



ASTRON

Netherlands Institute for Radio Astronomy

Apertif DR1

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and the Apertif team

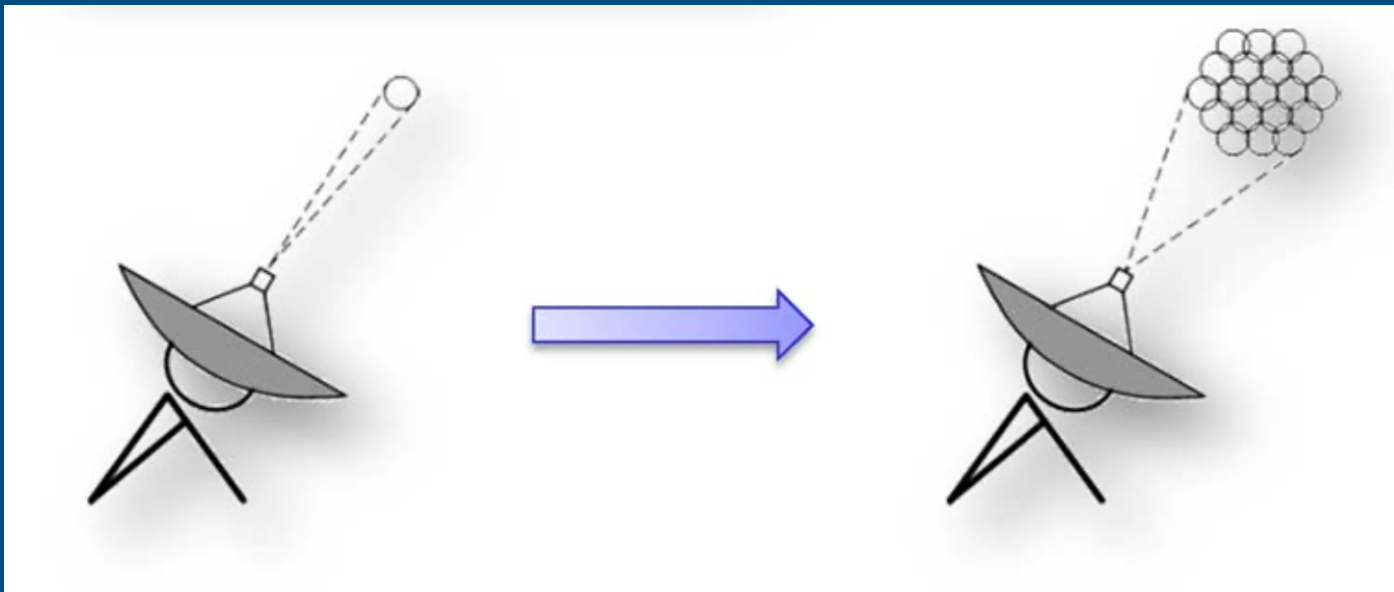
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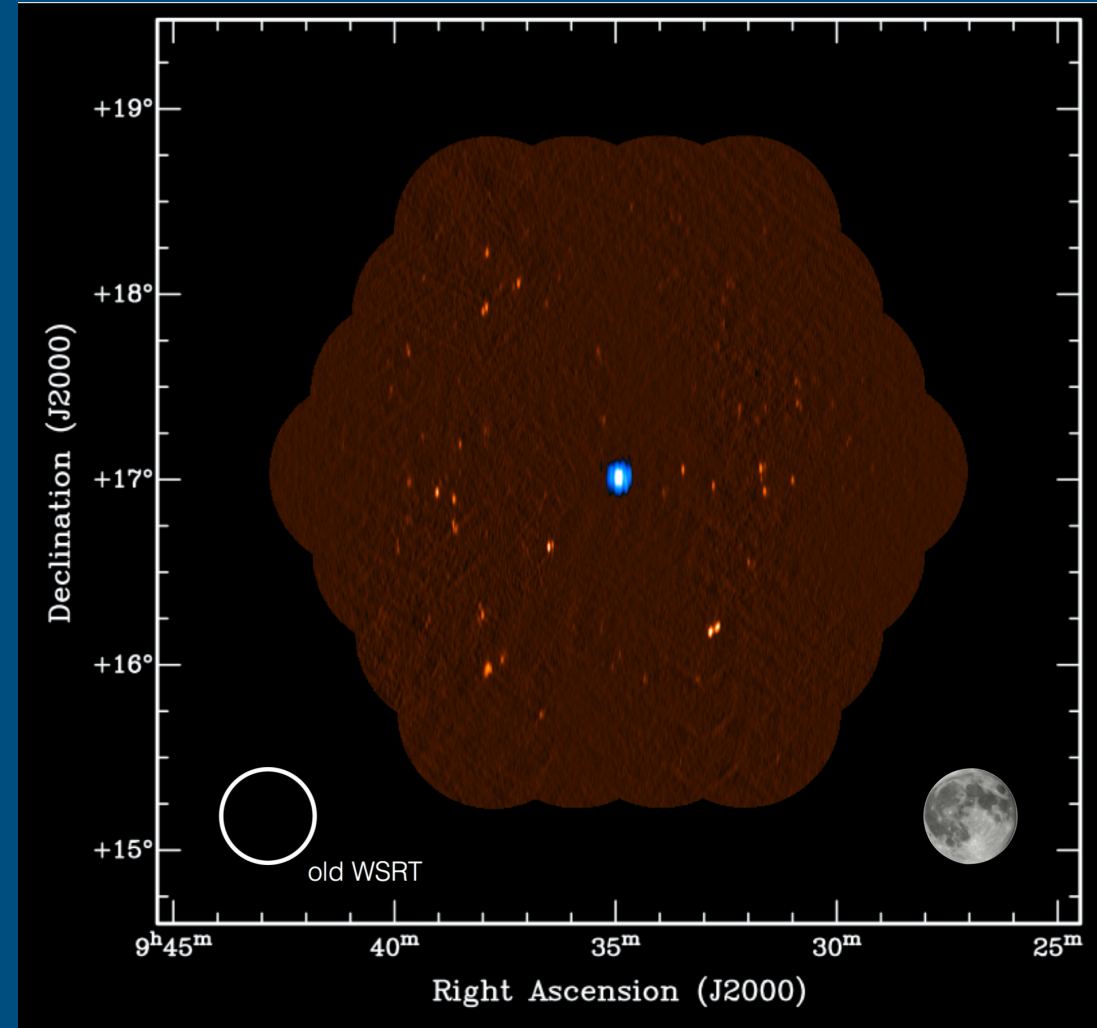
Apertif, Phased Array Feeds for the WSRT

- Transform the WSRT into an efficient 21-cm survey facility
 - 17x Survey speed increase
 - SKA Pathfinder



Apertif

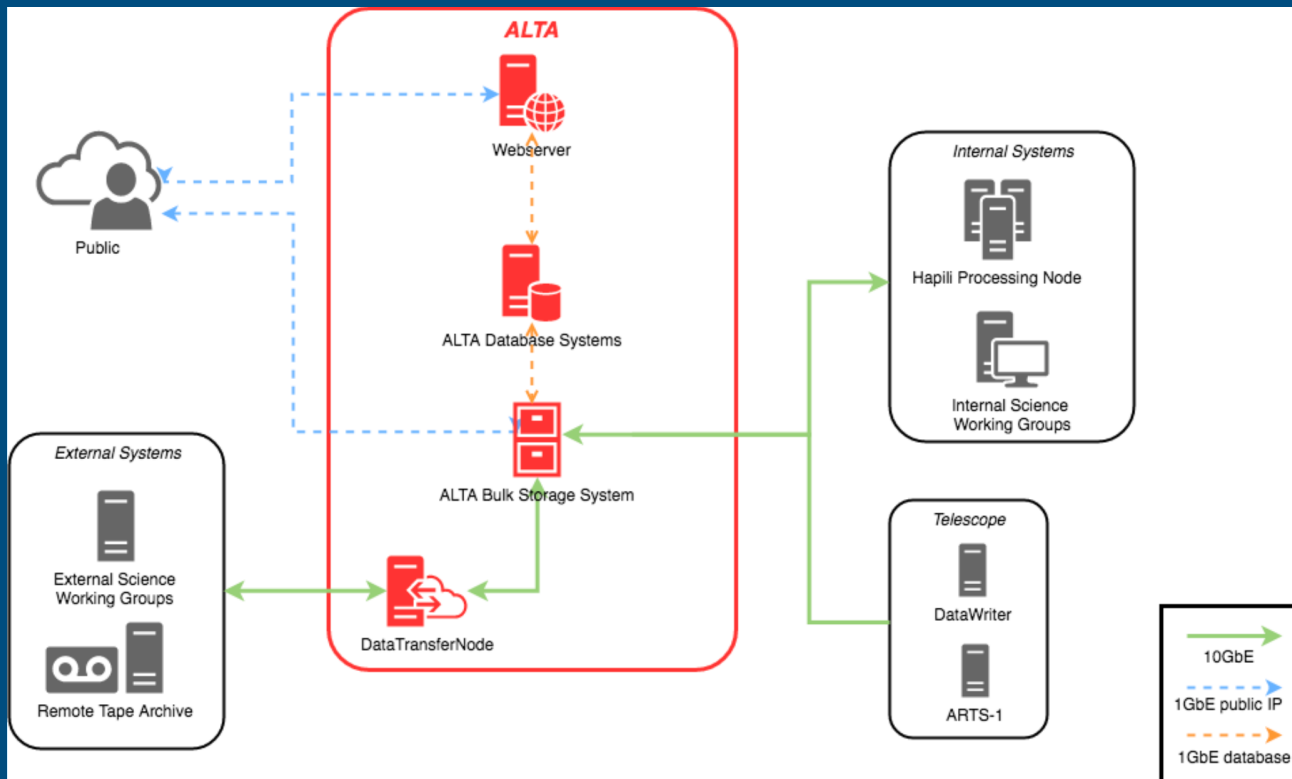
- Original 4 year legacy survey program
 - ~350 sq degree medium deep survey (10x12 hr)
 - ~3500 sq deg wide survey (1x12 hr)
- Radio continuum, polarization and neutral hydrogen
- Role of environment & interaction on galaxy processing
- Finding the smallest galaxies
- Connecting cold gas to AGN
- Understanding the faint radio population
- Studying magnetic fields in galaxies



First Apertif image
Tom Oosterloo, copyright ASTRON
<http://www.astron.nl/dailyimage/main.php?date=20170131>

ALTA – the Apertif Long-Term Archive

- ALTA is the main source of Apertif data.
 - VO interface is the main interface for external user access



- ASTRON (Dwingeloo): Online storage (including 'nearline' object store)



- SURFsara (Amsterdam): Offline storage (including online cache & staging)






































- Using IRODS for internal data management

<https://alta.astron.nl>

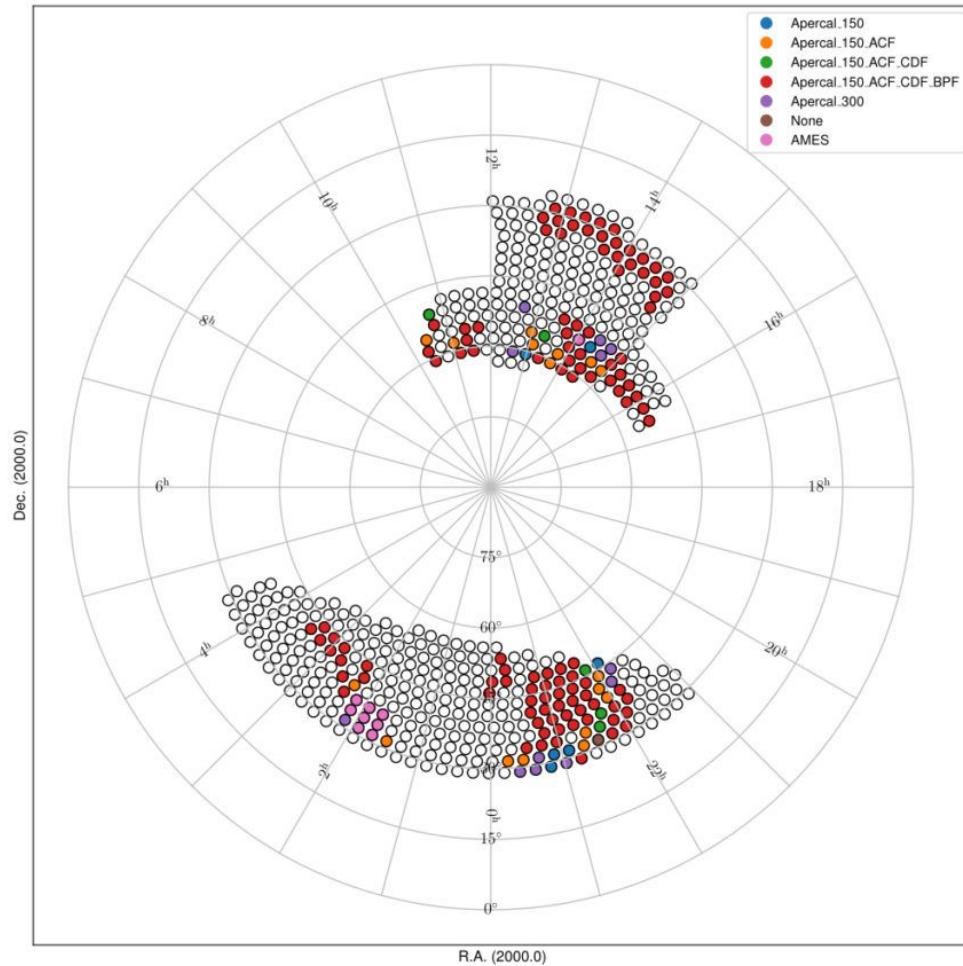


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The DR1 data set – the released products

Type	Discovery	Access
Raw observational data	 	  
Full time, Stokes and spectral resolution calibrated visibility data	 	  
Calibration tables	 	  
Inspection plots		
Continuum images	 	   
Polarization images and cubes	 	   
Line and dirty beam cubes	 	   

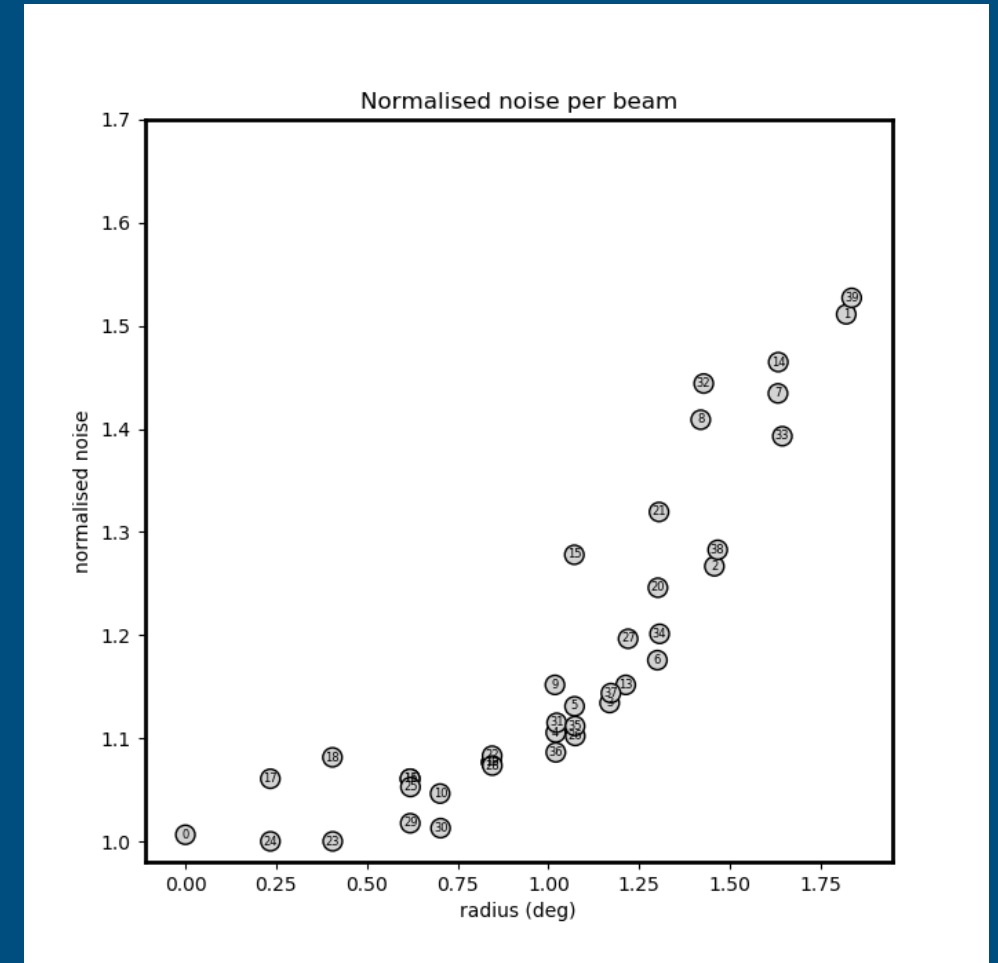
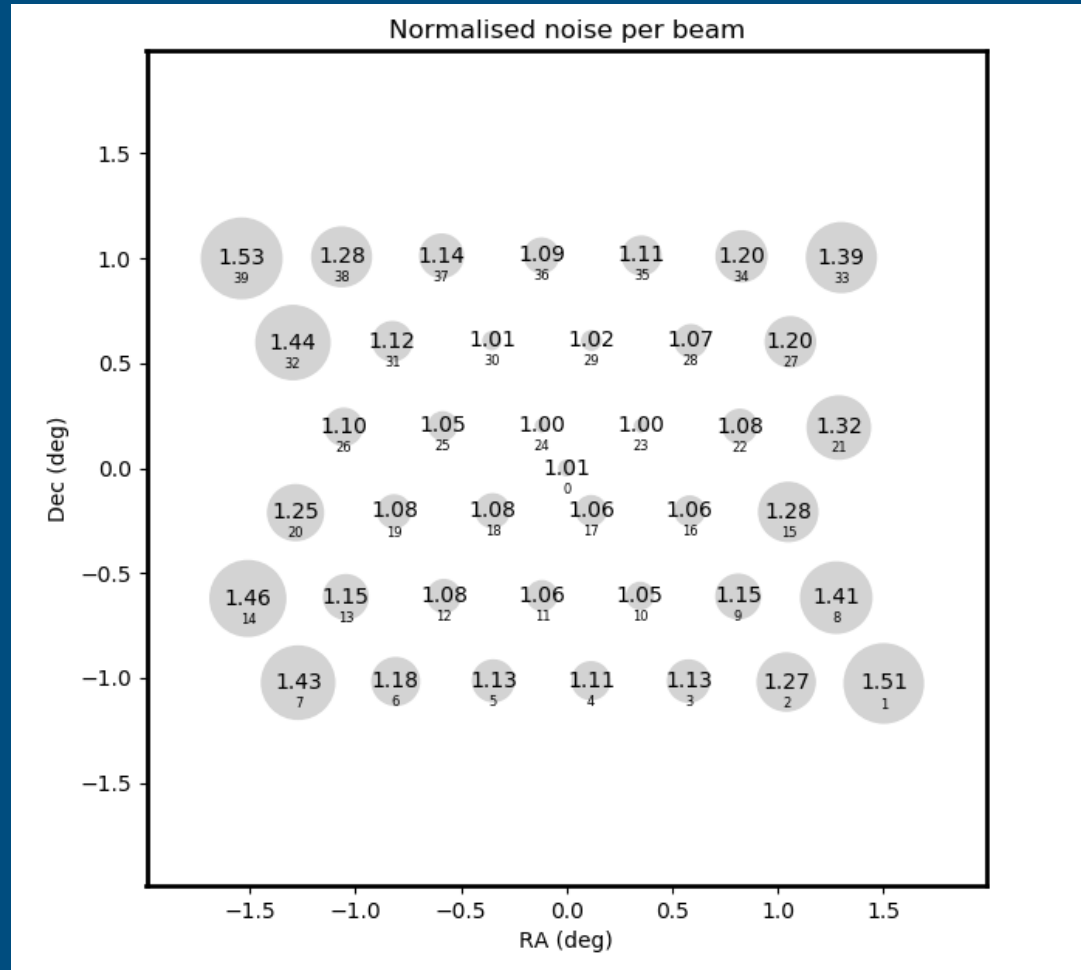
The DR1 data set – wrt the full survey



The DR1 data set – product quality

- Validation of continuum data products by assessing they meet the resolution requirements, minimum sensitivity requirements, and have no significant image artifacts.
 - Other data types released based on validation of corresponding continuum product!
- Most polarisation cubes also pass their own validation, since the criteria are comparable.
- Line cubes get flagged “G”ood, “O”kay or “B”ad depending on severity of artifacts.

The DR1 data set – beam noise



And now... The demo!

Closing remarks

- Currently we only have an image HiPS. Cube HiPS may be a good addition
- Looking forward to other surveys with externally-hosted data (e.g. LOTSS DR2).
- Any feedback is very welcome :)

<http://hdl.handle.net/21.12136/B014022C-978B-40F6-96C6-1A3B1F4A3DB0>