

# Studying low-mass stars in the VO... and other things! (yet another change of title)

Amelia Bayo (transiting from MPIA to Valparaiso)

Disclaimer: Blames are more likely on me... good ideas, if any, most likely coming from Enrique Solano or Carlos Rodrigo



# Outline

From a hybrid-user perspective (very biased, and not pretending AT ALL to be complete):

- What kind of science cases have benefited HUGELY from the VO initiative?
- Where is ample room for improvement?
- Could I make some suggestions?

- Data related:
  - CDS wonders vs pain of getting, for example, ApJS IOP tables

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(	10	10	10 VVV CL015	14	300.967
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	12	12	12 [DB52003] 78	0.5	301.118
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	24	24	24 Teutsch 109	02	303.652
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	37	37	37 VVV CL022	14	305.362
	38	38	38 [MCM2005b] 36	09	305.383
	39	39	39 Danks 2	01.17	305.392
	40	40	40 VVV CL023	14	305.438

- Data related:
  - CDS wonders vs pain of getting, for example, ApJS IOP tables

				TOPCAT(33): Table	Browser	
04370+2559 (A, B) 04385+2550 (A, B)	2	4.3 \sim K3-M1 C A \gg B 20, 21 18.9 M0 C A \gg B 22, 2	0 ¥			
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FS Tau (Aa, Ab, B)	3	0.23 (Aa, Ab), 20 (A-B) M1+M4 (Aa, Ab) C Aa > Ab 1, 24		[DBS2003] 77	05	300.966
FV Tau (A, B) 2	0.72	$K5+K6$ C A \sim B, $FV > FV/c$ 1	10	VVV (1015	14	300.984
FX Tau (A, B) 2	0.89	M1+M4 C+W A > B 1, 6	12	(DES2003) 78	14	301.118
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GG Tau (A, B) 4	10.3	\ldots C A \gg B 1	14	[FSR2007] 1622	11	301.416
GG Tau (Aa, Ab) 2	0.25	K7+M0.5 C Aa \gtrsim Ab 1	15	G3CC 6	17	301.643
GG Tau (Ba, Bb) 2	1.48	M5.5+M7.5 C Ba > Bb 1	16	NGC 4609	01	301.895
GH Tau (A, B) 2	0.31	M1.5+M2 C A \sim B 1	17	G3CC 7	17	301.947
GI Tau 2 12.9	K6	C GI \sim GK 5, 6	18	Hogg 15	01	302.047
GK Tau (A, B) 2	2.5		19	VVV CL018	14	302.158
Haro 6-37 (As Ab B)	0.33		20	[MCM2005b] 34	09	302.433
HE Tau (A B) 2	2.34	2.52(R, D), 0.53(Ra, RD) KITAL C $Ra > RD, R > D 1, 11$	21	[FSR2007] 1630	11	302.612
HN Tau (A, B) 2	3.11	K54M4 C A \g B 1	22	[DB52003] 79	05	302.64
HP Tau (A, B) 2	0.017	K2 C A > B 7, 8, 15	23	[DB52003] 80	05	302.806
IS Tau (A, B) 2	0.22	K7+M4.5 C+W A > B 1	24	Teutsch 109	02	303.652
IT Tau (A, B) 2	2.39	K3+M4 C A \gtrsim B 1, 6	25	G3CC 8	17	303.927
RW Aur (A, B, C)	3	1.42 (A-BC), 0.12 (B-C) K1+K5 (A, B) C A > B \gg C 1, 10	26	G3CC 9	17	304.002
T Tau (N, Sa, Sb)	3	0.70 (N-S), 0.1 (Sa-Sb) K0 C N \sim Sa \sim Sb 1, 3	27	VVV CL019	14	304.805
UX Tau (A, B, C)	4	5.86 (A-B), 2.63 (A-C) K5+M2+M5 C+W+W A > B, A \gg C 1	20	[MCM20050] 35	0.9	304.87
UX Tau (Ba, Bb) 2	0.138	M2 W Ba > Bb 11	30	C3CC 10	17	304.887
UY Aur (A, B) 2	0.88	NO+M2.5 C A \gtrsim B 1, 17	31	[DB\$2003] 82	05	304,928
UZ TAU (A, BA, BD)	4	SB (A), 3.54(A-Ba), 0.37 (Ba-BD) M1+M2+M2 C A > B, Ba \sim Bb 1, 13	32	[DRS2003] 131	05.17	305,259
V/10 Tau (A, B) Z	3.17	NU.5TM2 (TW A 1518 B 1 CR (30) 0.12 (30:0) 0.24 (30:0) 22400 (30:0) 840 0 5.0 5.0 5.1	33	[DB52003] 130	05	305.269
V807 Tau (AD, C, D)	3	SO $(AD)$ , V.12 $(AD-C)$ , V.24 $(AD-D)$ A2TAU $(AD, C)$ WTC $D > C > AB 1, 4$ 0.20 $(AD, B)$ , 0.4 $(Ba-B)$ V74M3 $(AD, C)$ A $B$ , $Ba$ leim $Bb$ 25 1	34	VVV CL021	14	305.277
V892 Tau (As. Ab. B)	3	0.06 (4.10 B9402 W As by b A or B 16.10	35	[DB52003] 132	0.5	305.321
V955 Tau (A. B) 2	0.33	KSewi C A > B 1	36	Danks 1	01,17	305.338
VY Tau (A, B) 2	0.66	NO W A > B 1	37	VVV CL022	14	305.362
XZ Tau (A, B) 2	0.30	N3+M1.5 C B > A 1	38	[MCM2005b] 36	09	305.383
ZZ Tau IRS 2	35	N4.5 C 22 IRS > 22 2	39	Danks 2	01,17	305.392
22 Tau (A, B) 2	0.04	N3 C A \gtrsim B 9	40	VVV CL023	14	305.438
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- Data related:
  - CDS wonders vs pain of getting, for example, IOP tables
  - The "sasmirala" atlas



Asmus et al 2014



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able I	Browser for 31: TAP 1 sasmin	rala.objects		Description
	name	raj2000	dej2000	
1	3C 390.3	280.5375	79.77139	PKS 2158-380/MCG-6-48-13 is a radio-loud lanticular calaxy at a radiohit of $z = 0.0334$ ( $D = 140$ Mpc) with a Sy 2 nucleus fueron-
2	NGC 1275	49.95083	41.51167	cetty catalogue 2010 and was first studied in deal by flosbury very 1982. HST observations revealed three compact but resolved
3	NGC 6251	248.13333	82.53778	sources in the nucleur instead of one central sources (total extend 1 arcsec 0.7 km; PA 90'; house faint 1998
4	3C 305	222.33989	63.27055	zirbel utraviolet 1998), in addition, water maser emission was detected in this object Kondratko discovery 2006L No. Solizer data
5	NGC 5866	226.62292	55.76333	are available for PKS 2158-380, which was imaged with VISIR in the SIC filter in 2006 (van der wolk dust 2010). A compact MIR
6	Mrk 266NF	204.57414	48.27806	nucleus is weakly detected in the image. The low S/N prevents a quantitative analyses of the source morphology but the latter seems
7	Mrk 266SW	204.57213	48.27556	different than that seen in HST, as only one source was detected. Our nuclear photometry is consistent with the value in
8	M51a	202,46958	47.19528	(van der wolk dust 2010).
9	NGC 4258	184.73958	47.30389	[boyce_faint_1996] P. J. Boyce, M. J. Disney, F. Macchetto, A. Boksenberg, J. C. Blades, and C. D. Mackay. Faint object camera
10	Mrk 3	93,90167	71.0375	observations of complex nuclear structure in PKS 2158-380, A&A, 305 pp. 715, January 1996.
11	NGC 3147	154.22375	73.40083	Tosbury_very_1982] K. A. E. Fosbury, A. Boksenberg, M. A. J. Snigers, I. J. Danziger, M. J. Disney, W. M. Goss, M. V. Penston, W. Wanstaker, K. J. Wolfachard, and A. S. Wisson, Very extended invized as in radio galaxies i, a radio galaxie
12	4C +73.08	147.44108	73.23976	PKS 2158-380, MVRAS, 201 pp. 991–1008, December 1982.
13	M81	148.88833	69.06528	[kondratko_discovery_2006] P. T. Kondratko, L. J. Greenhill, J. M. Moran, J. E. J. Lovell, T. B. H. Kuiper, D. L. Jauncey, L. B.
14	UGC 5101	143.965	61.35306	Cameron, J. F. Gómez, C. Garcia-Miró, E. Moll, I. de Gregorio-Monsalvo, and E. Jiménez-Bailón. Discovery of water maser emission
15	NGC 3690E	172.14012	58.56294	in explicit ALSNS with r/u m antennas of NASA's deep space network. ApJ, 638 pp. 100–105, February 2006.
16	NGC 3690W	172.12925	58.56131	February 2010.
17	NGC 3998	179.48375	55.45361	[veron-cetty_catalogue_2010] MP. Véron-Cetty and P. Véron. A catalogue of guasars and active nuclei: 13th edition. A&A , 518 pp.
18	NGC 3982	179.11708	55.12528	10, July 2010.
19	NGC 3718	173.14542	53.06806	[Zrbe] utraviolet 1998] Esther L. Zirbel and Stell A. Baum. Ine utraviolet continuum emission of radio galaxies. L description of sources from the hubble source tolescore architecture. 4 or 5, 114 pp. 127. Exercutery 1998.
20	IRAS 08572+3915	135.10583	39.065	startes non on house space enstance archives, Apro, 114 pp. 177, Pedidary 1860.
21	PKS 2158-380	330.32125	-37.77333	1
22	NGC 7130	327.08125	-34.95111	
23	NGC 7172	330.50792	-31.86972	Images
24	IC 1459	344.29417	-36.46222	
25	NGC 7496	347.44708	-43.42806	
26	NGC 7552	349.045	-42.58472	
27	NGC 7582	349.59792	-42.37056	
28	NGC 7590	349.72833	-42.23917	
29	NGC 7314	338.9425	-26.05056	
30	PKS 2354-35	359.25292	-34.75917	50
31	ESO 602-25	337.85625	-19.03444	
32	MR 2251-178	343.52417	-17.58194	
33	Mrk 915	339.19375	-12.54528	
34	3C 445	335.95625	-2.10361	e
35	Mrk 926	346.18125	-8.68583	
36	NGC 7592W	349.59084	-4.41574	
37	ESO 297-18	24.655	-40.01139	
				URL: http://dc.zah.uni-heidelberg.de/sasmirala/q/prod/qp/PKS%202158-380

- Data related:
  - CDS wonders vs pain of getting, for example, IOP tables
  - The "sasmirala" atlas
- Tool related (development)
  - The final AVO science demo







√ 17657 sources
 with good quality
 MSX photometry
 8-14 micron

✓ 3278 with
 SIMBAD class.
 ✓ 155 known PNe
 or Post-AGB stars
 ✓ Confusion with
 other type of
 sources

#### 2005 AVO demo: the PPNe case



If we harden our selection criteria:  $\sqrt{|b|} \ge 2$  degrees  $\sqrt{[A]-[C]} \ge 0.7$  $\sqrt{[C]-[D]} \ge 0.7$ 

Large majority of PNe and Post-AGB stars... and many new candidates!

#### 2005 AVO demo: the PPNe case



#### 2005 AVO demo: the PPNe case





- Data related:
  - CDS wonders vs pain of getting, for example, IOP tables
  - The "sasmirala" atlas
- Tool related (development)
  - The final AVO science demo
  - The birth of VOSA (and its continuous development)

#### Warning! self-promotion

#### Cool objects: From SED fitting to age estimation.

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<sup>1</sup>Laboratorio de Astrofísica Espacial y Física Fundamental (LAEFF-INTA), P.O. 50727, E-28080 Madrid, Spain <sup>2</sup>Spanish Virtual Observatory, Spain <sup>3</sup>Centre de Recherche Astronomique de Lyon (CRAL), Ecole Normale Supérieure de Lyon, 69364, Lyon, France

#### Abstract

One of the typical tools to estimate physical parameters of almost any kind of astronomical object is to perform a fitting of synthetic spectra or photometry extracted from theoretical models to observational data. This process usually involves working with multiwavelength data, which is one of the cornerstones of the VO philosophy. From this kind of studies, when combining with theoretical isochrones one can even estimate ranges of ages. We present the results from a code designed to perform  $\chi^2$  tests following two different methodologies to fit observational data: using grids of models (on their synthetic photometry), and combinations of blackbodies (including modified blackbodies). In particular, we use the models by the Lyon group. Some steps in this process can already be done in a VO environment, and the rest are in the process of development. We must note that this kind of surveys in star forming regions, clusters, etc. produce a huge amount of data, very tedious to analyse using the traditional methodology. Therefore this is an ideal example of the VO capabilities.

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Building the multiwavelength SEDs (B) SPECVIEW

V March 2007

LACFF

#### Warning! self-promotion



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1	LOri001	10,228	0,003	10,255	0,004	10,214	0,009	10,206	0,01	-
2	LOri002	9,935	0,003	10,042	0,003	9,93	0,009	9,88	0,008	
3	LOri003	10,262	0,003	10,318	0,004	10,239	0,01	10,171	0,01	-
4	LOri004	10,287	0,003	10,249	0,004	10,185	0,009	10,127	0,009	
5	LOri005	10,204	0,003	10,321	0,004	10,218	0,009	10,158	0,009	
6	LOri006	10,454	0,003	10,454	0,004	10,399	0,011	10,319	0,01	
7	LOri007	10,668	0,004	10,636	0,004	10,615	0,012	10,482	0,013	
8	LOri008	10,498	0,003	10,495	0,004	10,44	0,011	10,256	0,012	
9	LOri009	10,834	0,004	10,873	0,005	10,788	0,012	10,743	0,014	
10	LOri010	10,916	0,004	10,953	0,005	10,733	0,012	10,839	0,016	
11	LOri011	10,378	0,003	10,521	0,004	10,444	0,011	10,326	0,011	
12	LOri012	10,619	0,003	10,758	0,005	10,627	0,012	10,543	0,012	
13	LOri013	10,511	0,003	10,48	0,004	10,467	0,011	10,344	0,012	
14	LOri014	10,902	0,004	10,904	0,005	10,839	0,014	10,797	0,014	
15	LOri015	10,808	0,004	10,886	0,005	10,824	0,013	10,882	0,015	
16	LOri016	10,833	0,004	10,817	0,006	10,378	0,011	10,7	0,014	
17	LOri017	11,165	0,005	11,206	0,006	11,173	0,017	11,072	0,019	
18	LOri018	10,804	0,004	10,798	0,005	10,722	0,012	10,636	0,014	
19	LOri019	10,88	0,004	10,866	0,005	10,767	0,013	10,788	0,018	
20	LOri020	10,676	0,003	10,609	0,004	10,573	0,012	10,485	0,012	
21	LOri021	11,129	0,004	11,107	0,005	11,081	0,016	11,065	0,019	
22	LOri022	11,01	0,004	10,985	0,005	10,895	0,014	10,683	0,014	
23	LOri023	11,09	0,004	11,114	0,005	11,071	0,015	10,928	0,018	
4	1.000.004	44.040	L 0.004	44.040	0.00F	40.070	0.04F	40.077	A 640	

Photometric data in four bands.



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#### Warning! self-promotion



### And VOSA came to life!



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		Files TSAP	LOHOO1	83:26:47	9:55:38	N	0.31	83:26:48	9:55:38	10.595±0.022	11.297±0.022	10.426±0.021							
		Photometry fit Isochrones	LO4003	83:58:51	9:58:31	N N	0.35	83:58:51	9:56:30	10.32940.023	11.416±0.023	10.524±0.023							
Bavo et al. (	2008)		LOn004	83:56:53	9:45:50	1	0.49	83:56:53	9:45:49	10.780±0.023	11.359±0.022	10.548±0.021							
			LO4006	83:28:24	9:43:08	V	0.14	83:28:24	9:43:08	10.549±0.022	11.378±0.022	10.354±0.023							







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		TSAP P	L0x001	83:26:47	9:55:38	¥	0.31	83:26:48	9:55:38	10.595±0.022	11.297±0.022	10.426±0.021						
		Photometry fit I Isochrones	LOH003	83:58:51	9:58:31	V.	0.35	83:58:51	9:58:30	10.725±0.022	11.418±0.023	10.524±0.023						
Bavo et al	(2008)		LOH004	83:56:53	9:45:50	<b>v</b>	0.49	83:56:53	9:45:49	10.780±0.023	11.359±0.022	10.548±0.021						
Dayo or al.	(2000)		LOx005	83:28:24	9:43:08	<b>v</b>	0.14	83:28:24	9:43:08	10.549±0.022	11.378±0.022	10.354±0.023						

## And VOSA came to life!



IT'S ALIVE



## And there was room for improvement

- "Limited to" / "conceived for" stars and brown dwarfs, what about older sources? and more massive? and science-fiction uhmm extragalactic studies?
- Reflected in the available collections of models: Kurucz, NextGen, COND, DUSTY and not many more
- Brute force fitting but no study of the relevance of the individual parameters to the fit
- No A<sub>V</sub> estimation
- Not design to work with a single object (input format)
- Variety of catalogs offered but you can always do better and also look for more than photometry
- No Isochrone interpolation, make it even more VO!
- Anything else in the wish-list?

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~200 regular users, cited in ~ 50 papers





2PMASS has uniformity scanned the entire sky in three near-initiared bands to detect and characterize point sources brighter than about 1 It consists of a set mJy in each band, with signal-to-noise ratio (SNR) greater than 1. DENIS survey and 12 arcmin More Info. Filters: 2MASS/2MASS.J 2MASS/2MASS.H ZMASS/2MASS.Ks Search radius: 5 arcsec Show magnitude limits **IRAS Catalog of Point Sources, Version 2.0** This is a catalog of some 250,000 well-confirmed infrared point sources observed by the Infrared Astronomical Satellite, i.e., sources with angular extents less than approximately 0.5, 0.5, 1.0, and 2.0 arcmin in the in-scan direction at 12, 25, 60, and 1. More Info. Filters: VIRAS/IRAS.12mu VIRAS/IRAS.25mu Search radius: 5 arcsec Show magnitude limits AKARI/IRC mid-IR all-sky Survey (ISAS/JAXA, 2010) The AKARI/IRC Point Source Catalogue Version 1.0 provides positions and fluxes for 870,973 sources observed with the InfraRed Camera (IRC). More Info. Filters: AKARI/IRC.S9W AKARI/IRC.L18W Info. Search radius: 5 arcsec Show magnitude limits

#### **C2D Spitzer and Ancillary Data**

C2D Fall '07 Full CLOUDS Catalog (CHA\_II, LUP, OPH, PER, SER). Spitzer/IRAC.12 Filters: Spitzer/IRAC.I1 Spitzer/IRAC.I3 Spitzer/IRAC.14 Spitzer/MIPS.24mu Spitzer/MIPS.70mu

Search radius: 5 arcsec Show magnitude limits

This catalogue is the latest incremental release of the DENIS project. svo2.cab.inta-csic.es/theorv/vosa Filters: DENIS/DENIS.I Search radius: 5 arcsec Show magnitude limits

#### MSX6C Infrared Point Source Catalog

Version 2.3 of the Midcourse Space Experiment (MSX) Point Source Catalog (PSC), which supersedes the version (1.2) that was released in 1999 (Cat. V/107), contains over 100,000 more sources than the previous version.. More Info.

Filters: MSX/MSX.A MSX/MSX.C MSX/MSX.D MSX/MSX.E

Search radius: 5 arcsec Show magnitude limits

#### **AKARI/FIS All-Sky Survey Point Source Catalogues** (ISAS/JAXA, 2010)

The AKARI/FIS All-Sky Survey Bright Source Catalog Version 1.0 provides positions and fluxes for 427071 point sources in the 4 far-infrared wavelengths centered at 65, 90, 140 and 160µm. More

Filters: AKARI/FIS.N60 AKARI/FIS.WIDE-S AKARI/FIS.WIDE-L AKARI/FIS.N160

Search radius: 5 arcsec Show magnitude limits

#### GLIMPSE Source Catalog (I + II + 3D)

The Galactic Legacy Infrared Midplane Survey Extraordinaire (GLIMPSE), is a survey of Galactic Plane central parts made with the Infrared Array Camera (IRAC) aboard the Spitzer Space Telescope (SST)., More Info.

Filters: Spitzer/IRAC.I1 Spitzer/IRAC.I2 Spitzer/IRAC.I3 Spitzer/IRAC.I4

Search radius: 5 arcsec Show magnitude limits



~200 regular users, cited in ~ 50 papers







This publication makes use of VOSA, developed under the Spanish Virtual Observatory project supported from the Spanish MICINN through grant AyA2008-02156.

#### Bayo et al. (2008, 2014a subm.)

~200 regular users, cited in ~ 50 papers

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