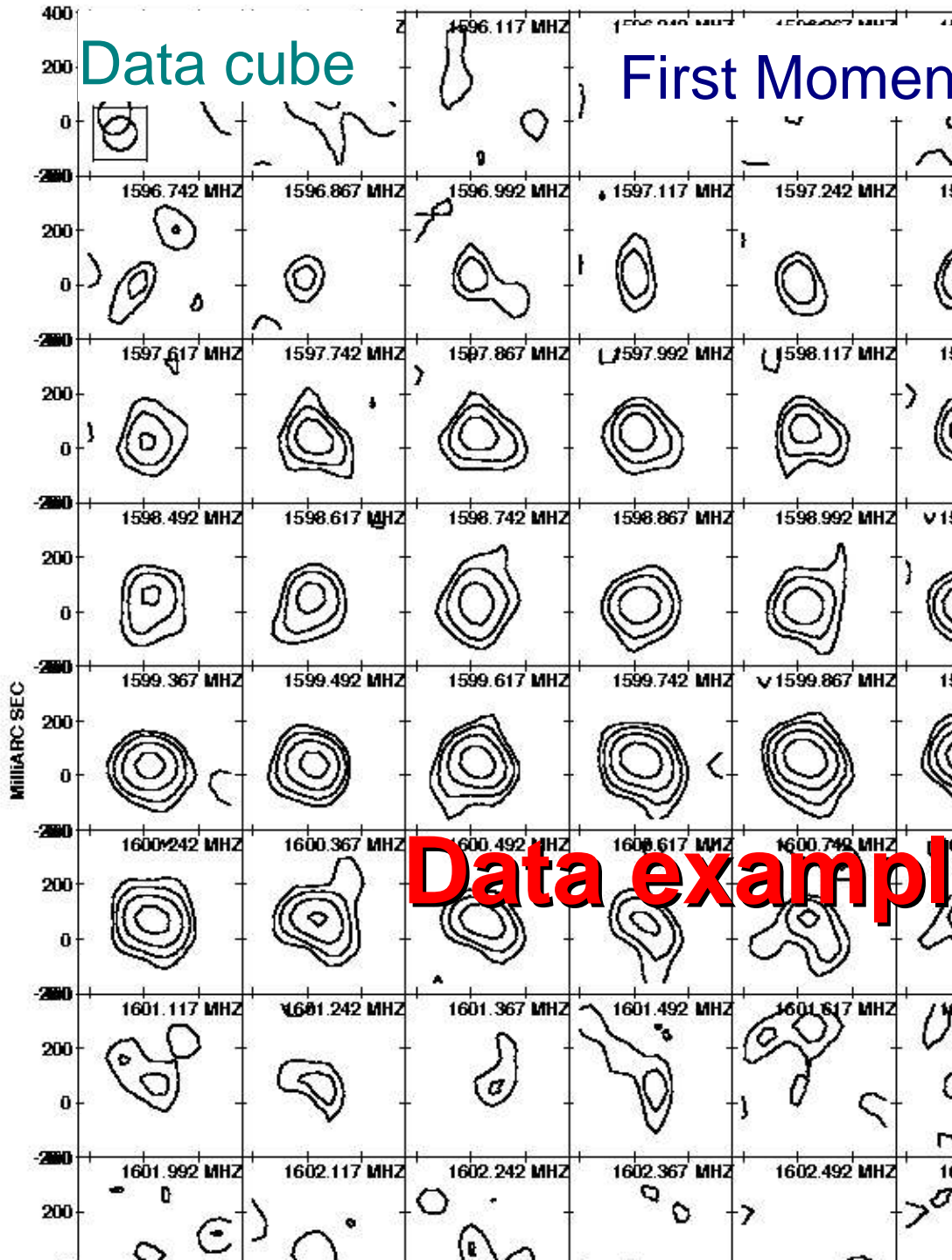




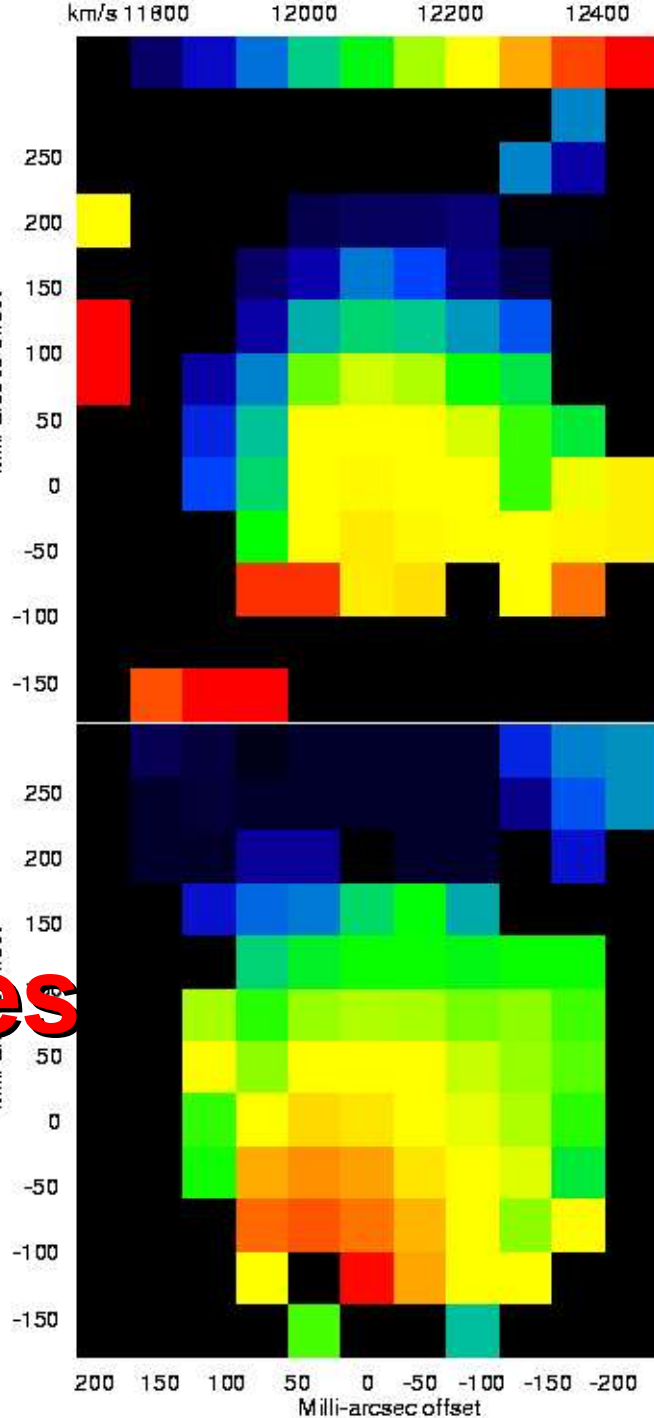
Data cube

First Moment

Janskys



Data examples



★ Non-expert access pre-2006

★ Ready-made FITS images

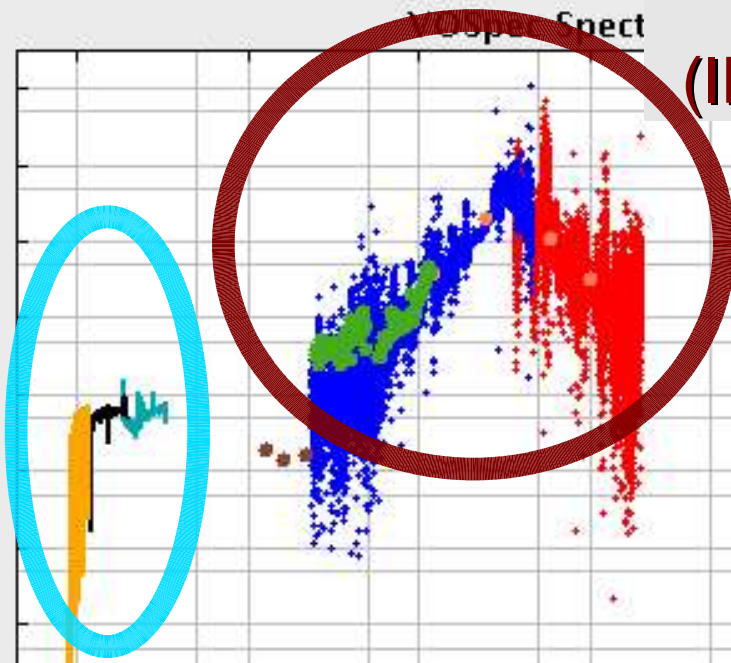
★ Aladin

★ Catalogues:

★ SpecFind

Target Ra D

IR from warm dust in remnant AGB shell (IRAS, ISO, 2MASS)



Hot stellar emission
FUSE, IUE

Non-thermal nebular radiocontinuum (various - Vollmer catalogue)

- Lines
- Dots
- Dots
- Dots
- Dots
- Dots

	Server	Title	Ra	Dec	Format	Select	Status
	Far Ultravi...	IC418 FUSE (IAP)	81.87	-12.6967	spectrum...	<input checked="" type="checkbox"/>	compl...
	Local File	Radio data: IC 418			spectrum...	<input checked="" type="checkbox"/>	compl...
	Local File	IRAS Photometry Data: IC ...			spectrum...	<input checked="" type="checkbox"/>	compl...
	Local File	ISO PHT IC 418			spectrum...	<input checked="" type="checkbox"/>	compl...

View


VLBI data – then and now

C20:
Search list by
experiment code

DIRAwww – SeaMonkey

http://db.ira.inaf.it/evn/

Back Forward Reload Stop Search




**EVN
Catalogues
Info**

**Working
Directory**

Select

**Select by
center**



NAME	ON	ID	RA1950	DEC1950	PI	AFFL	EVNPN	USPN	WL
0004+139		Q	0 4 23.500	13 59 12.000	Gurvits	Arecibo	GG14		6
0004+139		Q	0 4 23.500	13 59 12.000	Gurvits	Arecibo	GG14		6
0008+106	IIIZw2	G	0 8 0.000	10 42 0.000	de Waard	Leiden	83-09		6
0008+260			0 8 0.851	26 1 35.420	Lestrade	JPL	90-19G	L49G	6
0010+775			0 10 22.260	77 32 5.690	Polatidis	NRAL	GP7		18
0010+775			0 10 22.260	77 32 5.690	Polatidis	NRAL	GP7		18
0010+405		Q	0 10 54.200	40 35 1.000	Readhead	CIT	89-74G	R51G	6
0010+405		Q	0 10 54.200	40 35 1.000	Readhead	CIT	89-74G	R51G	6
0010+405		Q	0 10 54.300	40 34 55.810	Akujor	MPIFR	89-07G	A25G	6
0010+405		Q	0 10 54.300	40 34 55.810	Akujor	MPIFR	89-07G	A25G	6
0010+405		Q	0 10 54.342	40 34 56.310	Readhead	CIT	89-74G	R51G	18
0010+405		Q	0 10 54.342	40 34 56.310	Readhead	CIT	89-74G	R51G	18
0014+813			0 14 4.100	81 18 28.400	Kuehr	MPIFR	83-29G	K15G	6
0014+813		Q	0 14 4.460	81 18 28.500	Wehrle	CIT	90-18G	W59G	18
0014+813		Q	0 14 4.460	81 18 28.500	Wehrle	CIT	90-18G	W59G	18
0014+813		Q	0 14 4.462	81 18 28.680	Wilkinson	NRAL	GW7		6
0014+813			0 14 4.462	81 18 28.680	Vermeulen	CIT	GV014		6
0014+813		Q	0 14 4.462	81 18 28.680	Wilkinson	NRAL	GW7		6
0017+257	4C25.01	Q	0 17 3.415	25 46 13.500	Kapahi	JPL	89-21G	K25G	18
0017+257	4C25.01	Q	0 17 3.415	25 46 13.500	Kapahi	JPL	89-21G	K25G	18
0016+731		Q	0 16 54.100	73 10 52.000	Schalinsky	MPIFR	87-15G	S71G	6
0016+731		Q	0 16 54.100	73 10 52.000	Schalinsky	MPIFR	87-15G	S71G	6

Done

VLBI data – then and now

JIVE - Mozilla



EVN Data Archive at JIVE

C21: The EVN Data Archive at JIVE contains correlated data as processed at JIVE. The archive includes a growing database of data entered the public domain.

Friendly search In addition, the archive provides various correlator and pipeline products. Various correlator and pipeline products are also available. The archive allows these to be accessed in a natural way.

Pipeline products

VO links

Access to EVN archive

- [Show experiment RAH2](#)
- [Show catalogue experiment RAH2](#)
- [Search archive source or position](#)
- [The EVN archive](#)

Access to VO archives

- [Aladin Sky Atlas](#) (EVN image address)
- [Sloan Digital Sky Survey](#) (Image is available)



Ra	Dec	Source	Image	Image
250.7450	39.8103	3C345	sdss	evn
291.7001	11.3546	IRC10420		
293.3409	15.0796	J1933+1504		evn
343.4906	16.1482	3C454.3		evn

NAME	ON
0004+139	
0004+139	
0008+106	IIIZw2
0008+260	
0010+775	
0010+775	
0010+405	
0010+405	
0010+405	
0010+405	
0010+405	
0010+405	
0010+405	
0010+405	
0014+813	
0014+813	
0014+813	
0014+813	
0014+813	
0014+813	
0014+813	
0017+257	4C25.01
0017+257	4C25.01
0016+731	
0016+731	

MERLIN Archive Visibility Data Retrieval Form - Mozilla

MERLIN Archive Visibility Data Retrieval Form

Release 2.1 Jan 2006

Use this form to extract automatically processed visibility data for your observations. At present, most L- and C-band continuum data (around 1.4 and 5 GHz) taken between 1993-1999 inclusive is available and a calibration solution will be retrieved as a multi-source FITS file (SPLAT) exported from AIPS for the field requested in each processing block (up to one month long). Up to five data sets will be supplied per query. Associated calibration sources and solution, flag etc. tables will be included. The data for the field alone will also be written out as a single-source, single channel FITS file with any relevant calibration applied. See the archive query form for the source originally observed and the MERLIN User Guide for details of data products and other options. Please use the [acknowledgement](#) in any publication.

For further details contact merlin_archive@jb.man.ac.uk
 This is an experimental service so feedback and reports of any problems are welcome.

After submitting the complete form, you will be able to download your data. This data retrieval may take 2 min - hours depending on the amount of data requested. You can cancel the data retrieval at any time.

MERLIN 2005

Web forms: static images

Calibrated visibility data

Obligatory:

Either

R.A. and Dec. Enter in format 23:55:03.3820 +28:38:01.050 (sexagesimal) or 358.7640917 28.633625 (degrees)

in Equinox Enter J for J2000 or B for B1950 (default J)

or

Source name *If used, must be exact name used for MERLIN observations*

Optional (leave defaults if wished, do not leave blank)

Field size (arcsec) If you request a field of view <21 arcsec diameter at a position closer than 11 arcsec to the original pointing centre, data will be squashed to a single frequency channel; for wider fields multi-channel data will be supplied. The maximum MERLIN field of view is a few tens of arcmin, depending on frequency and configuration.

Frequency and frequency tolerance Enter in format 1420.0 100.0 (MHz) or leave default for all available bands: 1370-1430 MHz or 1550-1730 MHz (L-band) 4500-5200 MHz or 6000-7000 MHz (C-band)

Start and end times Use format YYYYMMDD or leave default for all epochs (most L- and C-band continuum data from 1993-1999 inclusive is available and an increasing amount outside those dates).

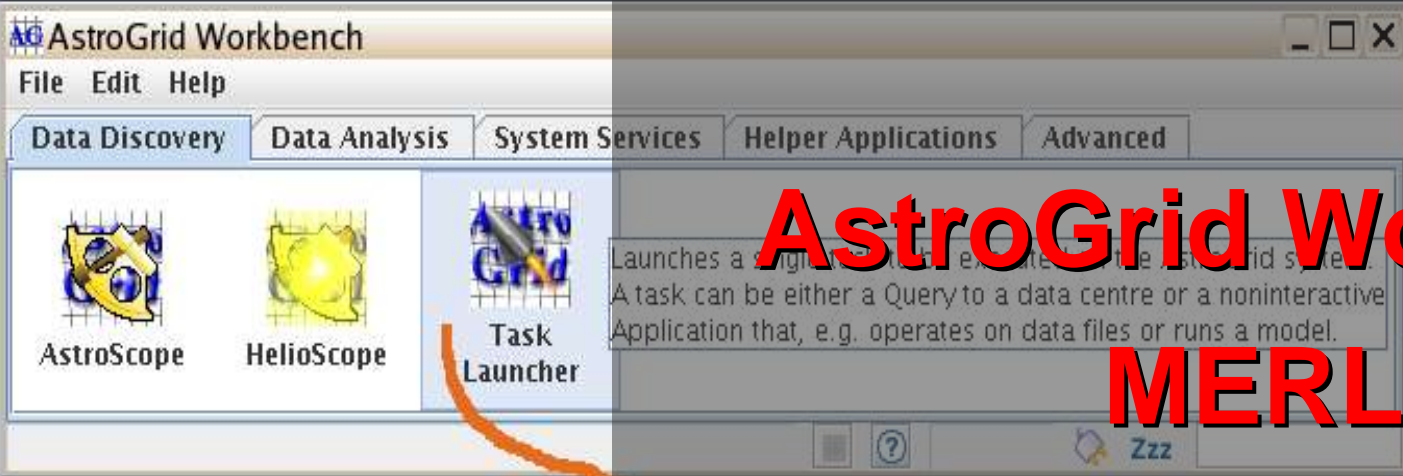
- ★ Visibilities live on disc, calibrated
- ★ Sample products
- ★ User version of pipeline
- ☼ Still need AIPS

What's wrong with that?

- ★ Dedicated web site fine for PIs of major projects
 - ★ Able to use domain-specific projects
 - ★ For simple observations, even PIs use pipeline images!
 - ✱ Not really friendly for non-experts
 - ✱ Not convenient for multi-wavelength
 - ✱ Can't be used easily in workflows
- ★ Need VO access
 - ✱ Databases of observations and metadata
 - ✱ Static quick-look products
 - ✱ Flexible data products
- ★ 2003-4 ATCA/CSIRO prototype imager using VO interferometry data model

2006

AstroGrid Workbench MERLIN Imager



*Simple
inputs
Name/
Pos
Pos
(Size
ResIn
Freq
Date)*

Task Launcher - MERLIN Imager

File

Created by Astrogrid

Chooser

Query

Parameter

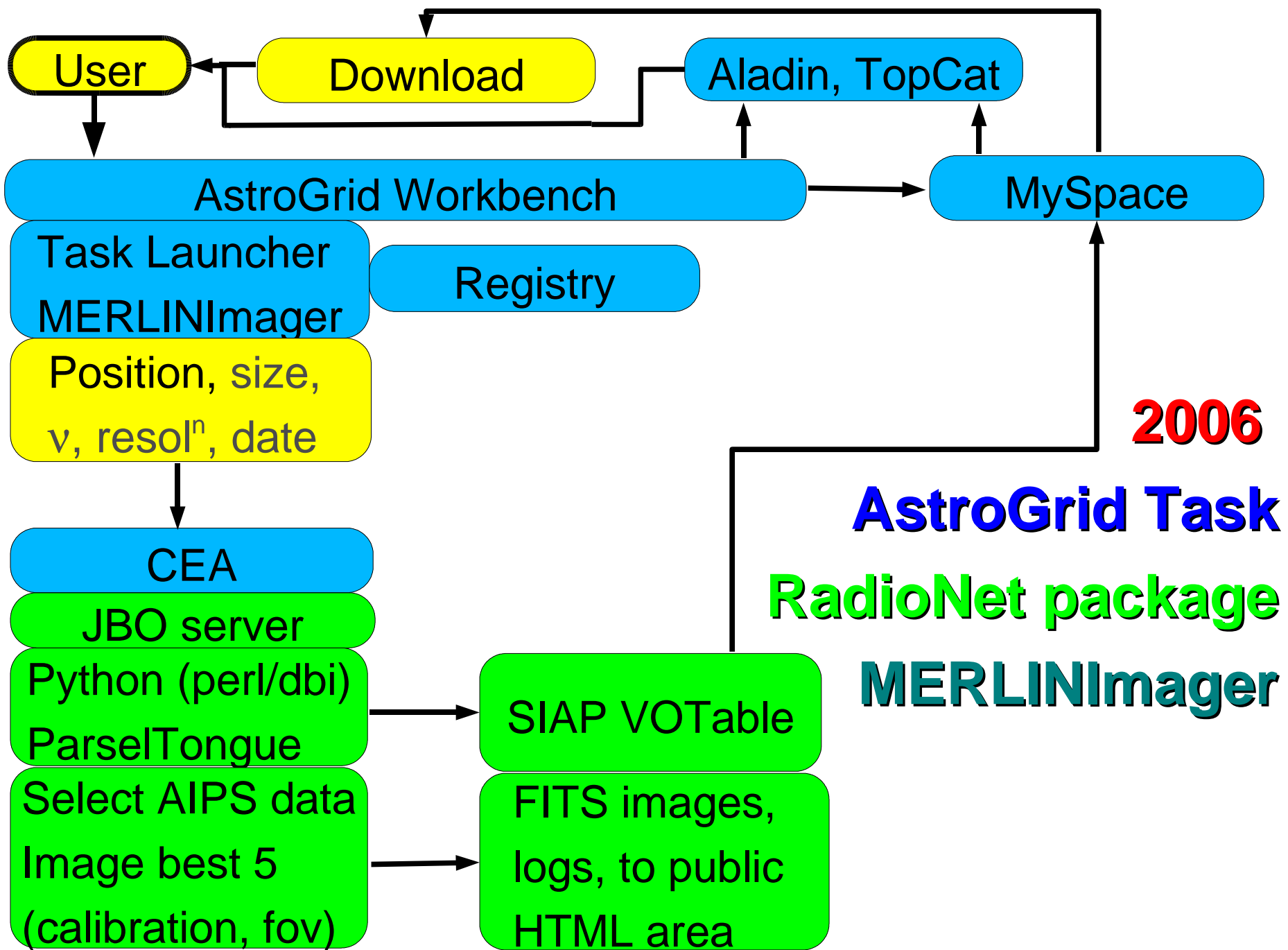
XML

Info

Security

Inputs		Name	Value
RA	0.0					
Dec	0.0					
Equinox	J					
MERLIN name	HMSGE					
Frequency	5000					
Ftol	5000					
Image Size	1					
Resolution	0.1					
Start	19910101					
End	20201212					

Outputs:		Name	Value	Ref?	Rep?	Del?
MERLIN Images List	ivo://uk.ac.le.star/anitarichards/MERLIN/HMSGE.vot			<input checked="" type="checkbox"/>		



Interoperability (Plastic)

★ Matching resolution

★ Spectral index image

Aladin v3.6 multiview *** PROTOTYPE VERSION (based on v3.611) ***

Buttons: Load... Save... Tools... PLASTIC... Print...

Position x,y Pixel 8 bits

Left panel: **18 cm** image with scale 0.2" x 1.464" x 1.56"

Right panel: **6 cm** image with scale 0.5" x 2.441" x 2.601"

Bottom left: **Steep spectrum** plot with scale 0.05" x 0.366" x 0.391"

Bottom text: [View A1] - .C.92NOVA4993_MKN273_206.175p55.887.ICLN.FITS

Footer: (c)1994-2006 ULP/CNRS - Centre de Données astronomiques de Strasbourg

MySpace Resource Chooser

Local Disk Myspace URL

- home
 - Query
 - Output
 - Workflow
 - Config
 - votable
 - CCut_RA56.75Dec23.867_1.0
 - TaskQuery
 - MERLIN

Resource URI: ivo://uk.ac.le.star/anitarichards#MERLIN/MKN273.vot

Buttons: OK Cancel

Server selector

Others: File all VO FOV MySpace Registry

Access to AstroGrid MySpace

Specify a MySpace resource identifier and press the SUBMIT button

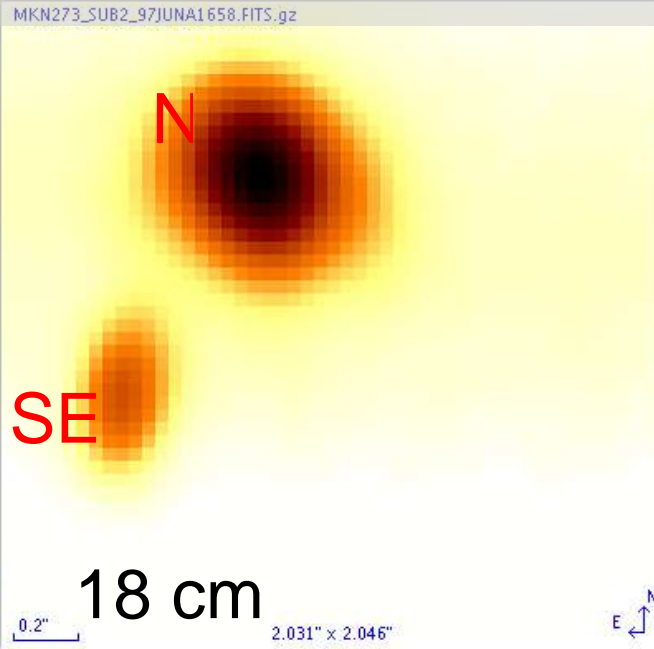
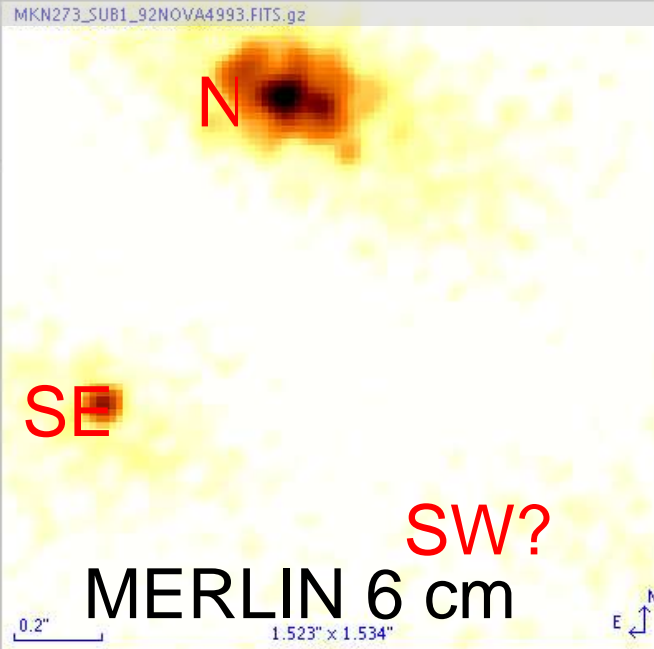
ivo://uk.ac.le.star/anitarichards#MERLIN/MKN273.vot

C	92NOVA4993_MKN273_206.175p55.887.ICLN.FITS	24.6	"x24.6
L	97JUNA1658_MKN273_206.175p55.887.ICLN.FITS	1.4	'x1.4'

Buttons: Reset Clear History SUBMIT Close

Steep spectrum
– SE not AGN

Hunt-the-AGN in Markarian 273



Compare:
old static
fixed
resolution
images



Server selector

Choose an image server or a data server and fill in the associated form drawn below

Image servers:

- Aladin server
- SSS...
- SkyView
- SDSS
- VLA...
- MAST
- Others...

Missions in VizieR

Specify a target and a mission to get the point so Select them in the view window to get the associa

Target: Markarin 273 Get all column

Mission: MERLIN Radius: 0.5

For your help, you can pick up mission names from:

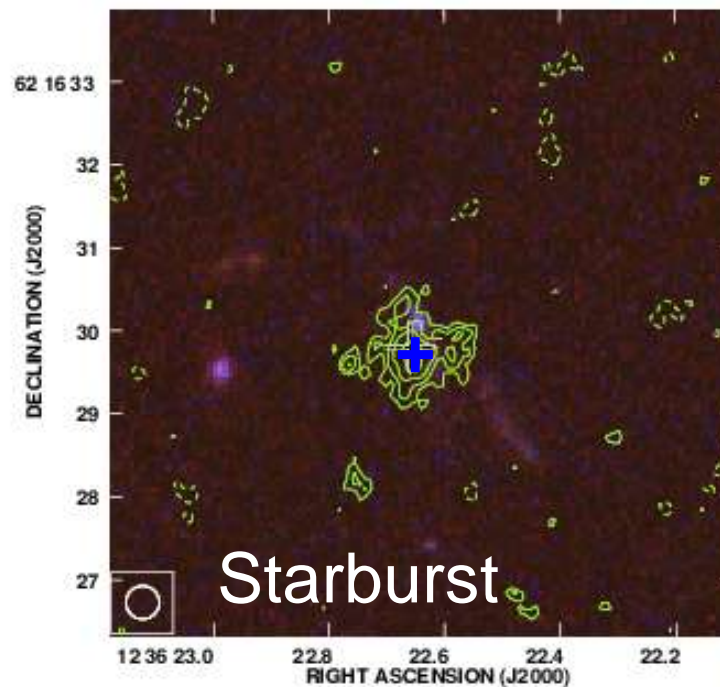
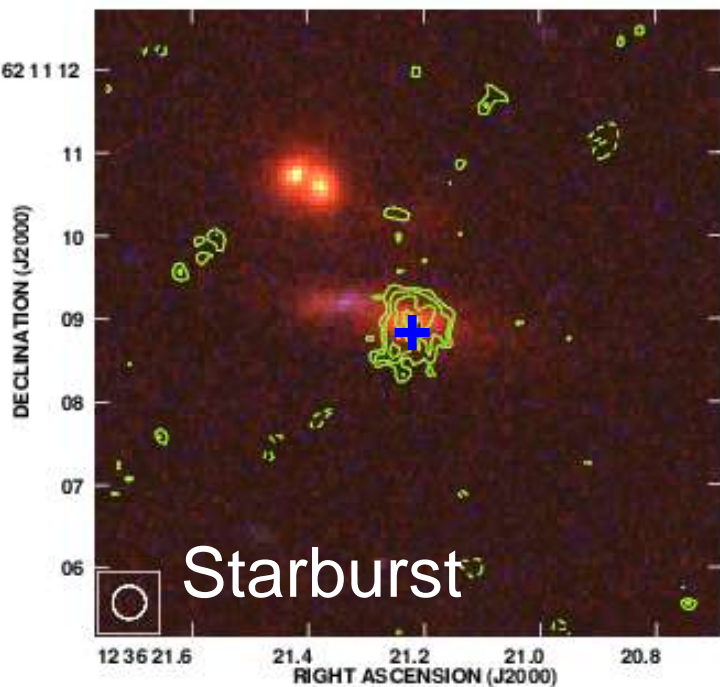
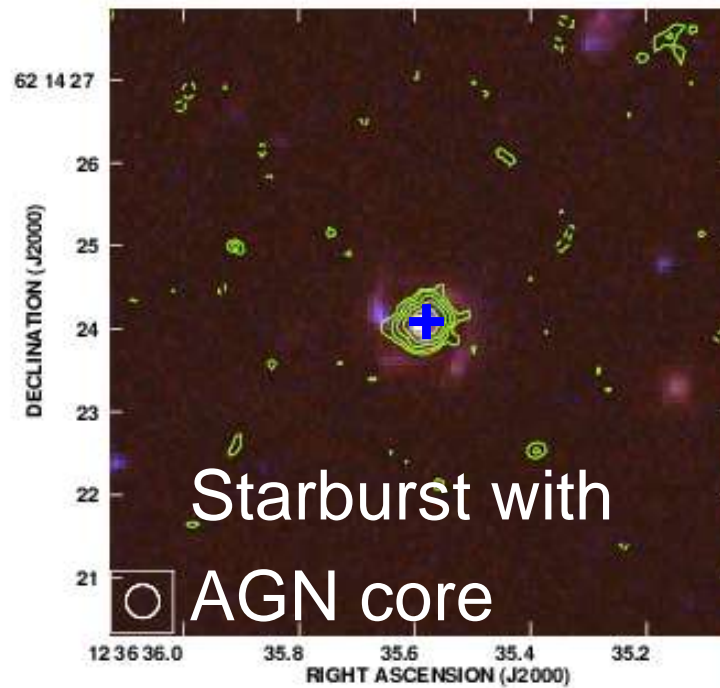
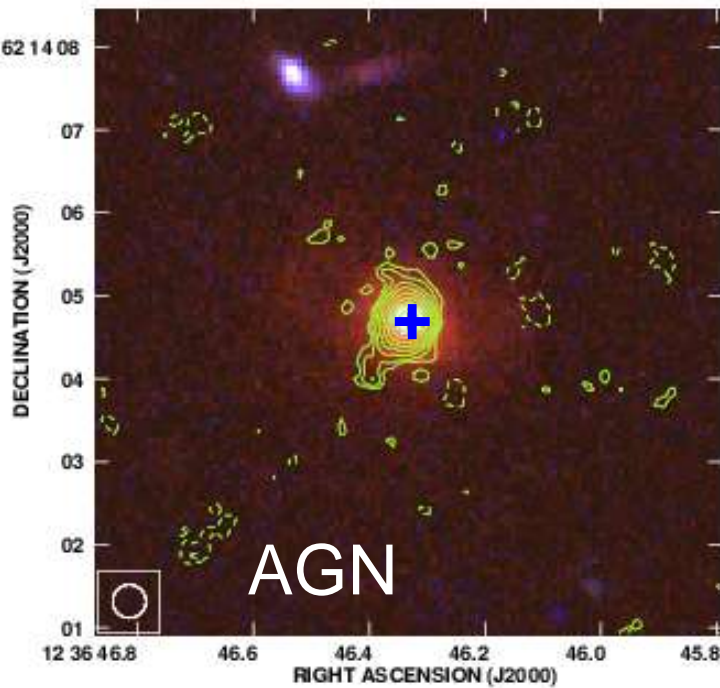
- FUSE - Log of FUSE (Far-UV Spectroscopic Explorer) -
- HST - Hubble Space Telescope Archives - 384Krow
- HUT - Log of Hopkins Ultraviolet Telescope - 1Krow
- INT - The Isaac Newton Telescope observation log - 8
- ISO - Log of ISO validated data - 23Krow
- IUE - IUE Ultraviolet Spectra - 109Krow
- MERLIN - The MERLIN interferometer at Jodrell Bank**

Simbad

NED

Others..

SDSS.



**Coming soon:
M+VLA
HDF(N)
cutouts**

1"

10 μ Jy
contours

+ X-ray
selected
AGN2

$1 < z < 2.5$

R. Beswick, S.T. Garrington, A.J. Holloway, T.W.B. Muxlow, **A.M.S. Richards**, H. Thrall, N. Winstanley (JBO);
 M.A. Garrett, M. Kattenis, H.J. van Langevelde (JIVE); E. Gonzalez-Solarez, N.A. Walton (IoA)
 P.A. Harrison, P. Padovani, P. Rosati (ESO); M.G. Allen (CDS)

1 Introduction and Data

We use Virtual Observatory methods to investigate the strong association between the presence of radio and X-ray emission at high redshifts.

A region of 100 arcmin² around the Hubble Deep Field North (HDFN 10-arcmin field) was observed by the UK radio interferometer MERLIN and the USA VLA [13]. 92 sources brighter than 40 μ Jy were found in the combined images, which reach a noise level of $< 4 \mu$ Jy using beam sizes of $0''.2 - 2''$. These are the only observations apart from the HST images (this background shows the UDF) to resolve almost all the objects detected. There is a close relationship between radio and FIR emission of starburst origin which extends to $z \sim 3$ [10, 12]. Many of the higher-redshift radio sources have sub-mm SCUBA counterparts (e.g. [16, 7, 8]), corresponding to the rest-frame FIR peak. 55 of the radio sources are among the 100 X-ray sources detected by Chandra in the 10-arcmin field [4]. 18 of these are among the obscured AGN identified by [14]; the radio emission is starburst dominated in at least 11 cases.

We employed data access, crossmatching and manipulation tools now available via the AstroGrid Workbench [1] and the EuroVO [2]. These include the Vizier and Aladin services, TopCat and a cut-out extractor for radio images which uses the parseTongue scripting interface developed by RadioNet [3]. We compare the radio data with the Chandra CDF(N) source list and other GOODS [11] HDFN observations to investigate the properties of active galaxies at $z \gg 1$:

1. Can radio emission originate from a process different from the X-ray source in the same galaxy? If so, how much radiation has a common origin and how much is separate?
2. Are high-redshift starbursts just analogues of local ULIRGs or are they a distinct phenomenon only seen in the early universe?

2 Classification of radio+X-ray sources

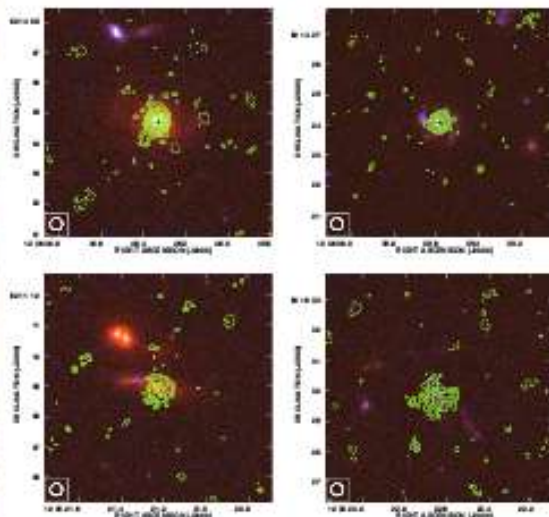


Fig. 1 Radio contours at $(-1, 1, 1.4, 2.1) \times 10 \mu$ Jy/beam overlying astrometrically aligned ACS images in the HDFN. Black crosses show the Chandra source positions; all are X-ray type-2 AGN.

Fig. 1. illustrates the radio classification criteria developed by [13]. J123646+621404 ($z = 0.96$) is a bright AGN with a compact core overlying the optical peak and lobes extending outside the host. J123635+621424 ($z = 2.01$) shows a radio AGN core within an extended starburst and significant X-ray emission. J123621+621100 ($z = 2.0$)

3 High-redshift radio starbursts host X-ray-selected AGN

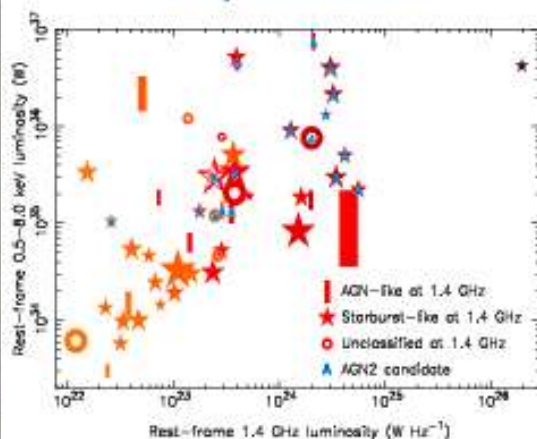


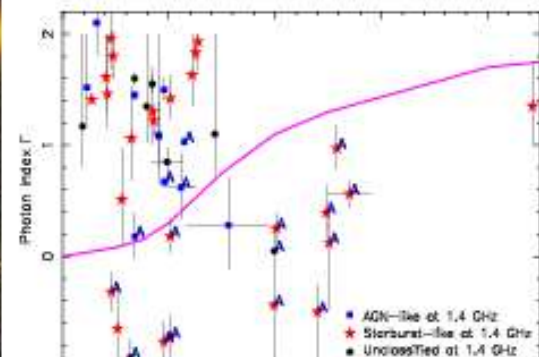
Fig. 3 Resolved radio-X-ray sources in the HDFN. Symbol height is proportional to the radio source largest angular size which ranges from $0''.3 - 12''$. Symbol shade from orange to dark red represents the redshift range 0.2-4.4.

Fig. 3. shows the classifications and luminosities of 1.4-GHz radio sources in the HDFN with X-ray counterparts and measured redshifts. 3 out of every 4 μ Jy sources are starbursts, including those with X-ray counterparts. Over half of these X-ray sources have $L_{0.5-8} > 10^{22}$ W (the unshaded area in Fig. 4.), indicating an AGN, 17 of which are classified as type-2 AGN. [14]. These are shown by a blue Δ in Figs. 3., 4. and 5.

5 A distinct population of high-z starburst AGN hosts

At $z < 1.1$, under 1/3 X-ray sources in the HDFN have radio counterparts. This rises to over 1/2 at higher redshift and in most of these the high X-ray luminosity indicates the presence of an AGN. The $L_{0.5-8} - L_{1.4}$ relationship for starbursts at $z < 1.3$ [6] shown by the green line in Fig. 4 breaks down at higher z where we see a dramatic increase in the scatter. This is the case even if only the soft-band or the de-absorbed hard-band X-ray luminosity is considered.

The mean angular size of the radio sources in the HDFN is $1''.3$ (Fig. 3), corresponding to a typical extent of 8-10 kpc for starbursts; all of which are larger than 2 kpc. The inferred star-formation rates are 1-2000 $M_{\odot} \text{ yr}^{-1}$ for the higher-redshift sources. In contrast, local ULIRGs, with FR luminosities $\geq 10^{22} \text{ W m}^{-2}$, have star-formation rates typically 50-200 $M_{\odot} \text{ yr}^{-1}$, concentrated inside the inner kpc of the galaxy. The near and far starbursts do have one feature in common; their optical counterparts very frequently show signs of severe disturbance and merger.



Next Generation Interferometers

- ★ e-MERLIN, eVBLI, EVLA
ATCA/LBA, LOFAR,
GMRT, ALMA, SKA etc
- ✿ Committed to VO access
- ✿ Software development
 - ★ RadioNet – ParselTongue
(Python over e.g. AIPS)
 - ★ NRAO – CASA (aips++
based, Python scripting).
 - ★ ESO – XML binary and
metadata storage
- ★ 'B2E' obs-to-archive
integrated data path
- ✿ Metadata retrieval
- ✿ Flexible remote
processing of products
- ★ VO components in
archive management?
- ✿ Cornish, GMRT, MMB
and other surveys

Char DM Markarian 273 @ 18 cm

General	Spatial	Temporal	Spectral	Observable
frame/units	ICRF, deg	ISOstring	Mhz	Jy
Location	13.123456 +55.987654	1997-06-14 T12:00:00	1658	undef
Bounds	12.92, +55.58 13.32, +56.38	...T06:00:00 ..T18:00:00	1650 1665	0.0002 0.5
Support	13.123456 +55.987654 0.4	(on-source scan listing URL)	1650 1665	undef
Sensitivity	$f(\tau_{\text{int}}, \delta\nu, 1\text{ary beam})$	undef	(bandpass LUT URL)	? dynamic range <5E+4
Filling Factor	1	0.7	1	undef
Resolution	0" 2 2".0 0" 2 2".0	PT00H05M	1000 kHz	5E-5 10E-6
Sampling	0" 04 0".0625 0" 04 0".0625	PT16S	1000 kHz	undef

Development

- ★ Expand/apply VO standards
 - ✻ Spatial frequencies, polarization...
 - ✻ Tools for radio data (Jy/beam, 3+D cubes...)
 - ✻ Authentication
- ★ Expand data products/services available
 - ✻ Robustness
 - ✻ Visibility data, spectra, time series
 - ✻ Single-dish data
 - ✻ Image cutouts for Deep Fields
 - ✻ Solve queuing/overload and other problems
- ★ Impliment at other sites/other radio datasets

AstroGrid/RadioNet workshop

- ◆ Radio data management (1400 5 Dec – 1600 8 Dec)
- ◆ Workshop for data providers/large surveys etc.
 - Data flow using archives and pipelines
 - ◆ ParseITongue, Common Proposal Tool etc.
 - Data delivery
 - ◆ Publishing data to VOs
 - ◆ Use and development of relevant VO tools
- ◆ Science use (0900 4 Dec – 1300 5 Dec)

Oxford w/c 4 December 2006

radiovo@jb.man.ac.uk

<http://wiki.astrogrid.org>