

Photometry DM 1.1: VO/DML Update

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Photometry DM 1.0

Recommendation since
5 October 2013

Implementations

- Filter Profile Service
- Cone Search with Photometry Extension
- Use of utypes in several TAP servers

Pending

Needed VO/DML view in order to allow proper annotation



Astronomical Photometry

Pros

- Most of the astronomical photometry shown as relative magnitude
- Easy to calibrate using, e.g. calibration stars like Vega

Cons

- Creation of SEDs is quite complex as magnitudes have meaning within the project
- Use of standard photometric systems or photometric filters is not totally accurate

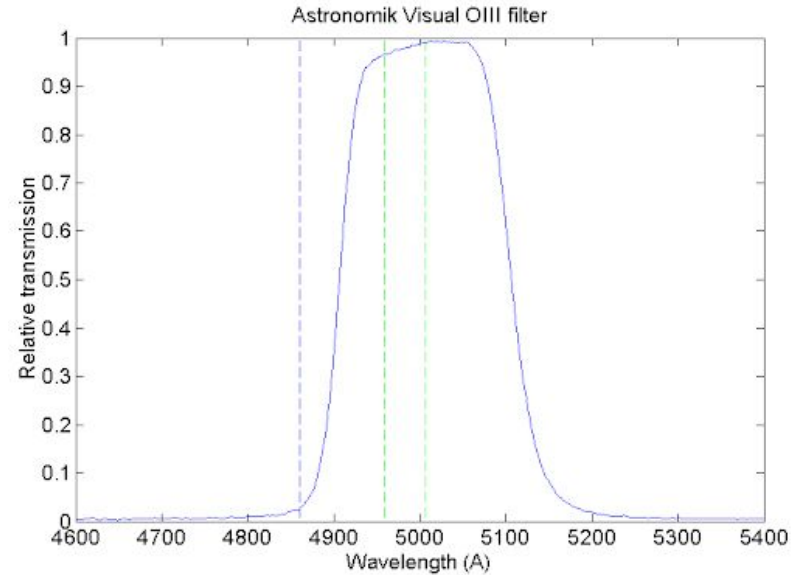
Transmission Curve

Filter Response

Evaluation of the filter Response only

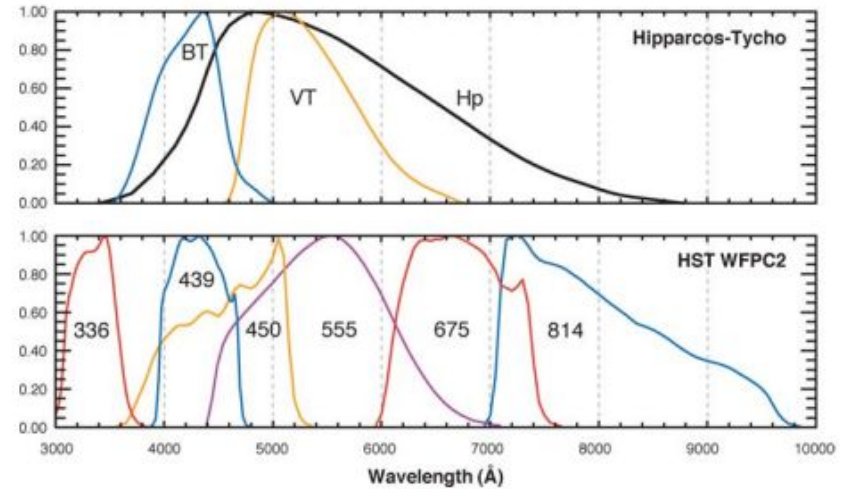
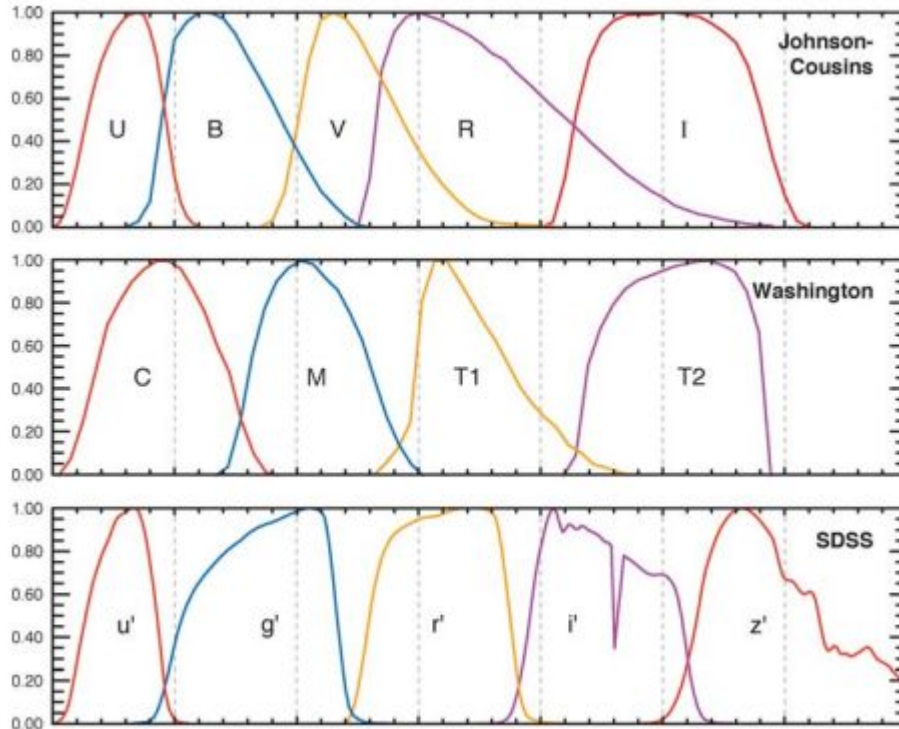
Instrumental Response

Instrumental response impacts the real final response of the system



$$\langle F_R \rangle = \int F_R(x) T(x) dx$$

Photometric Systems





Photometry based Reference Spectrum

Vega

Used (mainly) Vega star to create the reference spectrum

$$m_{AB,\nu} = 0$$

$$f_{\nu} = 3.63 \times 10^{-20} \text{ erg cm}^{-2} \text{ s}^{-1} \text{ Hz}^{-1}$$

AB ν and ST λ

Use constant flux density per unit frequency and constant flux density per unit wavelength

$$m_{ST,\lambda} = 0$$

$$f_{\lambda} = 3.63 \times 10^{-9} \text{ erg cm}^{-2} \text{ s}^{-1} \text{ \AA}^{-1}$$



Zero point type: Conversion to Flux

Pogson

Quite common in astronomy

$$f = f_0 10^{-(m-m_R)/2.5}$$

Asinh

Also called Iupitudes. They have a softening parameter

$$f = f_0 10^{-(m-m_R)/2.5} \left[1 - b^2 10^{2(m-m_R)/2.5} \right]$$

Linear

Quite common in high energy

$$f = f_0 \frac{m}{m_R}$$



Use cases

- Create SED with photometry from different catalogues
 - Convert to flux
 - TAP annotations - VO/DML!
- Synthetic photometry
 - For theoretical spectra
 - For future observations (e.g. for proposals)
- Photometry filters classification

Phot DM 1.1



Generation of XML VODML document / M.Louys

- Phot DM1.1 IVOA discussion page at <https://wiki.ivoa.net/twiki/bin/view/IVOA/PhotDMv1-1> with all products attached :
- *xmi*: New version of the model in Modelio 3.8
- translation to:
 - *xml*: new version of the VO-DML XML description of the data model.
 - *html*: documentation of the model in HTML



Changes PhotDM1.0 --> PhotDM 1.1

- Very few changes in order to preserve backward compatibility
- Simple types for attributes are now based on the VODML *ivoa:* template types.
- Deprecates the use of PhotDM1.0 **PhysicalQuantity** class, whose granularity does not fit the one adopted in VODML.
- Define components classes : Bandwidth , SpectralLocation used in **PhotometryFilter**
- include *referenceMagnitude* attributes directly in **ZeroPoint** class



Discussion points

- Keep Utypes tables in the specification (backward compatibility)
- Components classes of PhotometryFilter have been renamed
→ roles and leaves classes are now in sync
- PhotDM 1.0 uses more advanced types than what is defined currently in the *ivoa:* template
 - **UCD** → derived from *ivoa:string*
 - **ISOTime** → derived from *ivoa:datetime*
currently these are derived in the PHOTDMv1.1 schema
 - Question for future use:
shall we insert them in the *ivoa:* template model instead



Calendar ?

Please review the document currently at <https://github.com/ivoa-std/PhotDM>

a message to the list will be issued on dm list

proceed for PR and RFC before January 30, 2022