CSP pannel follow up ObsCore extension For radio data

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# Follow-up of pannel discussion

- Are you aware of the ObsCore extension for radio data ? Is that useful for discovery of the data you expose ?
- Access and processing of huge datasets, in which direction to go ?
  - Code to the data (platforms, jupyter notebooks, etc..)
  - SODA to extract/reshape regions of interest ?
  - HiPS cube for multi resolution access?
  - Any kind of combination ?

# Short summary of what the extension is all about

- ObsCore allows to discover datasets by constraining datasets standardised metadata
  - Instrumental Provenance (facility, instrument)
  - Identification
  - Product type (image, cube, spectrum, etc...)
  - Curation
  - Charactérisation of physical axes (spatial, spectral, time, polarisation)
  - Data access mode (url, format, datalink, cutout, etc..)
- Is that sufficient for all kind of datasets ?
  - Not always

# Spatial axis addition Uv coverage characterisation

- s\_fov\_min and s\_fov\_max (each end of the spectral window)
- s\_resolution\_min and s\_resolution\_max (each end of the spectral window)
- s\_maximum\_angular\_scale (because large scales are filtered in interferometry)
- uv\_distance\_min , uv\_distance\_max (for scale filtering and resolution)
- uv\_distribution\_exc (distribution excentricity data regularity)
- uv\_distribution\_fill : (distribution filling factor -data sampling)

## Spectral axis additions Product types additions

- Addition of f\_resolution (as a counterpart to em\_res\_power)
- f\_min and f\_max beside em\_min and em\_max
- Addition of « spatial\_profile » dataproduct\_type
- velocity/position profiles

# Sky scan modes as additional parameters tracking modes



Figure 1: Single Dish Observation Sky scan modes

## **Proposed Instrumental parameters**

- Antenna typical diameter (for all)
- •And for interferometry
  - Number of antennae
  - Minimal distance between antennae
  - Maximal distance between antennae

#### CDS prototype demo (implemented in Dachs) :

frequency between 1 and 2 Ghz (upper left – 3 results) freq. between 1 and 2 Ghz and spectral resolution better than 100 Mhz (lower right – 2 results)



#### CDS prototype demo (implemented in Dachs) :

s\_resolution better than 0.6 arcsec (upper left - no result) s\_resolution \_min better than 0.6 arcsec(lower right - 1 result)



Aladi

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resolution better than 0.6 arcsec

SELECT obs publisher did, target name, s ra, s dec, s fov, s resolution, em min, em max, s fov min, s fov max, s resolution min, s resolution max, f resolution, f min, f max, instrument ant max dist, instrument ant diameter FROM rucio.obscore where s resolution < 0.6

> Server sele SFile W... Wols...

> > MIRAD TAP Gaia ACK SKYB

yperLED Aladir SIAv2

No

CIRCLE - Add

SYNC - Async jobs>>

• vo-proto-debian.cds.unistra.fr

Check.

#### CDS prototype demo (implemented in Dachs) : s\_fov larger than 7.5 deg (upper left - two results) s\_fov\_min\_larger than 7.5 deg(lower right - 1 result)



 $\rightarrow$  all channels of the dataset have a field of view larger than 7.5

## CDS prototype demo (implemented in Dachs) : 16 datasets of ObsCore discovery service of S(ka)RCnetwork



## **ObsCore** for complex data

- Observation made of several datasets
  - Distinction obs\_id/obs\_publisher\_did
  - Raw per obs\_publisher\_did, not per observation
- Characterisation at the dataset level
- Example : raw interferometry data with
  - Main target and calibrator
  - Two different spectral windows
  - → 4 datatasets with same obs\_id
- Reverse situation : dataset produced from several observations : combined obs\_id

- f\_min and f\_max beside em\_min and em\_max
  - Rationale :
    - Something natural for the users directly available for the users
    - Something natural for the users for the queries.
    - Parameters produced from basic ones in a view
  - Cons:
    - Don't duplicate information in the standard
    - And either :
      - Use « user defined function » in query and display
        - ivo\_specconv(f, « funit », « wlunit »)
        - 1 = ivo\_interval\_overlaps(em\_min, em\_max, ivo\_specconv(1.5, "GHz", "m"), ivo\_specconv(1, "GHz", "m"))
      - Let the clients do the transformation in both directions

- Instrumental details :
  - proposed because they give an hint on some data characterization (sensibility, resolution, data quality)
  - Cons :
    - They are very specific to each experiment and do not previde generic information
    - Except (maybe) if we have use cases for that
- Science cases were missing anyway
  - Partially done (but not for instrumental details)

- How to expose it in a TAP service
- Two possible ways in TAP:
  - 1 basic ObsCore table, + 1 table with ObsCore+ extension. New StandardID for the latter :

ivo://ivoa.net/std/ObsCore#table-1.0

ivo://ivoa.net/std/ObsCore#radioExt-1.0

- 1 basic ObsCore table + 1 table with extension only. (ivo://ivoa.net/std/obsradio#table-1.%') User or client have to join the tables
- Solution 1 become complex if we have several extensions which may be combined or not
- Solution 2 masks that the extension is meaninless wthout the baisc table. Doesn't tell us where the baisc table lies.

- Compromise :
  - Set the standarID on a schema containing the Obscore table and the extension(s) with standarID ivo://ivoa.net/std/ObsCore
  - Have specific standardID on the tables (ObsCore basic, extensions)
  - Possibility to build view providing the joins in the same schema (but without standardID)