Enhancing your publications with DOIs from MAST

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for

MAST: Karen Levay, Amanda Marrione

STScI Library: Jill Lagerstrom

AAS Journals: Gus Muench, Julie Steffen, Ethan Vishniac

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To build on your work, readers may need your data. Getting data from tables is slow and difficult.

Table 1. Sample of PNe with *HST* WFPC2 or WFC3 Hα and [O III] observations

PN G	Common Name	Hα Exp. Time (s)	[O III] Exp. Time (s)	Proposal ID	PN G	Common Name	Hα Exp. Time (s)	[O III] Exp. Time (s)	Proposal ID
000.3+12.2	IC 4634	1000	1000	6856	084.2+01.0	K 4-55	2460	2440	11956
001.2+02.1	Hen 2-262	280	280	9356	084.9−03.4	NGC 7027[c]	500	100	11122
001.7−04.4	H 1-55	200	280	9356	089.8−05.1	IC 5117	240	320	8307
002.3−03.4	H 2-37	280	280	9356	096.4+29.9	NGC 6543	800	1600	5403
002.4+05.8	NGC 6369	640	640	9582	106.5−17.6	NGC 7662	200	500	6117, 6943, 8390
002.7−04.8	M 1-42	900	1800	11185	111.8−02.8	Hb 12	1600	1600	11093
002.9−03.9	H 2-39	280	280	9356	138.8+02.8	IC 289	2000	2000	11956
003.5−04.6	NGC 6565	160	320	11122	144.1+06.1	NGC 1501	1600	2000	11956
003.6+03.1	M 2-14	280	280	9356	189.1+19.8	NGC 2371-72	1600	1600	11093
003.8+05.3	H 2-15	280	280	9356	197.8+17.3	NGC 2392	400	400	8499
003.9−03.1	KFL 7	280	280	9356	215.2−24.2	IC 418	888	360	6353, 7501
004.0−03.0	M 2-29	200	160	9356	231.8+04.1	NGC 2438	2080	2080	11827
004.1−03.8	KFL 11	280	280	9356	215.6+03.6	NGC 2346	200	120	7129
004.8−22.7	Hen 2-436	200	160	9356	234.8+02.4	NGC 2440	1600	1600	11090
004.8+02.0	H 2-25	400	400	9356	249.0+06.9	SaSt 1-1	200	280	8332
005.2−18.6	StWr 2-21	280	280	9356	261.0+32.0	NGC 3242	100	1200	6117, 7501, 8773
006.1+08.3	M 1-20	200	160	9356	261.9+08.5	NGC 2818	1600	2000	11956
006.3+04.4	H 2-18	280	280	9356	272.1+12.3	NGC 3132	400	1200	6221, 8390
006.4+02.0	M 1-31	780	160	9356	285.6−02.7	Hen 2-47	1600	1600	11090
006.8−19.8	Wray 16-423	200	160	9356	285.7−14.9	IC 2448	200	320	11122
006.8+04.1	M 3-15	200	160	9356	294.6+04.7	NGC 3918	140	320	11122
007.5+04.3	Th 4-1	280	280	9356	305.1+01.4	Hen 2-90	2325	1210	8345, 9102
008.2+06.8	Hen 2-260	200	460	9356	307.5−04.9	MyCn 18	600	1400	6221
008.6−02.6	MaC 1-11	280	280	9356	309.1−04.3	NGC 5315	1600	1600	11090
009.3+05.7	Hen 3-1475	830	800	7285	312.3+10.5	NGC 5307	1600	1600	11090
010.0+00.7	NGC 6537	1240	1000	6502	319.6+15.7	IC 4406[d]	540	600	8726, 9314
010.8+18.0	M 2-9	1240	1000	6502	324.0+03.5	PM 1-89	4900	2900	5404, 5864
010.8−01.8	NGC 6578	160	320	11122	327.8+10.8	NGC 5882	140	380	11122
019.4−05.3	M 1-61	240	320	8307	331.1−05.7	PC 11	200	280	8332
025.3+40.8	IC 4593	1600	1600	11093	331.3−12.1	Hen 3-1357	240	368	6039, 8390
025.8−17.9	NGC 6818	520	1300	6792, 7501, 8773	331.7−01.0	Mz 3[e]	1260	1160	6856, 9050
027.6+04.2	M 2-43	520	1800	8307	341.8+05.4	NGC 6153	1000	1200	8594
034.6+11.8	NGC 6572	180	840	7501, 9839	349.5+01.0	NGC 6302[a]	2100	2220	11504
036.1−57.1	NGC 7293	1800	1800	5977	351.1+04.8	M 1-19	160	160	9356
037.7−34.5	NGC 7009	400	320	8114	351.9−01.9	Wray 16-286	200	280	9356
037.8−06.3	NGC 6790	160	200	8307	352.6+03.0	H 1-8	200	280	9356
043.1+37.7	NGC 6210	320	320	6792	353.5−05.0	JaFu 2[f]	3600	2000	6780
054.1−12.1	NGC 6891	1280	320	11122	354.5+03.3	Th 3-4	280	280	9356
054.2−03.4	Necklace Nebula[a]	2000	2000	12675	354.9+03.5	Th 3-6	280	400	9356
057.9−01.5	Hen 2-447	520	1800	8307	355.4−02.4	M 3-14	200	160	9356
060.1−07.7	NGC 6886	1120	1020	7501, 8345, 8773	355.9+03.6	H 1-9	280	280	9356
060.8−03.6	NGC 6853	2000	1000	8726	356.1−03.3	H 2-26	280	280	9356
063.1+13.9	NGC 6720	480	720	7632, 8726	356.5−03.6	H 2-27	360	400	9356
064.1+04.3	M 1-92	680	2080	6533	356.9+04.4	M 3-38	280	280	9356
064.7+05.0	BD+30°3639	484	900	8116, 8390	357.1−04.7	H 1-43	200	280	9356
065.0−27.3	Ps 1[b]	11420	1040	6751	357.2+02.0	H 2-13	280	280	9356
071.6−02.3	M 3-35	520	1000	8307	358.5−04.2	H 1-46	160	160	9356
073.0−02.4	K 3-76	6	18	6943	358.5+02.9	Wray 16-282	280	280	9356
074.5+02.1	NGC 6881	280	320	8307	358.9+03.4	H 1-19	200	280	9356
082.1+07.0	NGC 6884	1100	560	8345, 8390	359.2+04.7	Th 3-14	280	400	9356
082.5+11.3	NGC 6833	40	3	6943, 6353	359.3−00.9	Hb 5	1300	1000	6502
083.5+12.7	NGC 6826	100	100	6117					

Guerrero+ 2013

*We searched MAST for HST WFPC2 or WFC3 coeval Hα and [O iii] images of PNe available by March 2013. This search yielded Hα and [O iii] images for **103** PNe obtained through the F656N and F502N filters, respectively*

*We present a catalogue of photometric and structural properties of **228** nuclear star clusters (NSCs) in nearby late-type disc galaxies. These new measurements are derived from a homogeneous analysis of all suitable Wide Field Planetary Camera 2 (WFPC2) images in the Hubble Space Telescope (HST) archive.*

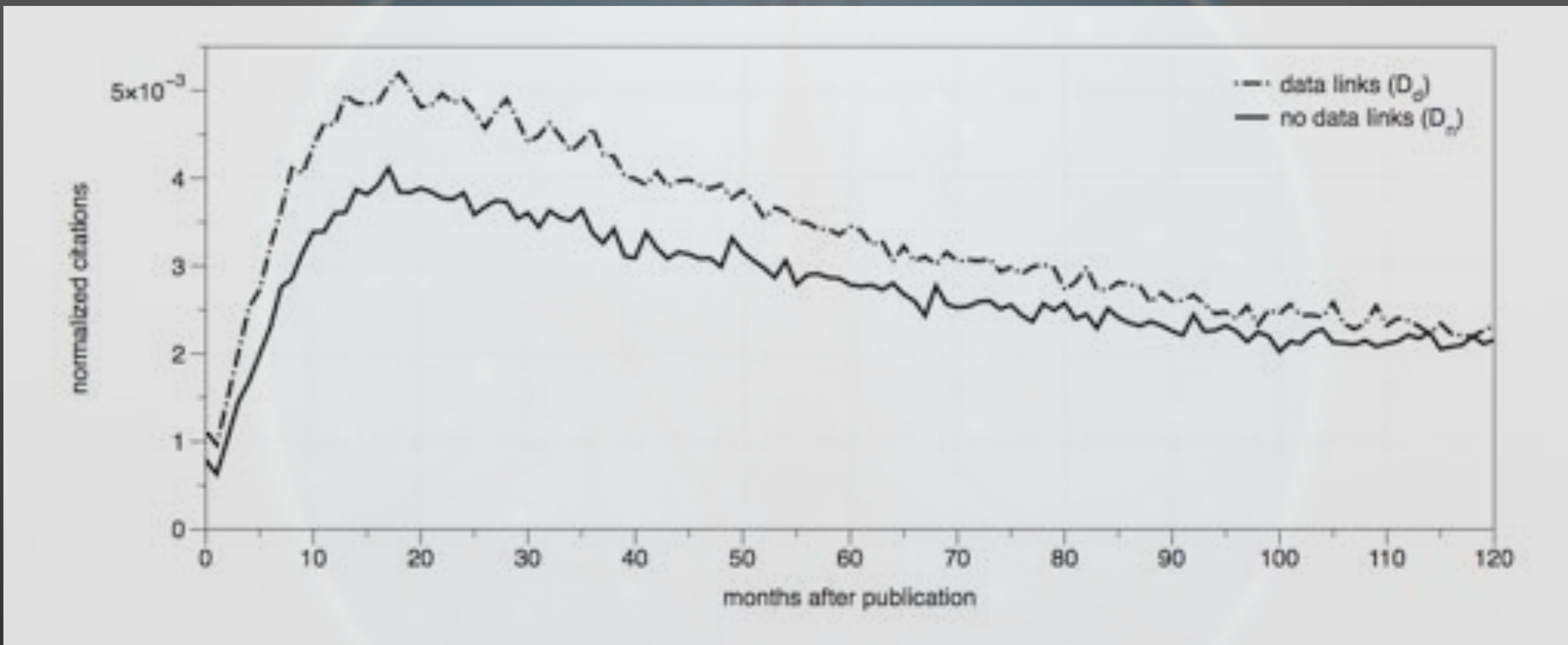
Table 1. Main properties of the galaxy sample with measured NSC properties. (All 228 galaxies are listed in the online version of the table).

Galaxy (1)	RA (hh:mm:ss) (2)	Dec. (dd:mm:ss) (3)	$m-M$ (mag) (4)	$E(B-V)$ (mag) (5)	B (mag) (6)	$B-V$ (mag) (7)	I (mag) (8)	R_{25} (kpc) (9)	ϵ (10)	PA (deg) (11)	Incl. (deg) (12)	Type (13)	t (14)
DDO078	10:26:27.78	67:39:25.1	27.82	0.018	15.8	–	–	1.063	0.00	–	0.0	I	10.0
IC 4710	18:28:37.95	−66:58:56.1	29.75	0.079	12.51	0.57	11.19	4.494	0.15	–	34.9	Sm	8.9
NGC 1258	3:14:05.50	−21:46:27.3	32.28	0.022	13.88	–	12.35	5.870	0.26	20.5	43.7	SABc	5.7
NGC 3319	10:39:09.47	41:41:12.5	30.7	0.013	11.77	0.41	11.46	7.289	0.51	36.	62.7	SBc	5.9
NGC 5334	13:52:54.44	−1:06:52.4	32.78	0.041	12.97	–	12.19	17.729	0.28	18.2	44.8	Sc	5.2
...	...	...	...	...	...	...	...	...	...	...	...	...	...

Notes: The values for all columns are taken from HyperLeda, except for columns 4 and 5, which are taken from NED. More specifically, the distance modulus $m-M$ in column 4 is the median value in NED. If the latter is not available, we adopt the redshift-derived distance modulus, modz, from HyperLeda.

Georgiev & Böker 2013

Papers with data links are more highly cited.



Henneken & Accomazzi 2012

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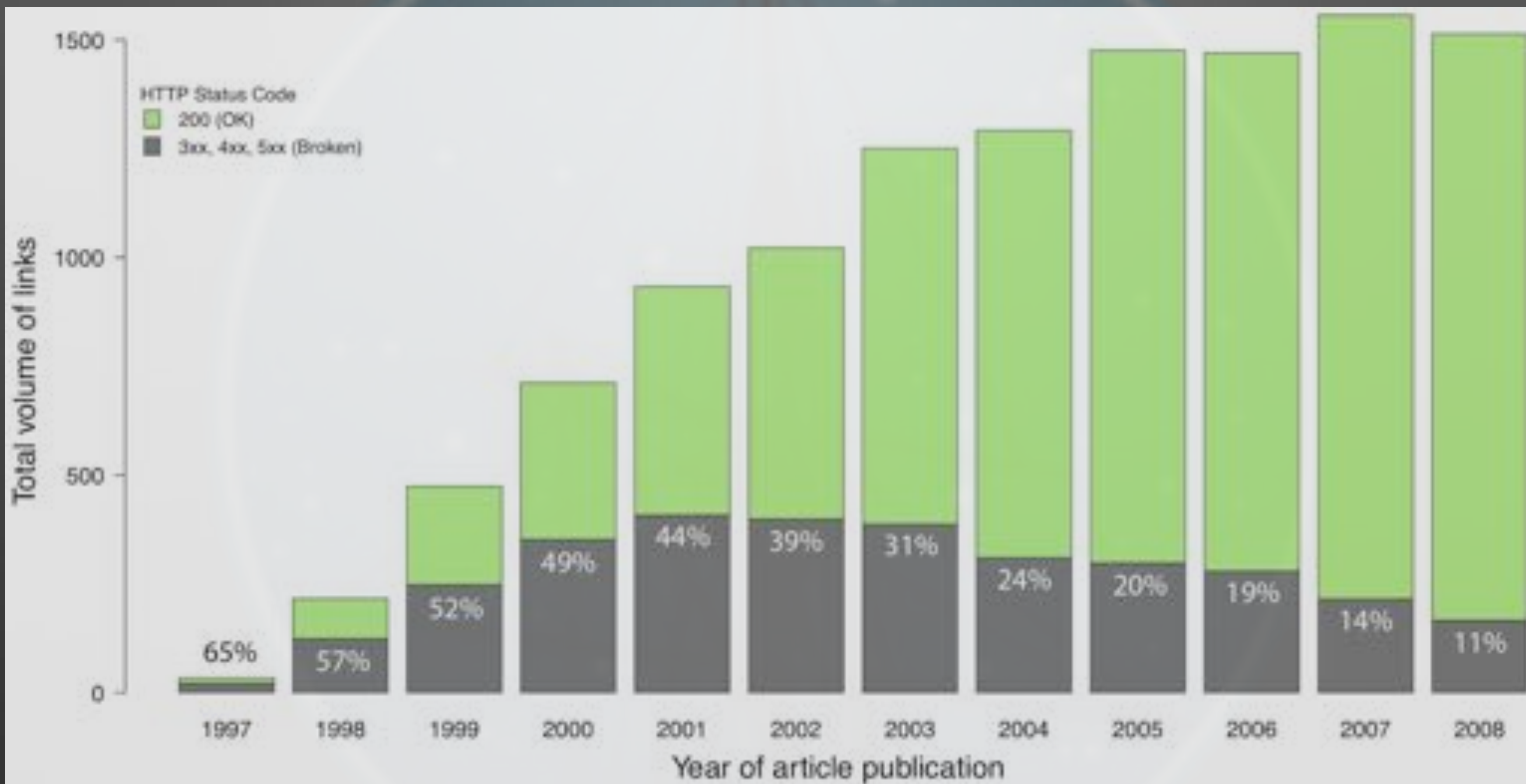

Pepe+ 2014

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(DOI: Digital Object Identifier
i.e. *"permalink")*

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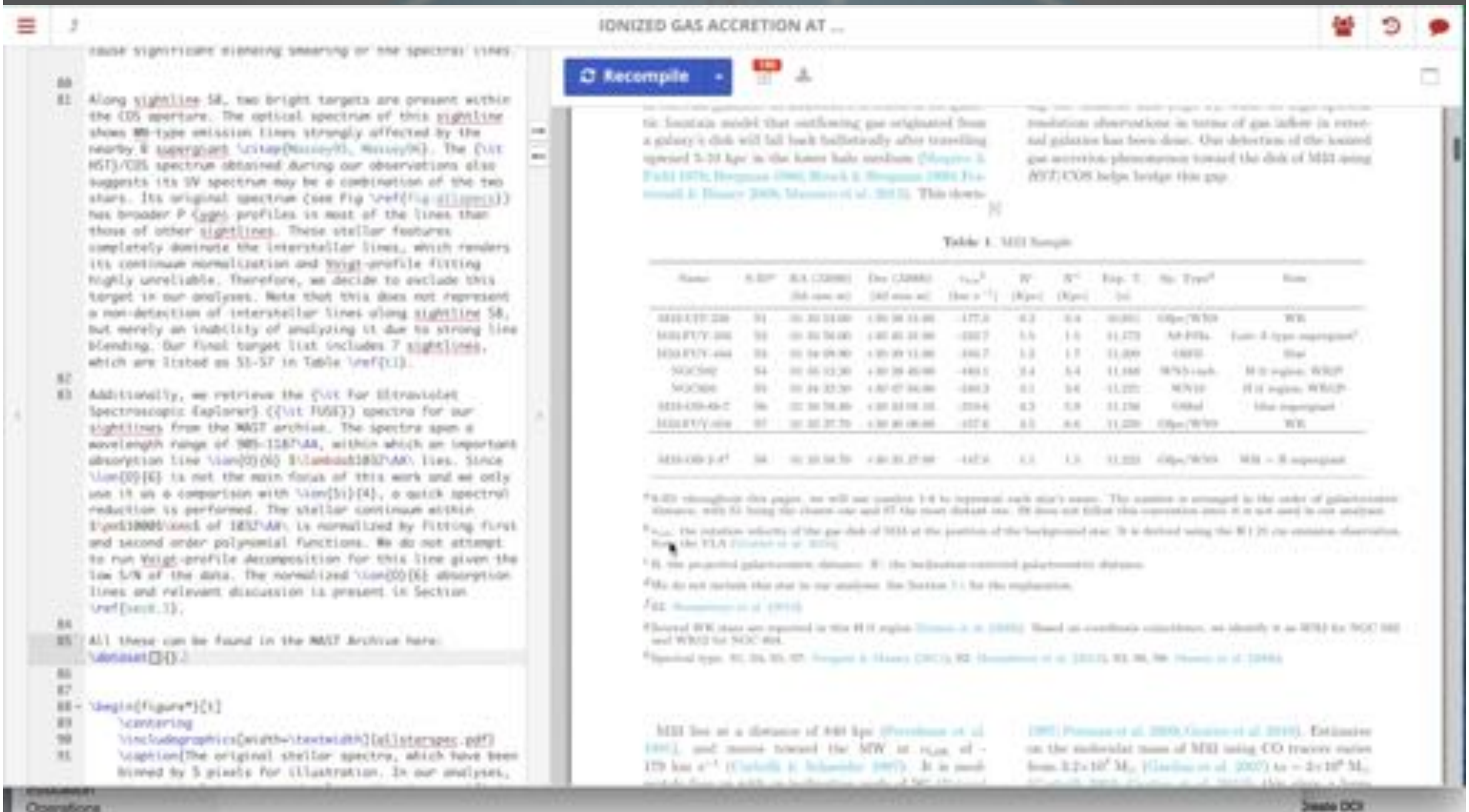

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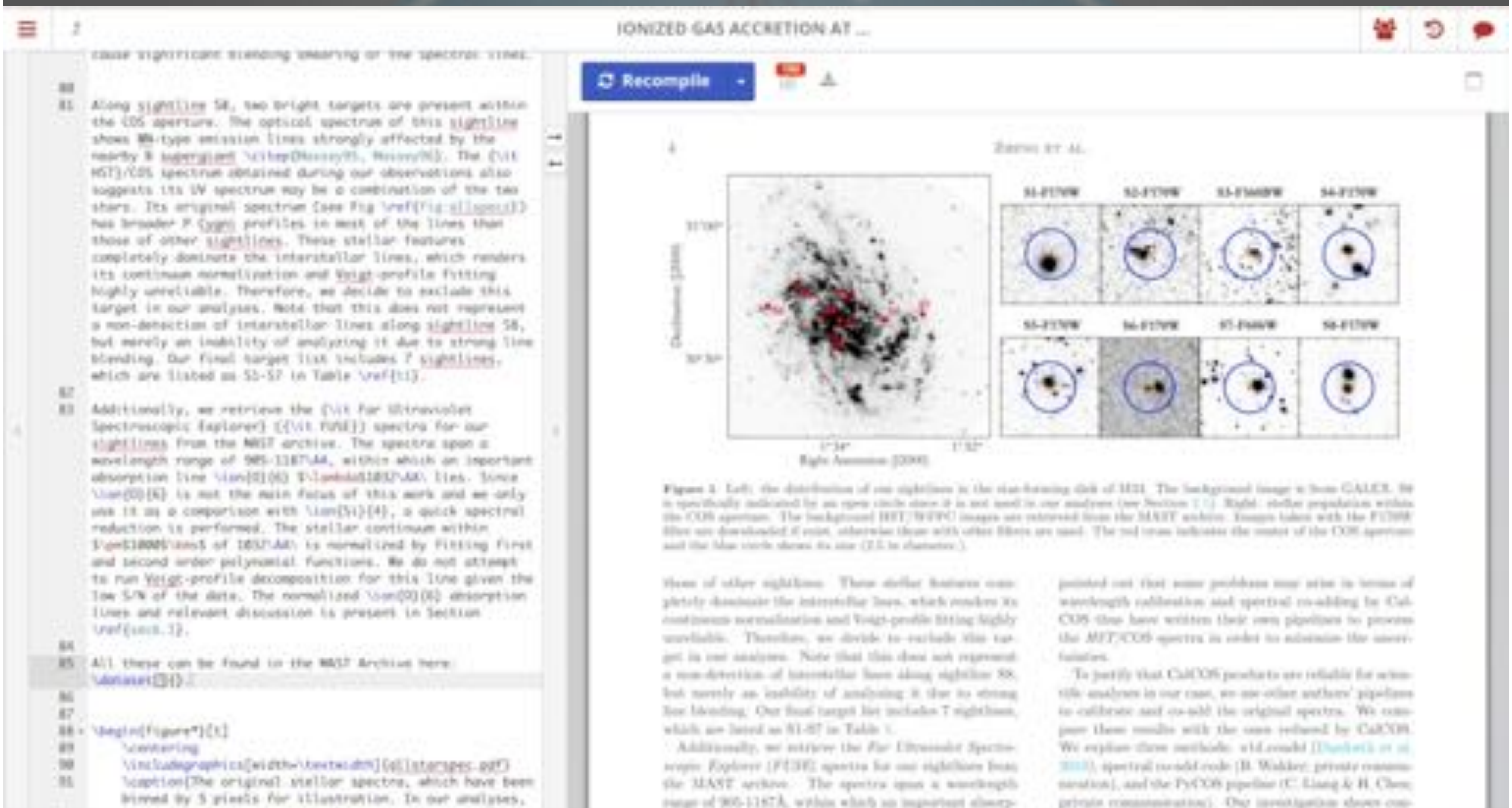

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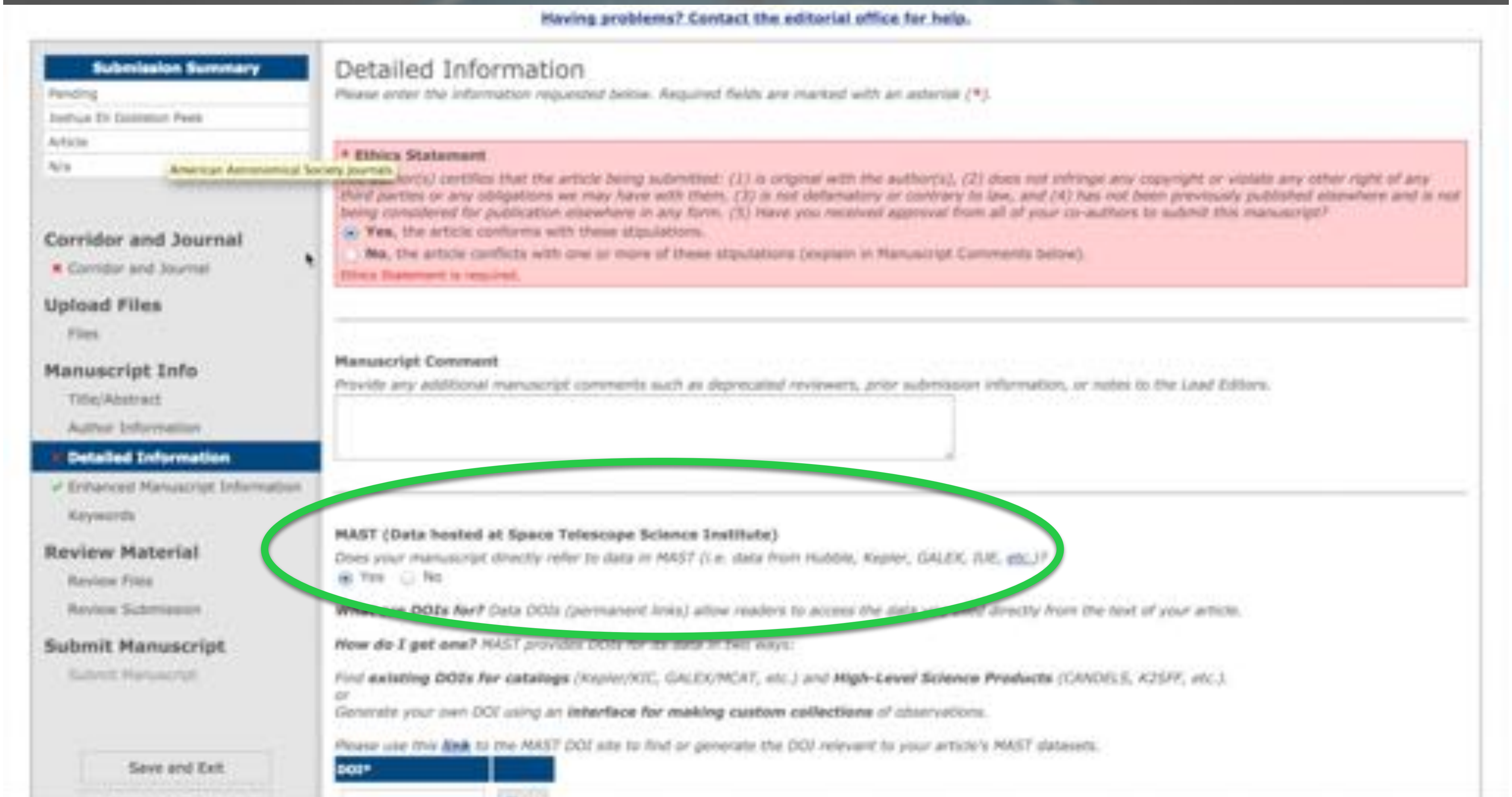

MAST DOI Page

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- GALEX/gPhoton
- Hubble Source Catalog
- Kepler/KIC
- K2/EPIC
- Guide Star Catalog 2
- Digitized Sky Survey

▶ An entire quarter or campaign from Kepler/K2?

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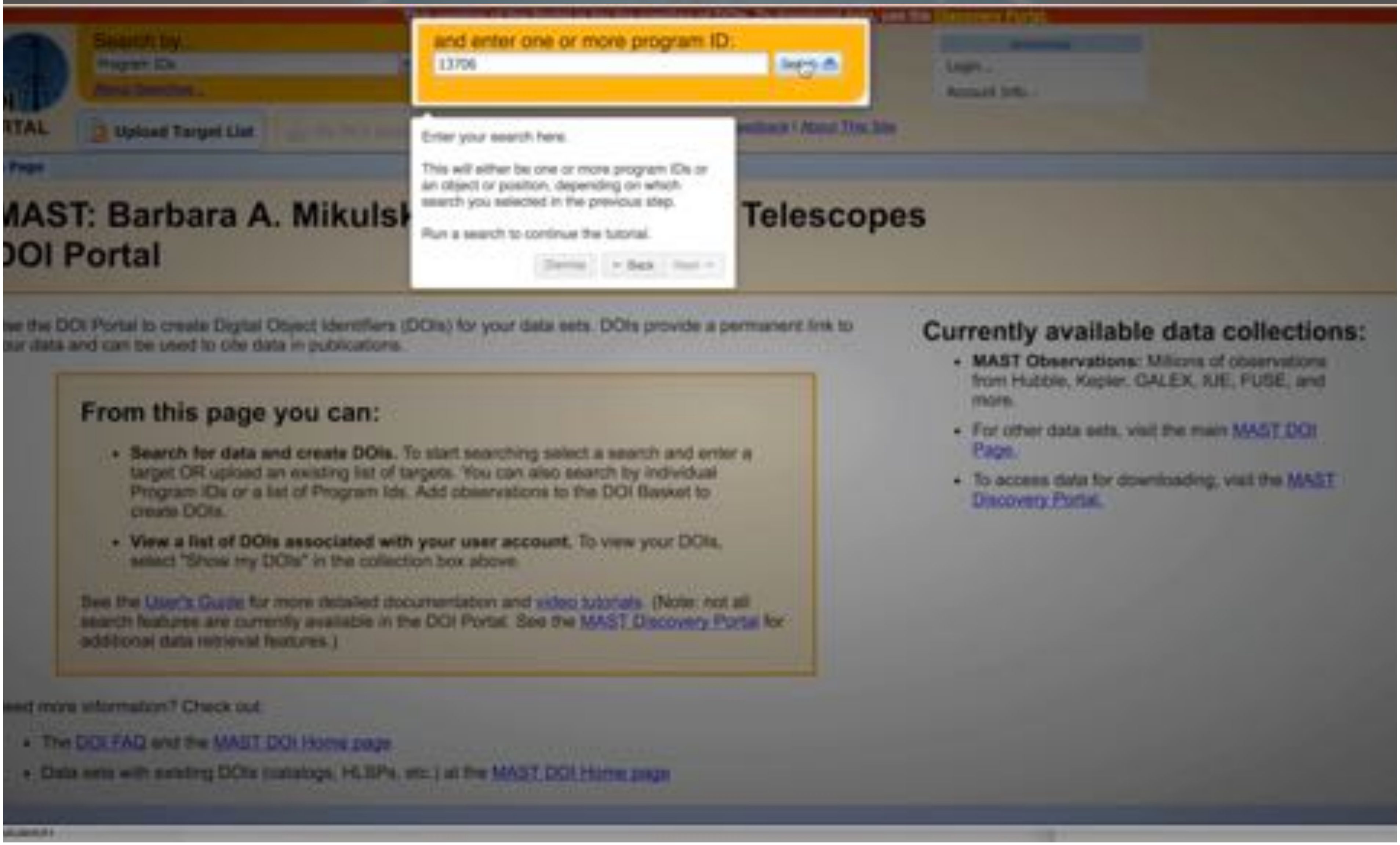

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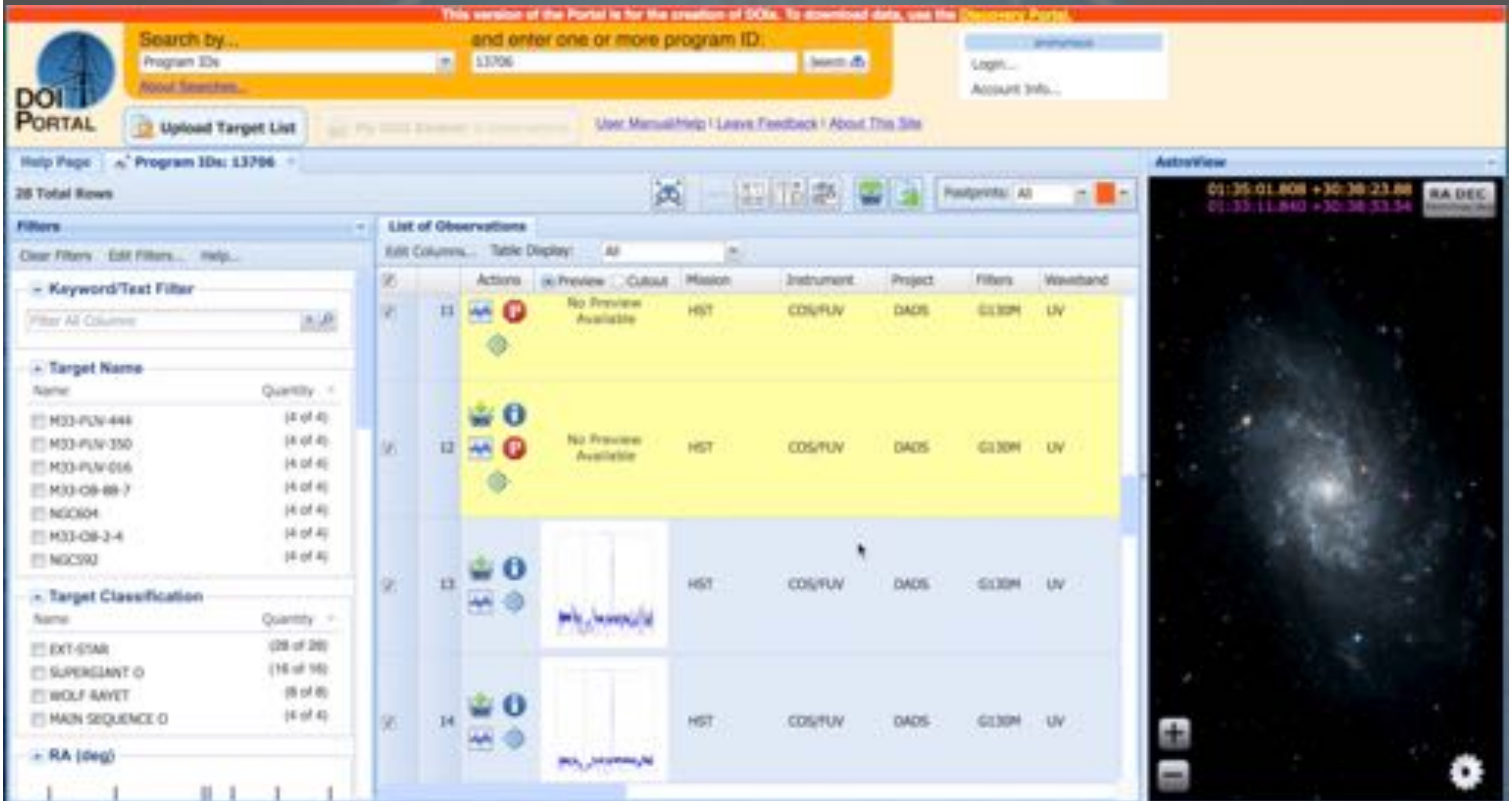

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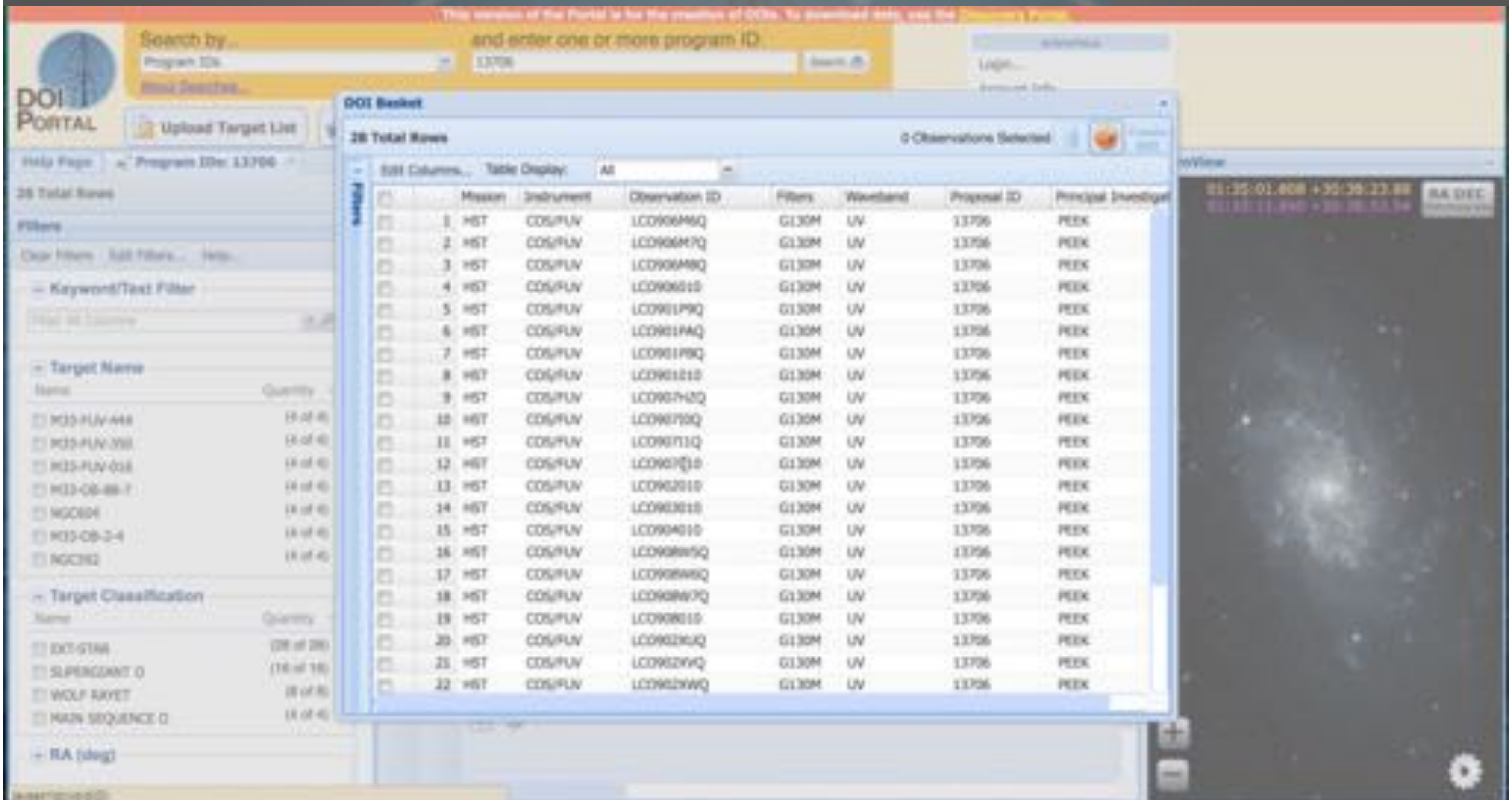

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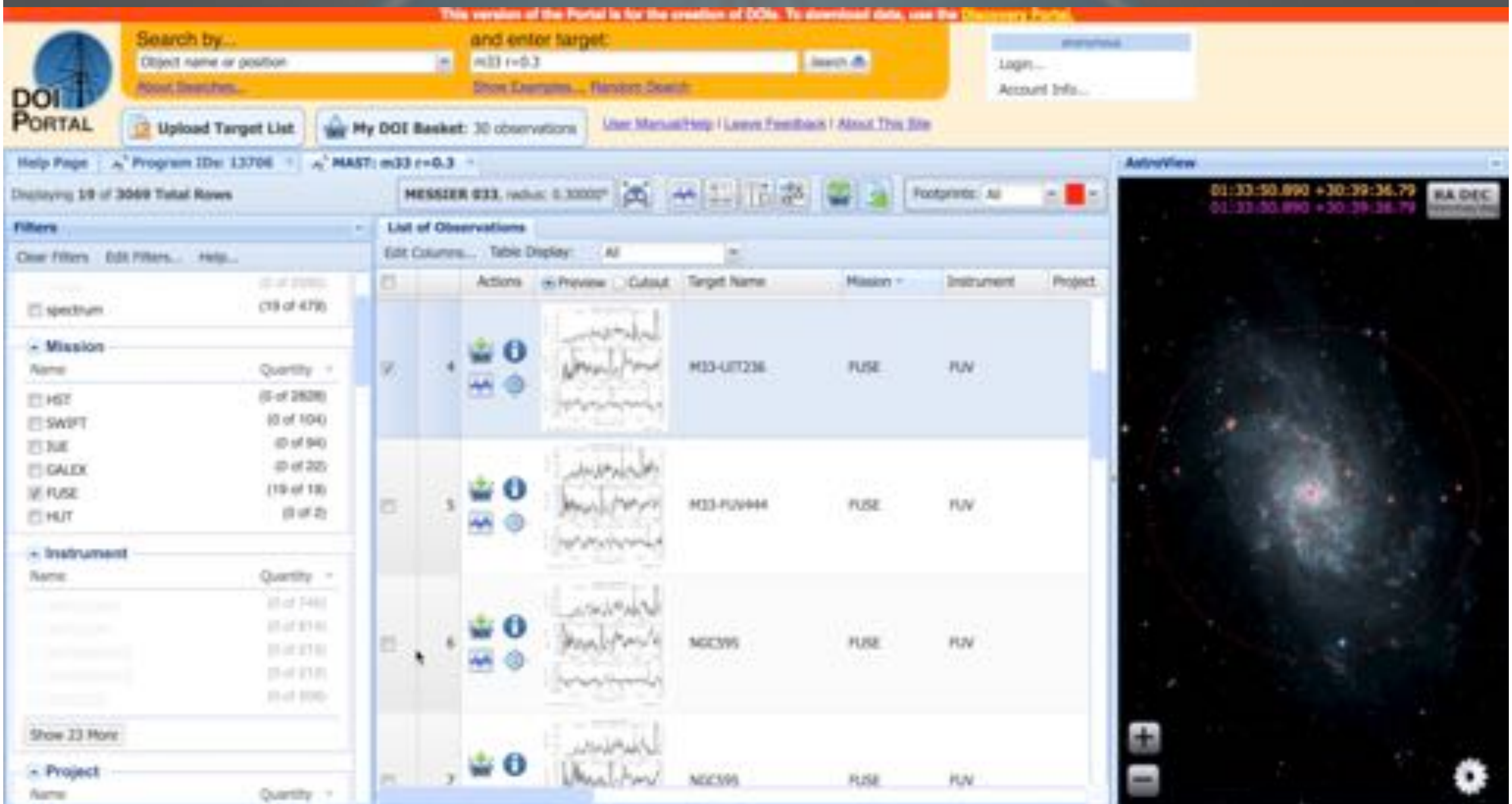

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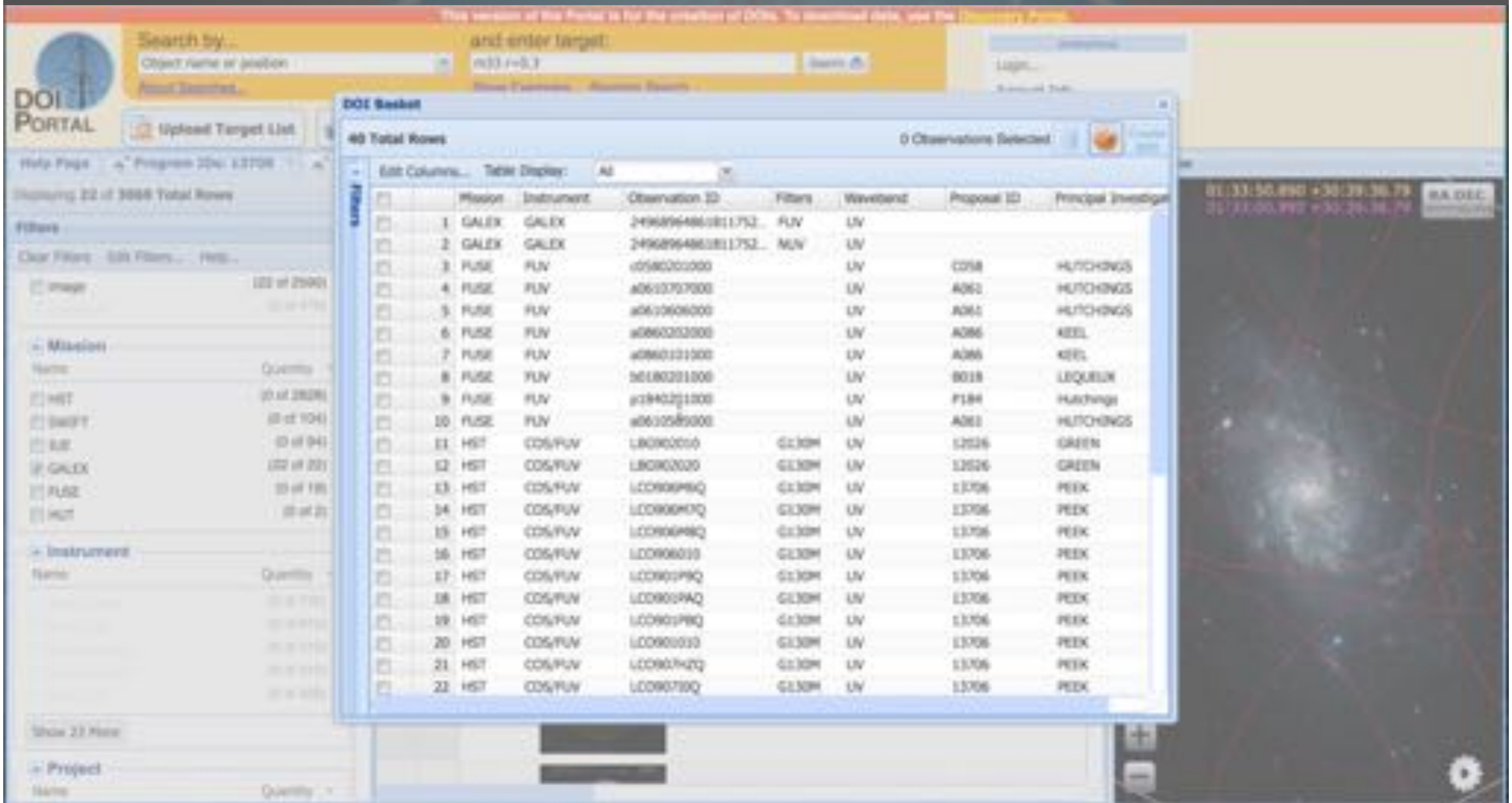

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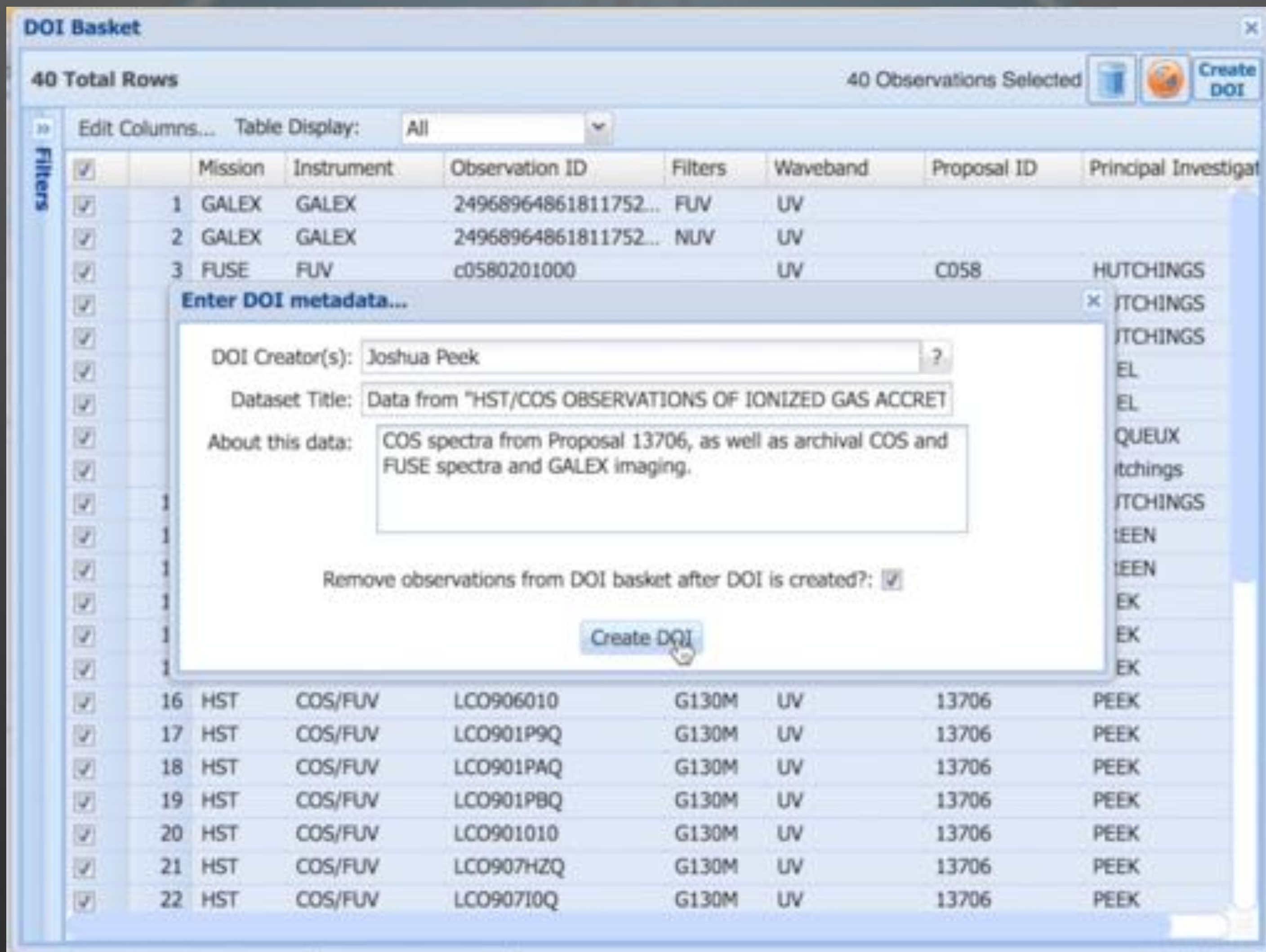

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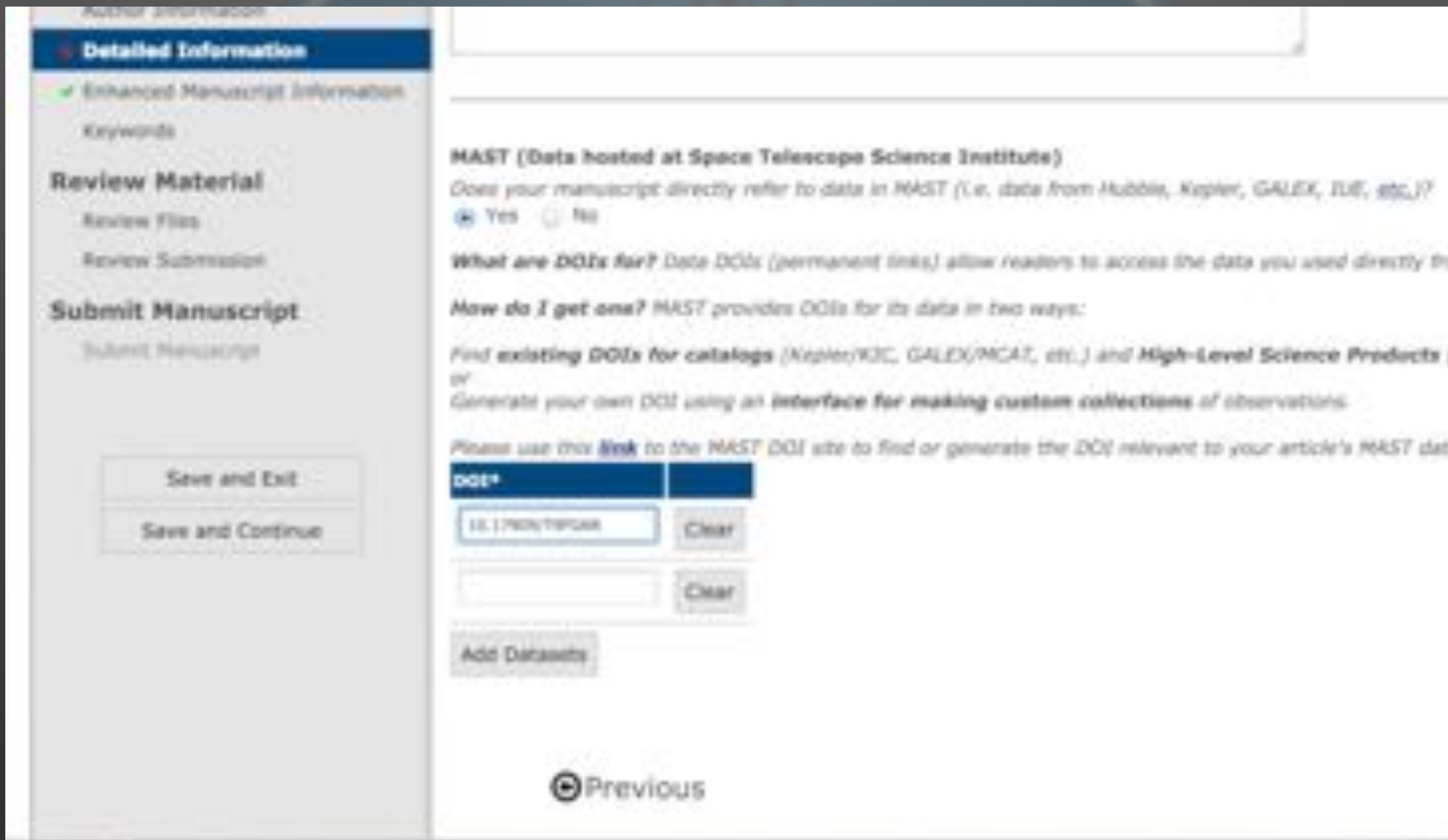

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scopic Explorer (FUSE) spectra for our sightlines from the MAST archive. The spectra span a wavelength range of 905-1187Å, within which an important absorption line O VI λ1032Å lies. Since O VI is not the main focus of this work and we only use it as a comparison with Si IV, a quick spectral reduction is performed. The stellar continuum within ±1000 km s^{-1} of 1032Å is normalized by fitting first and second order polynomial functions. We do not attempt to run Voigt-profile decomposition for this line given the low S/N of the data. The normalized O VI absorption lines and relevant discussion is present in Section 6.1.

All these can be found in the MAST Archive here: [10.17909/T9FG6R].

2.2. Wavelength Calibration and Spectral Co-addition

As mentioned in Section 2.1, each sightline is observed with four exposures which result in four spectra that need to co-add. The standard CalCOS pipeline provides data reduction for spectral co-adding and wavelength calibration, which has an accuracy of $\sim$ 0.06Å ($\sim$ 15 km s^{-1} at 1260Å). However, several authors have

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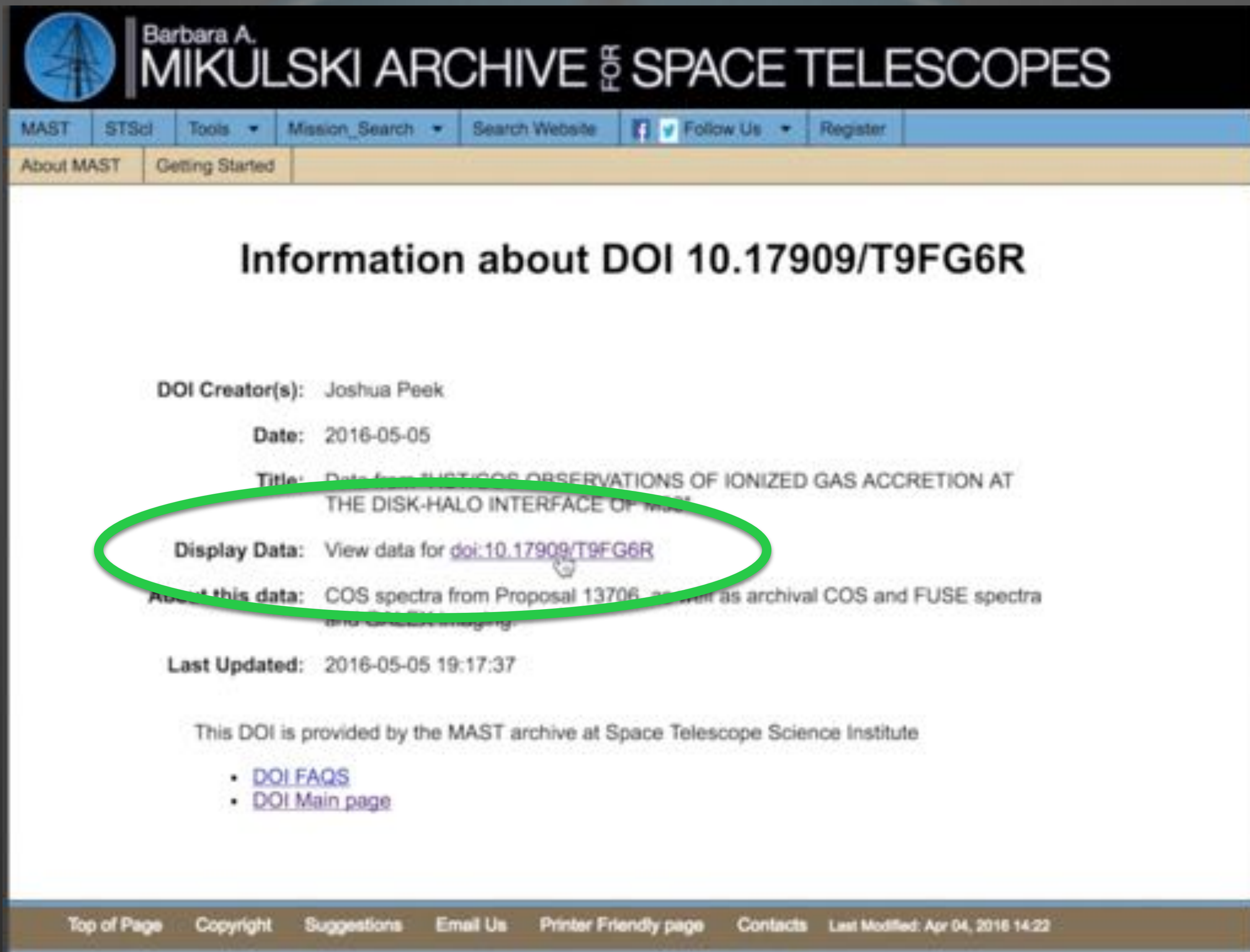

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